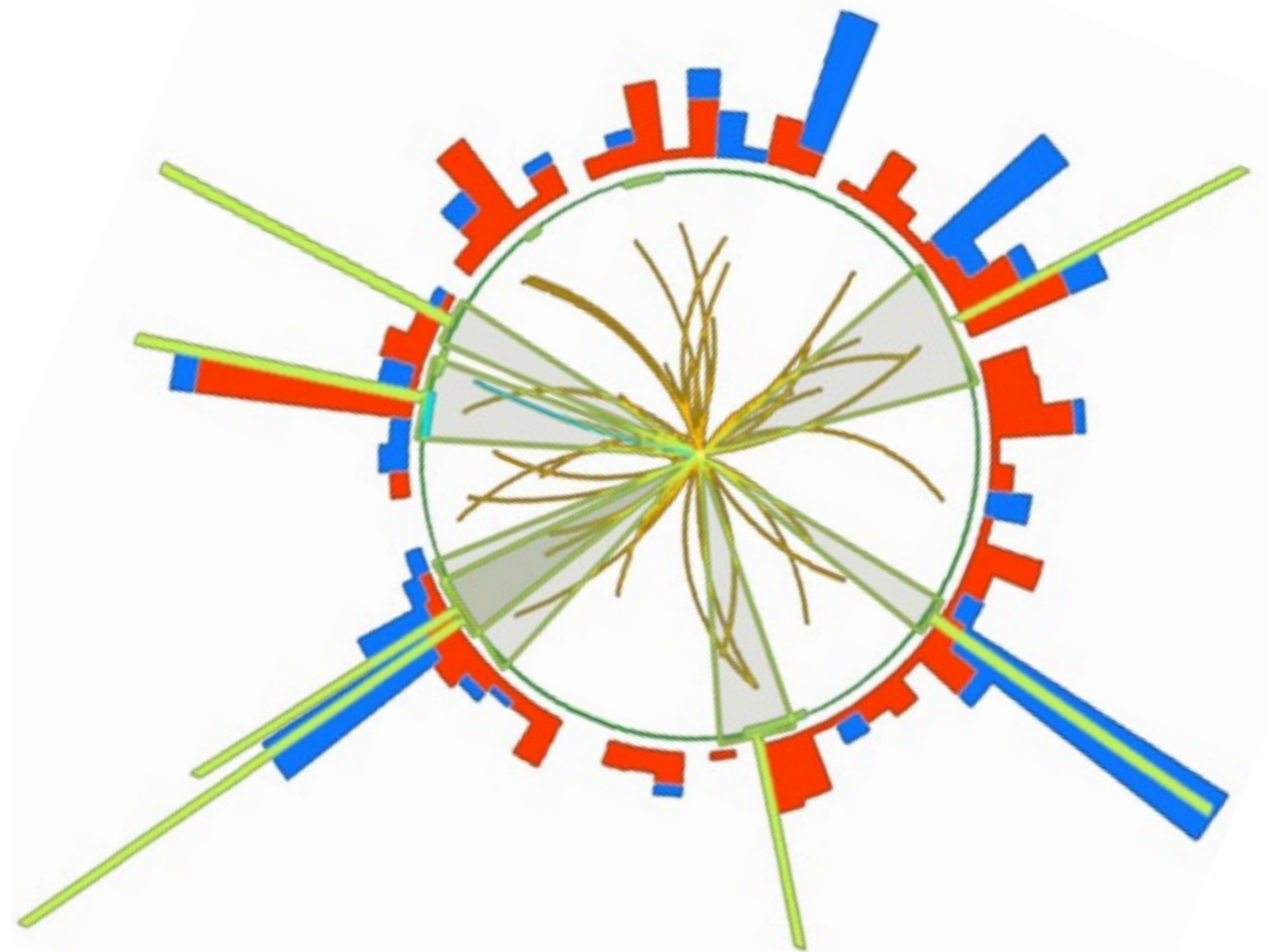
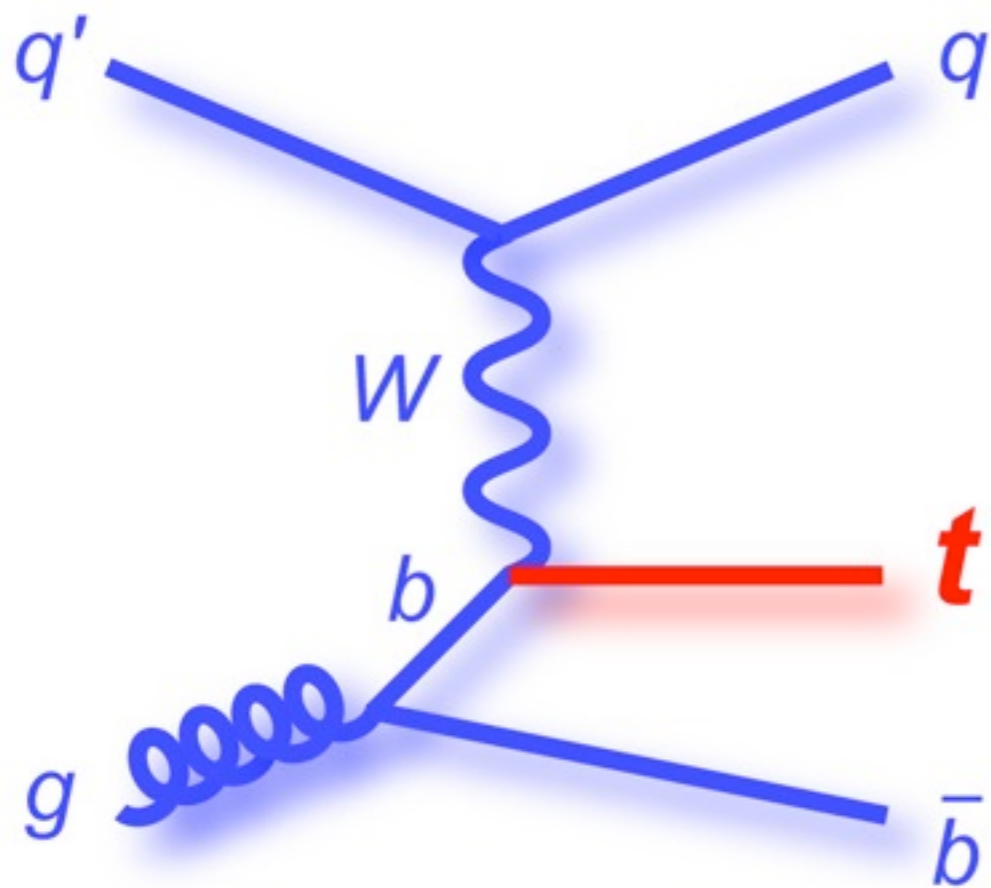


Introduction to Experimental HEP

Dave Newbold – University of Bristol



Overview

- ▶ Experimental HEP – **what we do, and why**
 - ▶ Nuts, bolts, and big toys
 - ▶ Where we stand today
 - ▶ LHC programme – first major results
 - ▶ The far future
-
- ▶ You will not hear the complete story in two lectures!
 - ▶ We will focus on **energy frontier physics at colliders**
 - ▶ Absolutely not the only game in town
 - ▶ Perhaps not even the most interesting one
 - ▶ **Please ask questions!**

Theory and Experiment

▶ HEP theorists:

- ▶ Think about Lagrangians
- ▶ Particles = “field quanta”
- ▶ (Can) work in small teams
- ▶ Can rapidly play with new ideas
- ▶ Admire elegance, simplicity (~ Dirac)

▶ But also...

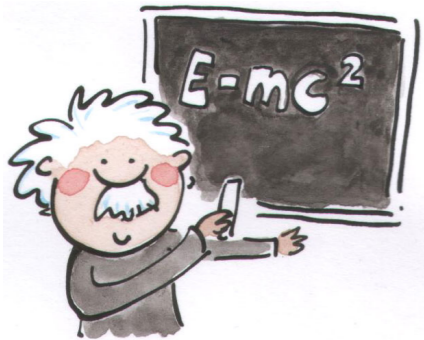
- ▶ Must invent new techniques
- ▶ Interact with other fields
- ▶ Get excited about new results

▶ HEP experimentalists:

- ▶ Think about measurements
- ▶ Particles = “tiny charged blobs”
- ▶ (Must) work in huge teams
- ▶ New ideas take years to test
- ▶ Admire ingenuity, effectiveness (~ Rutherford)

▶ But also...

- ▶ Must invent new technologies
- ▶ Interact with other fields
- ▶ Get excited about new results



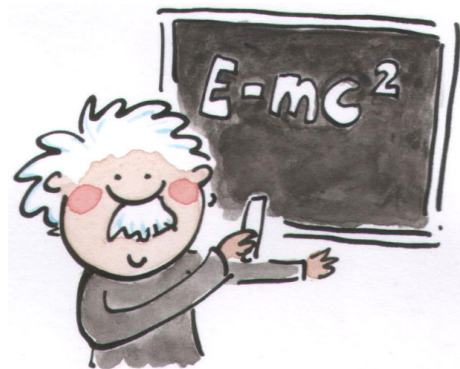
Real progress is **only** made when we work effectively together



Progress in HEP - The Ideal



Develop theoretical model with testable consequences



Devise experiment or observation to challenge model

Build and perform experiment

Scrap or further develop model

Unexpected results?



Progress in HEP - Reality Today

Ongoing pursuit of interesting / fashionable / promising problems

Form consensus on phenomena that might be observable (today)

Form consensus on next required large facility, carry out R&D

Finance (G\$), plan facility

Collaborations propose and design experiments (cash limited)

Perform experiments

Discoveries / anti-discoveries (perhaps unexpected!)

Precision measurements of parameters

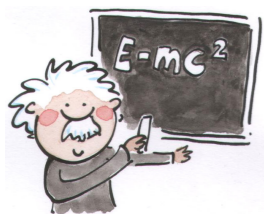
Scrap or further develop models

~ 5 - 10 years

~ 5 years

~ 5 - 10 years

~ 1 - 10 years



Motherhood

- ▶ Particle physics is a complicated business
 - ▶ In theory and experiment: “All the easy experiments have been done”
 - ▶ Nobody understands the whole field or all the techniques
 - ▶ You will meet ‘experimentalists’ who:
 - ▶ Know rather little about experiments
 - ▶ Know rather little about theory
 - ▶ All of the above
 - ▶ Advice to the budding professional physicist:
 - ▶ Understand what can be measured, what *can't* be measured - and why
 - ▶ Understand where the key sources of uncertainty are
 - ▶ Learn the language of the experimentalists and use it
 - ▶ Provide the tools to allow your models to be used
-
- ▶ Here endeth the lesson

What Can We Measure?

- ▶ Most often carry out statistical counting experiments
 - ▶ Where possible, prepare known initial state
 - ▶ Observe repeated collisions + try to measure the final state
 - ▶ Count rates of given final states \rightarrow cross sections
 - ▶ Examine distributions of parameters & compare with theory
- ▶ Static properties of bound states
 - ▶ Existence & mass
 - ▶ Quantum numbers (charges, J^{CP}) and couplings
 - ▶ Width / lifetime, branching ratios, mixing parameters
- ▶ Dynamic quantities
 - ▶ Cross sections as function of energy, momentum exchange, etc
- ▶ Often use ratios and derived quantities to cancel errors
 - ▶ Branching ratios, mixing angles, polarisations, decay parameters

Doing Experiments: Initial State

Decide on initial state



Build detector



Simulate performance



Record collisions



Reconstruct the events



Isolate a signal



Make measurements

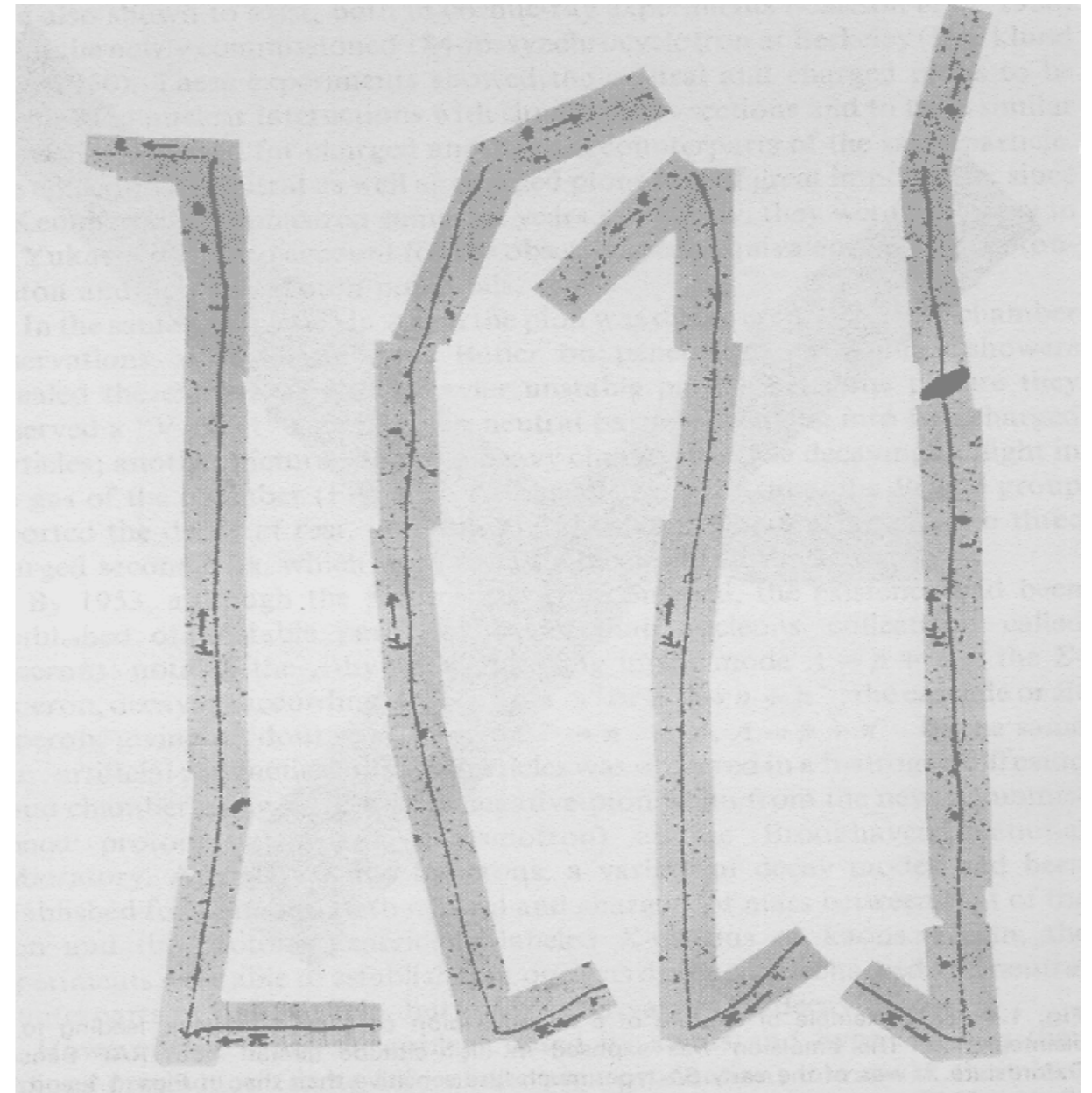
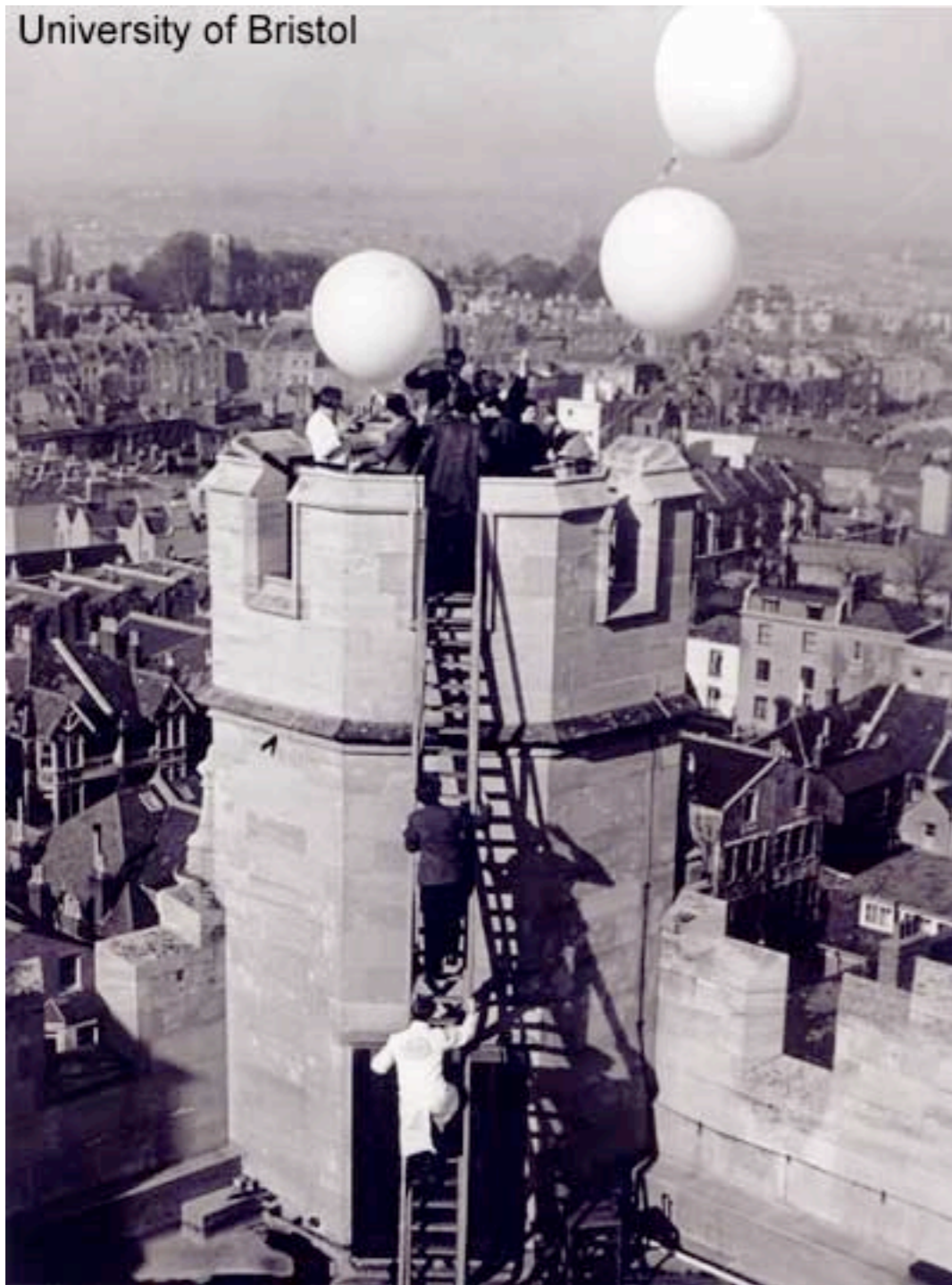


Compare with theory

- ▶ A source of high-energy particles is required
- ▶ Choices to make:
 - ▶ Identities of colliding particles
 - ▶ Energies: monochromatic / spread?
 - ▶ Colliding beams / fixed target?
 - ▶ Polarisation?
 - ▶ Particle flux?
- ▶ Realistically...
 - ▶ A free choice is not usually possible
 - ▶ Most experiments carried out at existing or new accelerator facility
 - ▶ Can also carry out 'observational experiments'

Natural Particle Sources

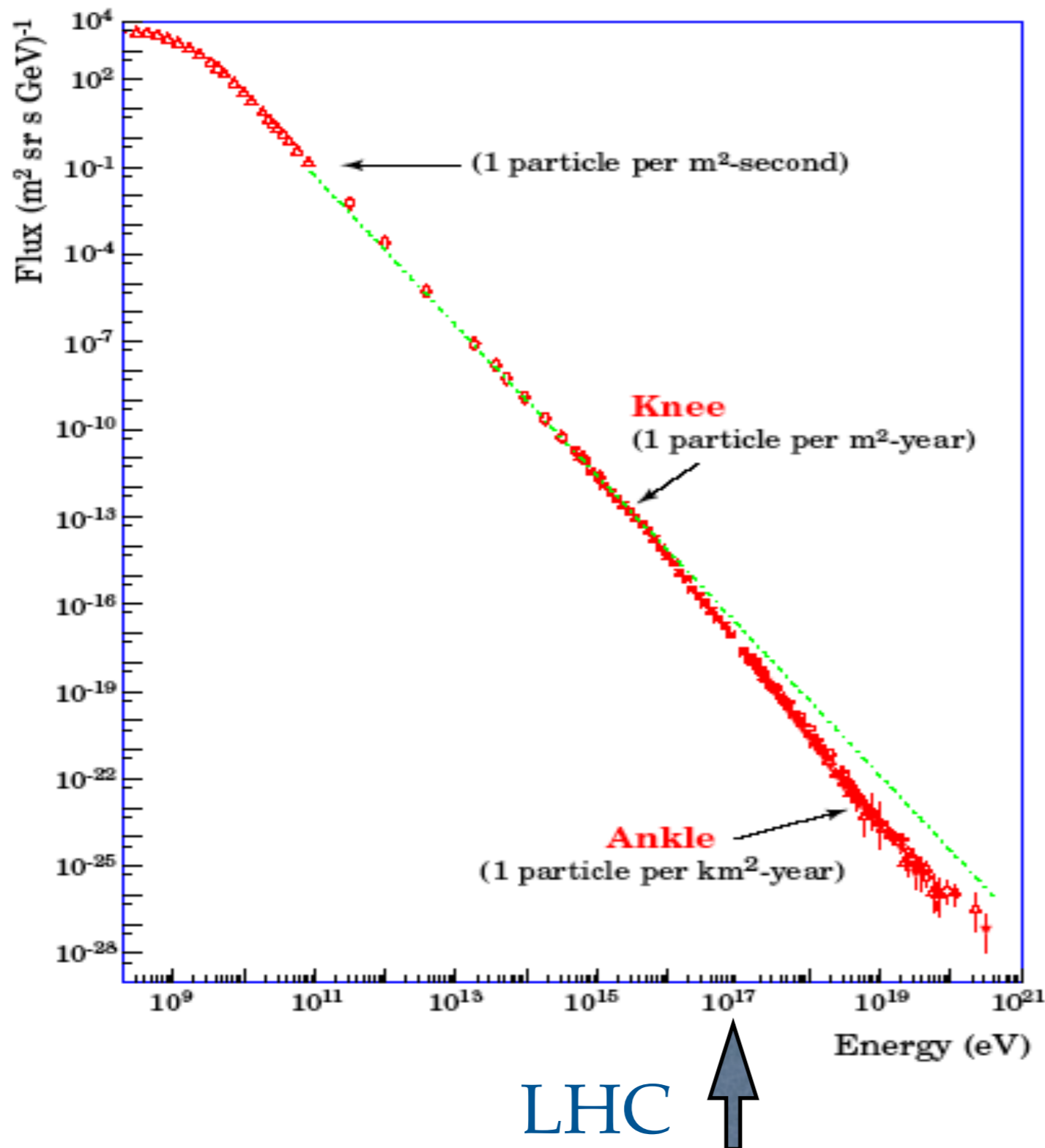
University of Bristol



Emulsions containing the first evidence of pion scattering - Powell, Nobel Prize 1950

- ▶ Cosmic rays main source of high-energy flux until 1950s

Cosmic Ray Spectrum



- ▶ Spectrum extends to (very) high energies
 - ▶ Subject of ongoing study
- ▶ Flux is low
 - ▶ Cannot be used to study rare processes
 - ▶ Is an important background for many experiments
- ▶ Natural sources also important for:
 - ▶ Solar neutrino studies
 - ▶ Reactor neutrino studies
 - ▶ Direct DM searches – ?

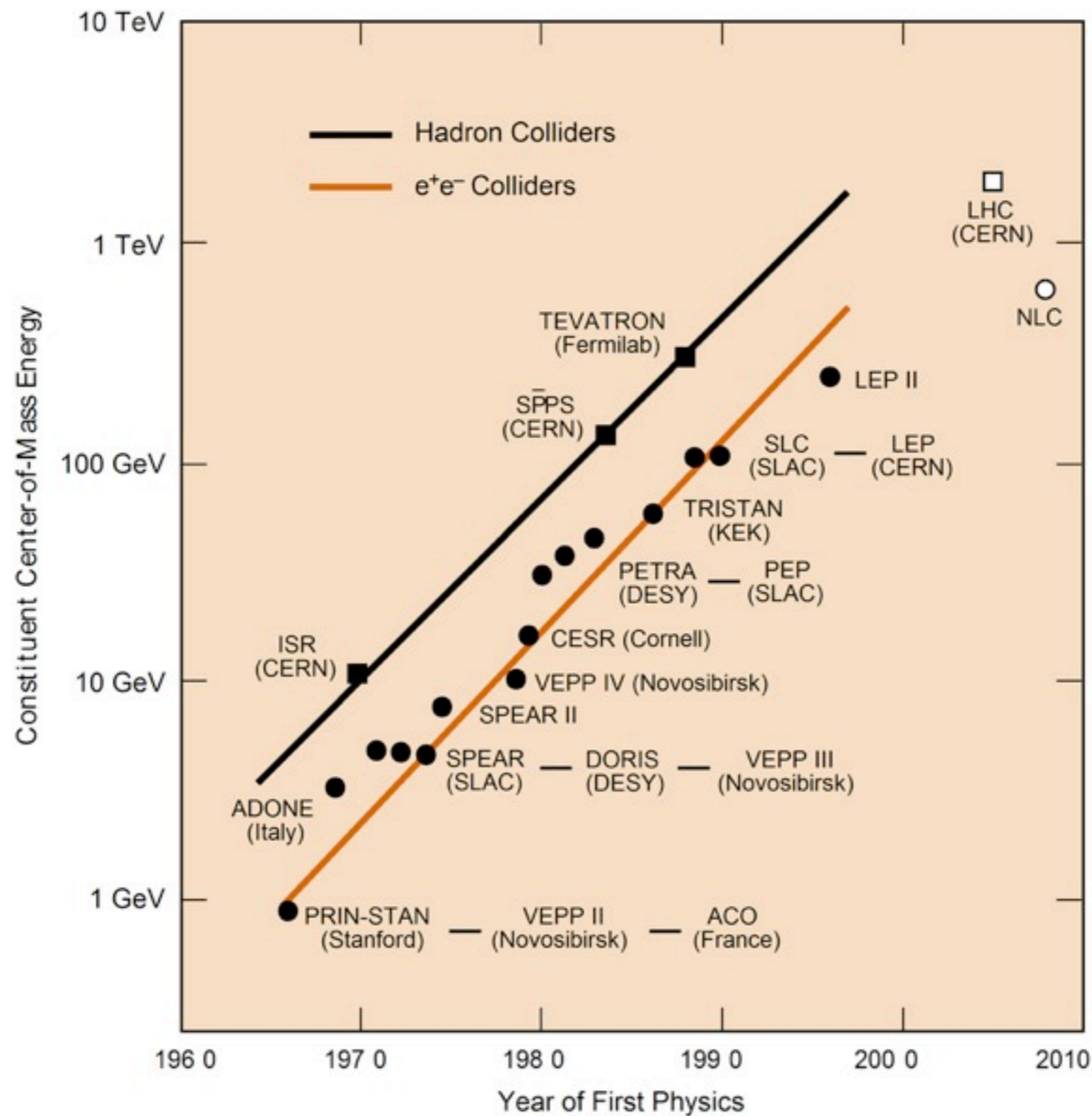
Particle Accelerators

- ▶ A 'reliable' and intense source of high-energy particles
 - ▶ Energies up to 3.5TeV per beam (record holder: LHC)
 - ▶ Colliding-beam luminosities up to $2 \cdot 10^{34} \text{cm}^{-2} \text{s}^{-1}$ (record holder: KEKB)
- ▶ Accelerator modes:
 - ▶ Fixed-target: an accelerated beam on a stationary target
 - ▶ Very high effective luminosities, usefully boosted collision frame
 - ▶ Secondary beams (e.g. neutral and / or unstable beams)
 - ▶ Colliding beams, equal energies
 - ▶ Maximum centre-of-mass energy obtained
 - ▶ Colliding beams, asymmetric energies
- ▶ What particles can be used?
 - ▶ Today: stable particles only
 - ▶ p, pbar, e-, e+, heavy ions (stripped nuclei)
- ▶ Accelerator physics is a sizeable discipline in its own right

Accelerators: Operating Principles

- ▶ Basic idea:
 - ▶ Accelerate charged particles through a potential gradient
 - ▶ Use resonant oscillating fields in (superconducting) **accelerating cavities**
- ▶ Circular accelerators
 - ▶ Life is easier if we reuse the same gradient repeatedly
 - ▶ Synchrotrons use **magnetic dipole field** to achieve closed orbit (~circle)
 - ▶ Synchrotron radiation losses $\sim m^{-4}$ – so no more circular e+e- machines
- ▶ Complexities
 - ▶ Beams will naturally diverge -> **cooling and focusing** necessary
 - ▶ Accelerators are the largest and most complex machines ever built
 - ▶ Practical + safety considerations -> often underground -> \$\$\$
- ▶ Basic figures of merit for accelerators
 - ▶ Collision energy, luminosity
 - ▶ *Integrated luminosities* often quoted (e.g. 1 / pb or 1 / fb)

Progress in Accelerators



See PDG review for parameters of current machines:
<http://pdg.lbl.gov/2010/reviews/rpp2010-rev-hep-collider-params.pdf>

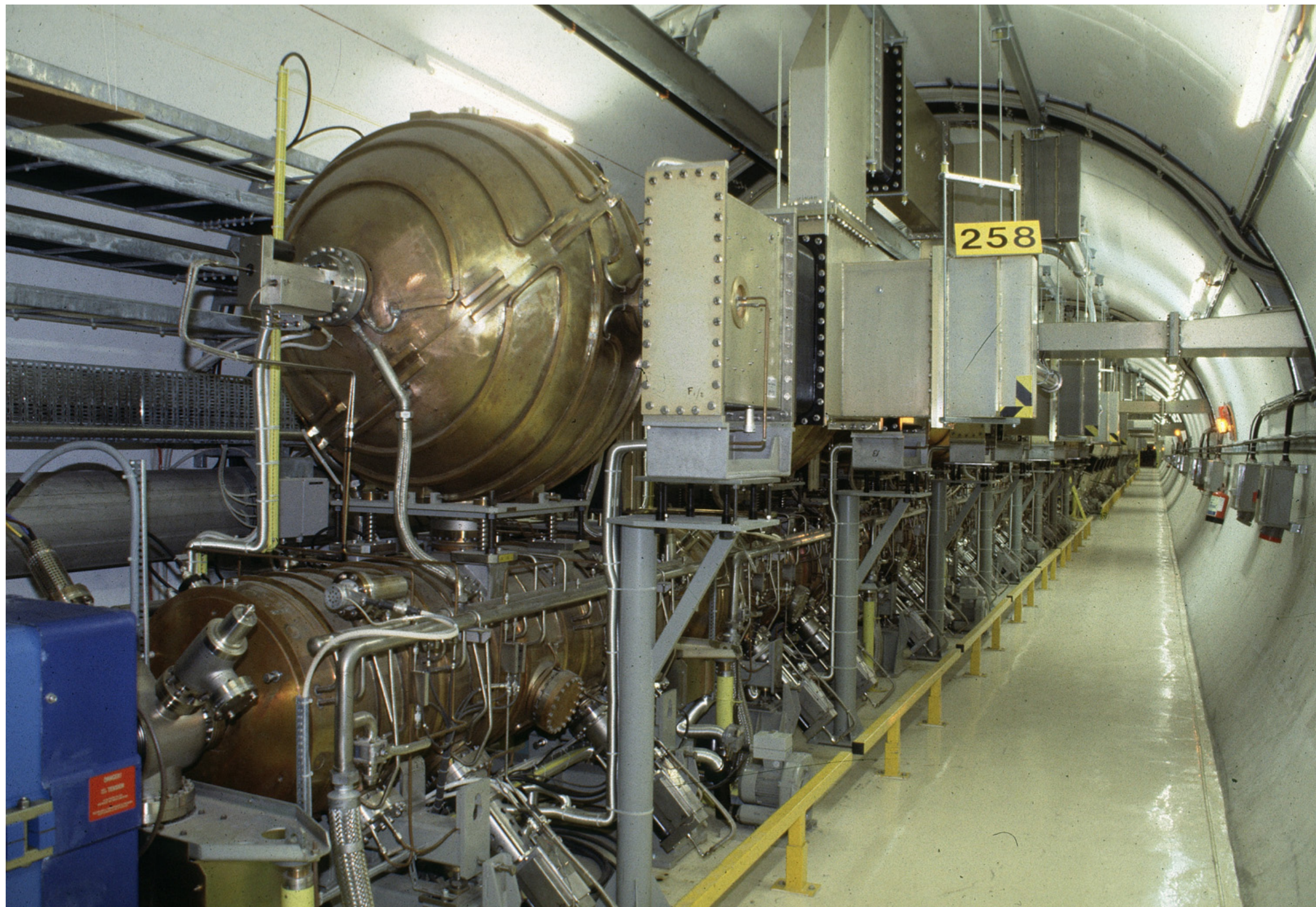
LEP / LHC @ CERN



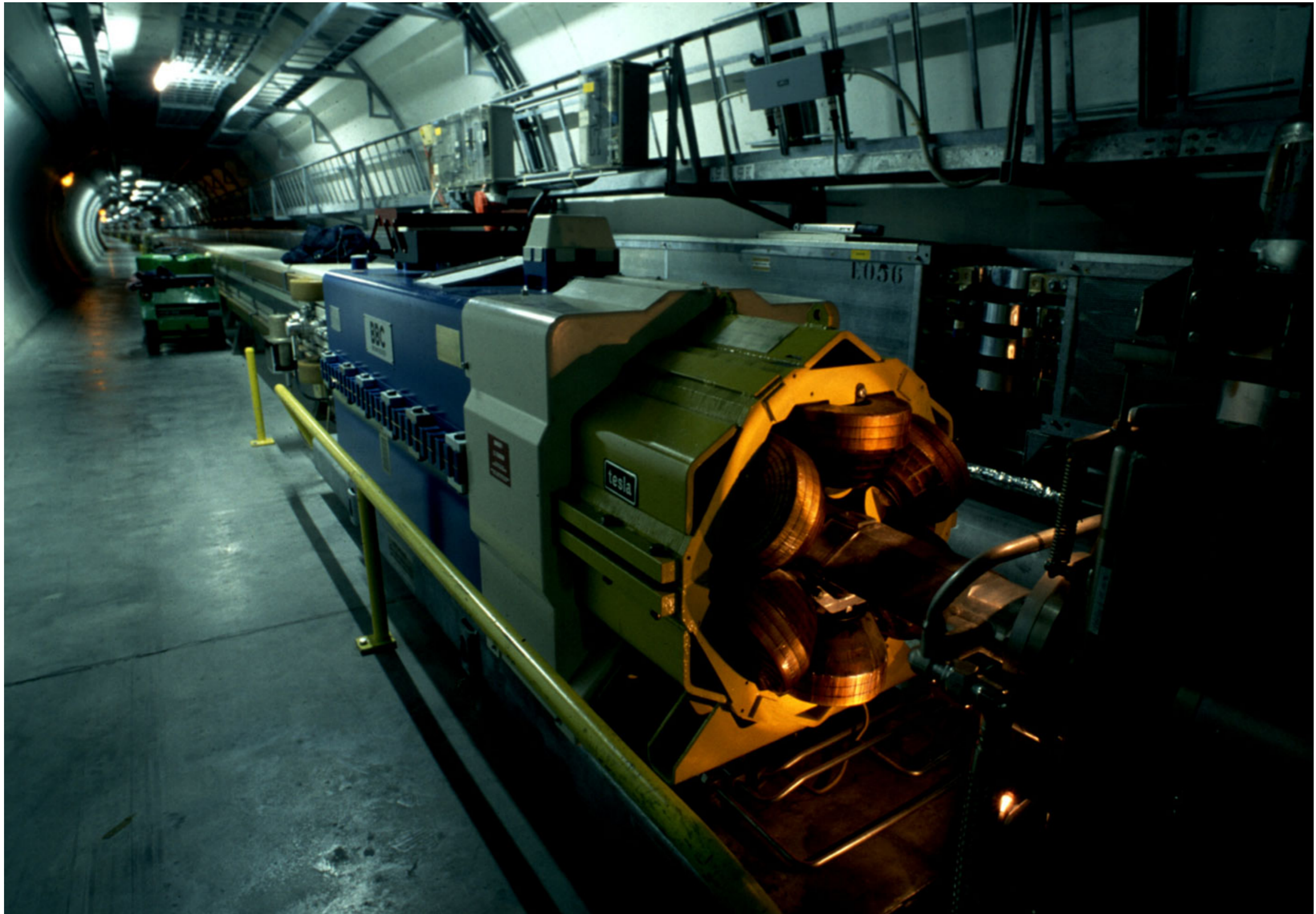
e+e- collider, 89-00; **LEP-1 'Z-factory'**: M_Z , 600 / pb , **LEP-2**: 200GeV, 2800 / pb

LHC: pp collider, 7TeV -> 14TeV (2013 -), 3000 / fb by 2030 [more later on LHC]

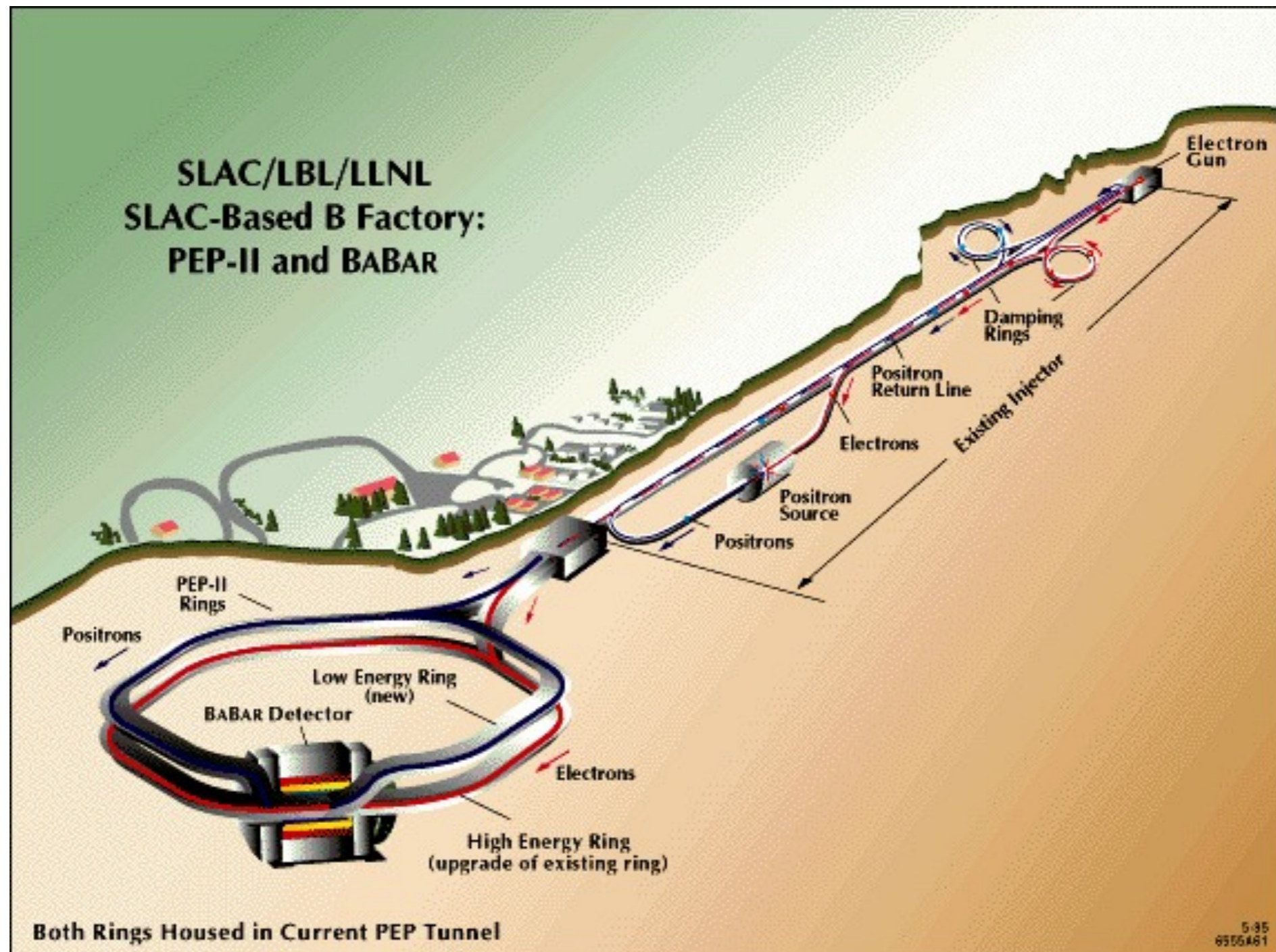
Inside LEP - Cavities



Inside LEP - Magnets



PEP-II @ SLAC



PEP-II: asymmetric e^+e^- b-factory (U_{4S} resonance), 240 / fb. Used SLC infrastructure.

HERA @ DESY



HERA: asymmetric ep collider (unique), 800GeV p on 30GeV e⁺ (or e⁻).

Tevatron @ FNAL



Tevatron: $p \bar{p}$ collider, 1TeV, 240 / pb (Run I), 2TeV, 7 / fb (Run II, ongoing)

Doing Experiments: Detector

Decide on initial state



Build detector



Simulate performance



Record collisions



Reconstruct the events



Isolate a signal



Make measurements



Compare with theory

- ▶ Experiment design / optimisation / construction
 - ▶ Needs predictions (MC) of physics signatures & detector performance
- ▶ Which experiment to build?
 - ▶ General purpose detectors are flexible, but expensive
 - ▶ Specialised detectors are optimised for one set of channels / studies
 - ▶ (almost) always have more than one detector studying the same physics
- ▶ Cost, complexity, timescales
 - ▶ All now very large. LHC GPDS:
 - ▶ ~0.5GCHF, 80M channels, 20 years
 - ▶ New technologies in continuous devt.

Basics of Particle Detection

▶ Charged particles

- ▶ Ionization – basis of most techniques (gaseous, solid state detectors)
 - ▶ Liberated charge is amplified in a potential gradient and detected
 - ▶ Photographic emulsions still used...
- ▶ Scintillation: excitation of a molecule or crystal lattice causes light emission
- ▶ EM effects: Cherenkov / transition radiation
 - ▶ Interaction of particle with dielectric medium causes light emission
 - ▶ Uniquely, can be directly sensitive to particle *velocity*

▶ Neutral particles

- ▶ Much more difficult - can detect only after interaction in material
- ▶ Weakly interacting neutrals can be inferred by their absence...

▶ Radiochemical effects

- ▶ Used (e.g.) for measurement of neutrino, WIMP fluxes

▶ Many other techniques exist, esp. in low background expts

- ▶ e.g. bolometric measurements, superconducting detectors

Detector Subsystems

▶ Tracking system:

- ▶ Measure trajectories of charged particles in a ~uniform magnetic field
- ▶ Curvature measurement -> **particle momentum**
- ▶ Position measurement -> **vertex reconstruction**
- ▶ NB: **Experimental magnets** are large, expensive and dangerous
 - ▶ e.g CMS 4T solenoid, diameter 6m, temperature 5K, current 18kA, stored energy 2.3GJ

▶ Calorimeters

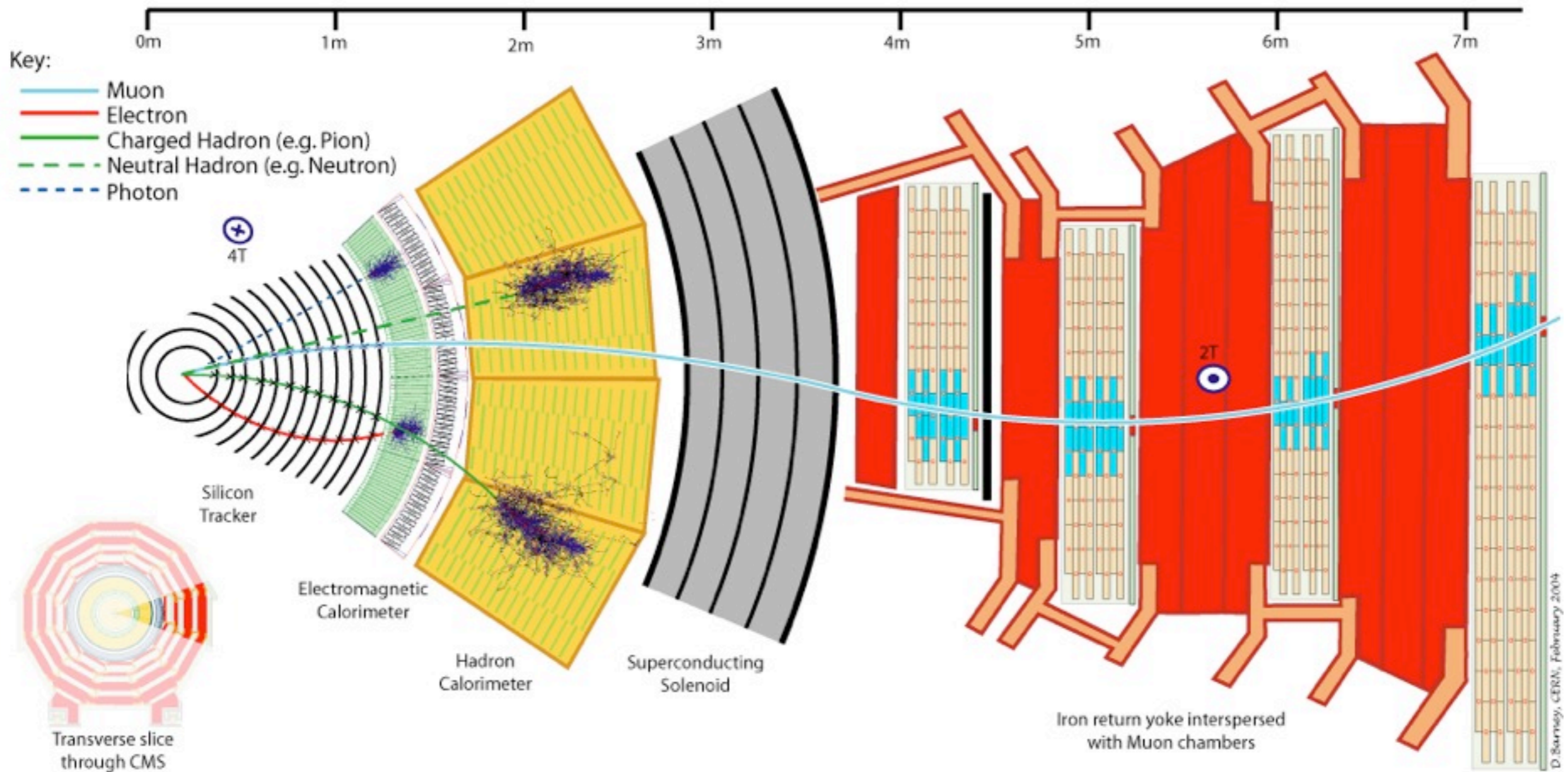
- ▶ Heavy material causes particles to deposit entire energy in a small volume
- ▶ Ionisation or light output proportional to **total energy**
- ▶ Works for charged or neutral particles (incl. photons)

▶ Particle ID

- ▶ Cherenkov / transition radiation / time-of-flight detectors
- ▶ Sensitive to velocity, and therefore **mass** (combine with momentum)

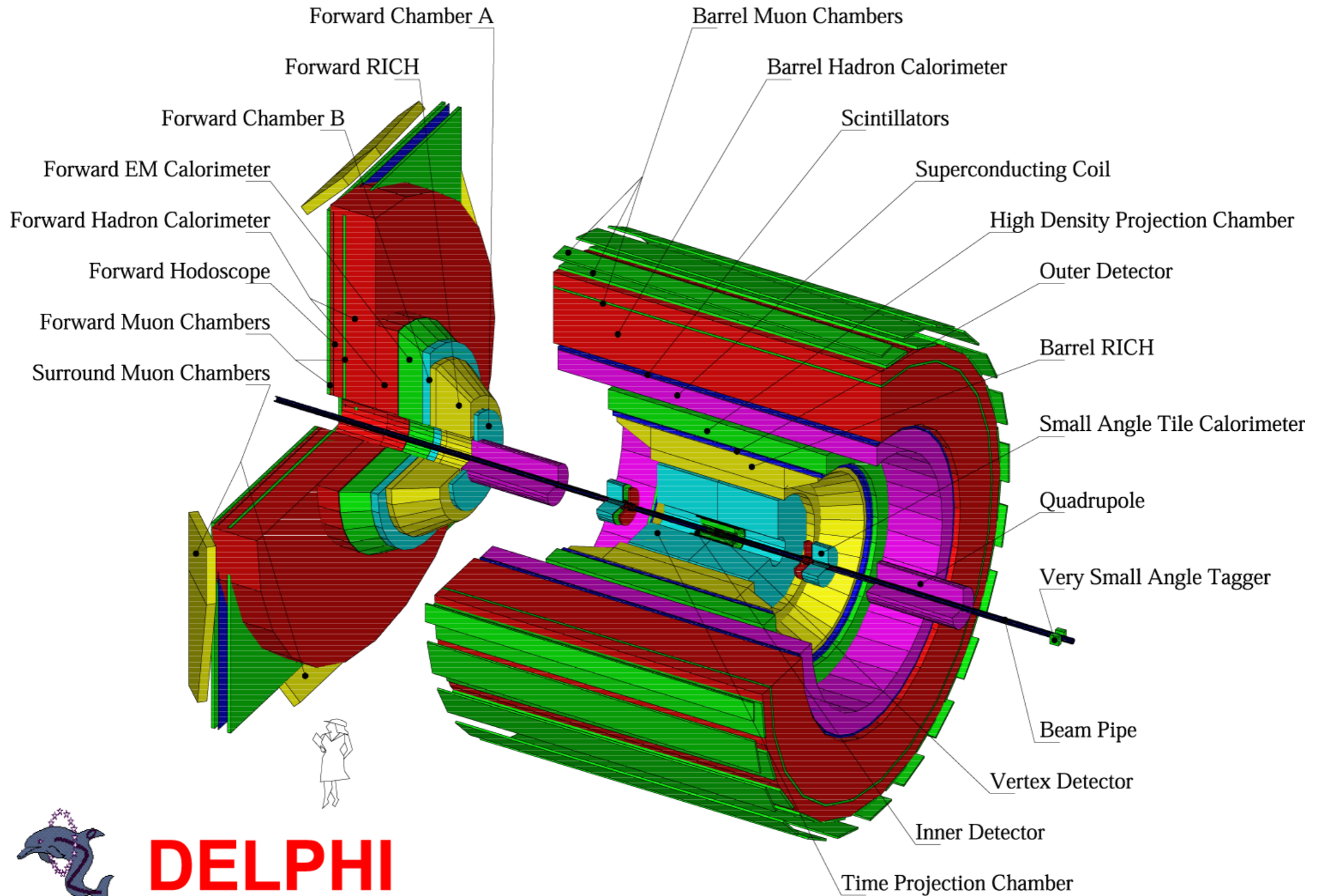
▶ Modern detectors use most or all of these techniques

Putting It All Together



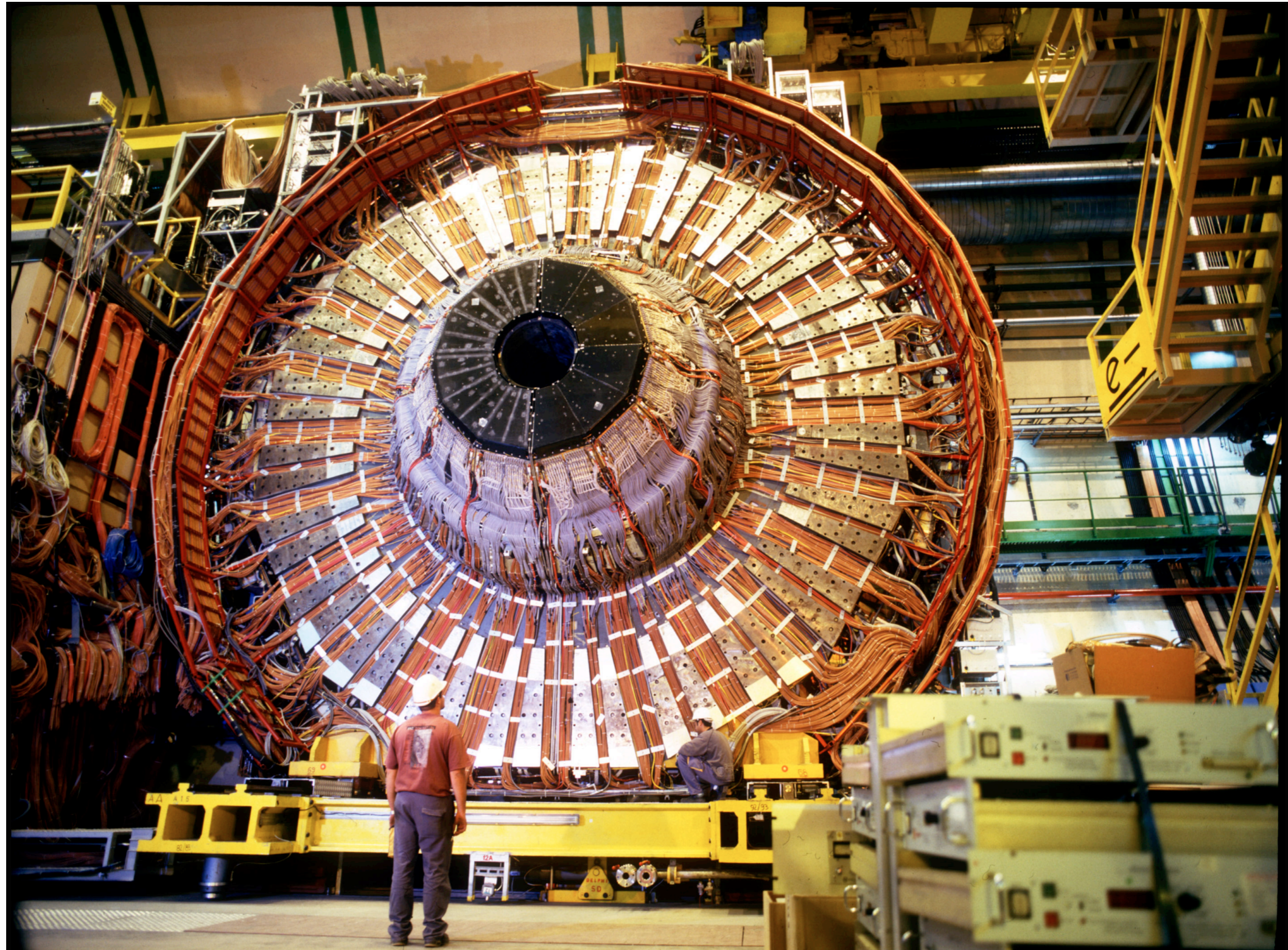
- ▶ CMS detector (GPD) layout (NB: no PID in CMS)
- ▶ Detector overall size scales with secondary particle energy

DELPHI (LEP)

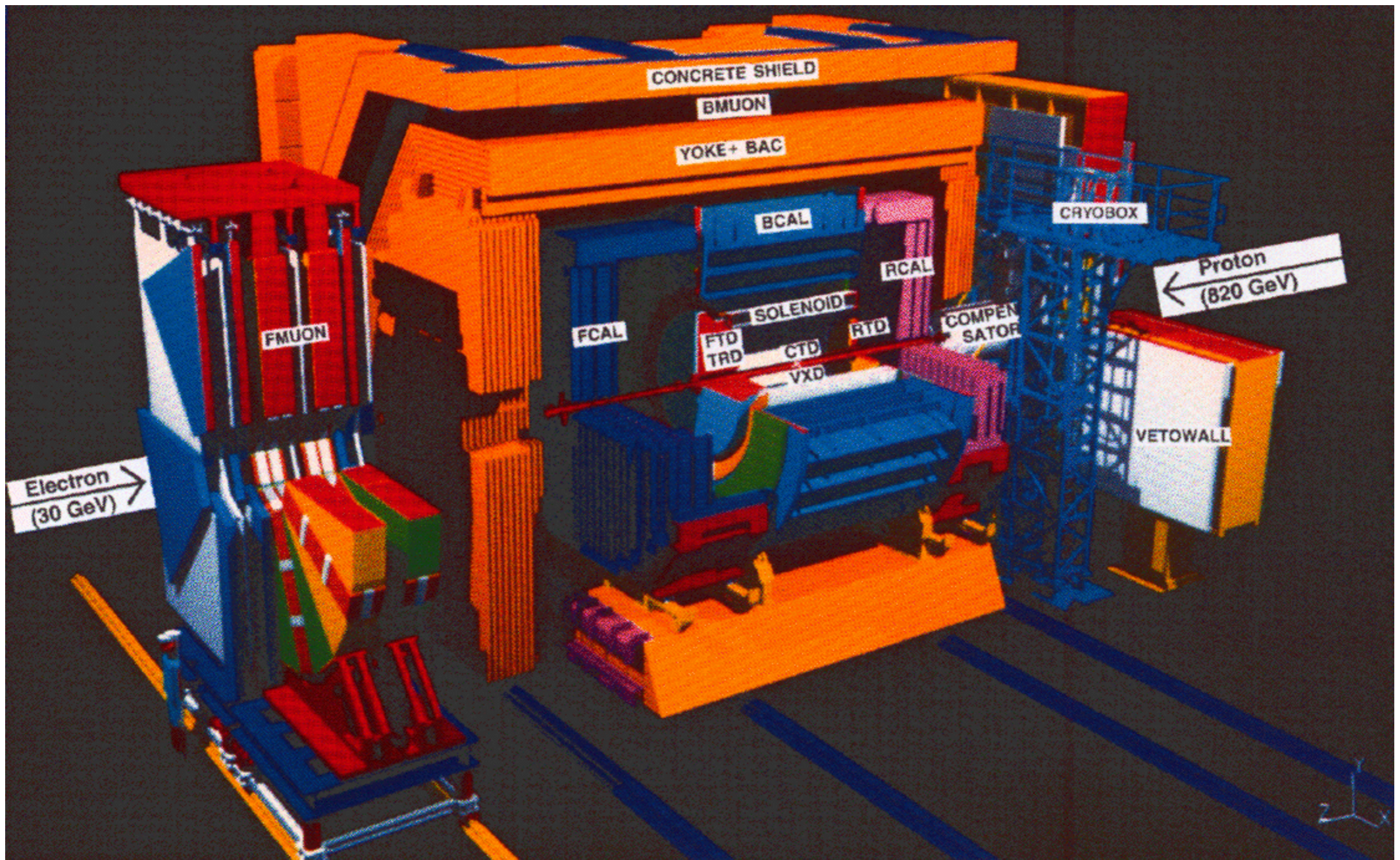


DELPHI

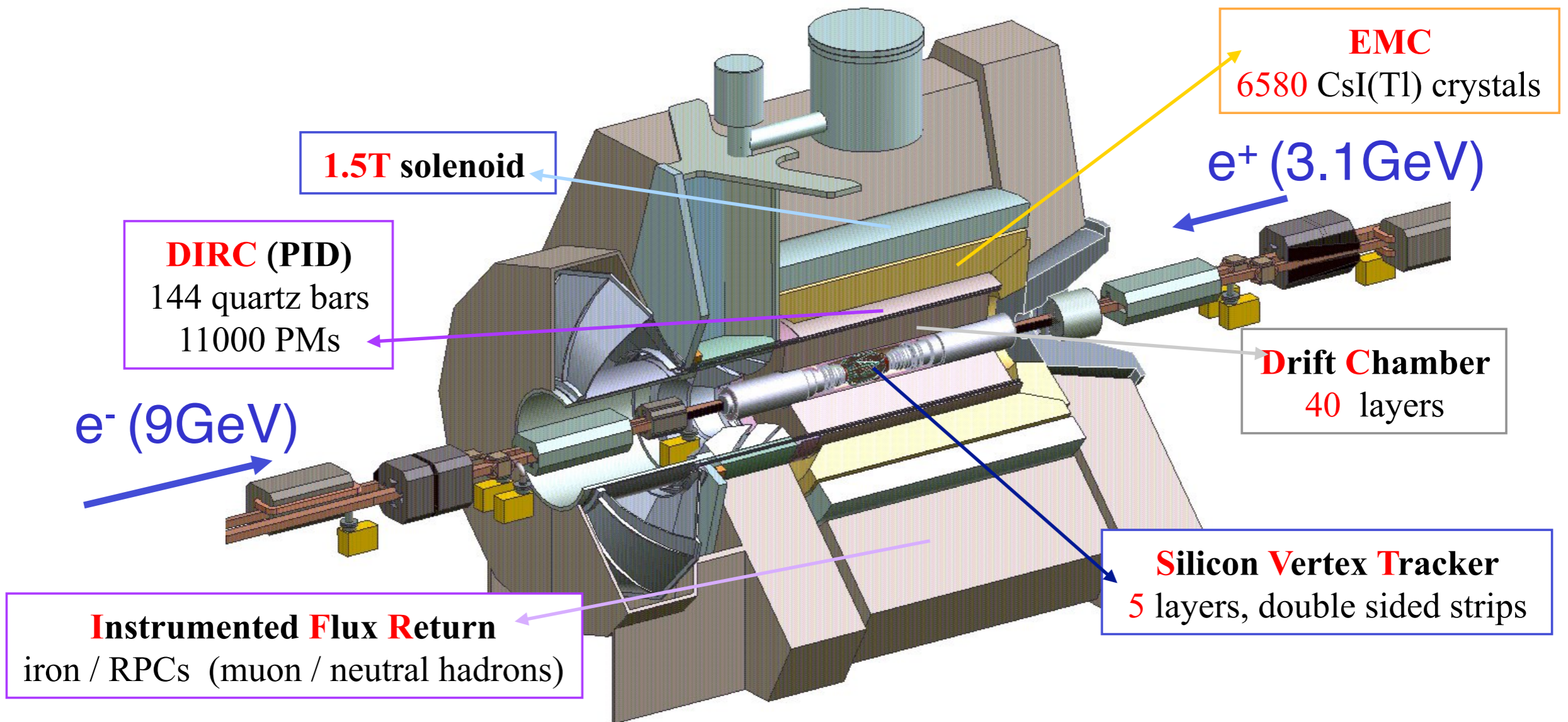
DELPHI (LEP)



ZEUS (HERA)



BaBar (PEP-II)

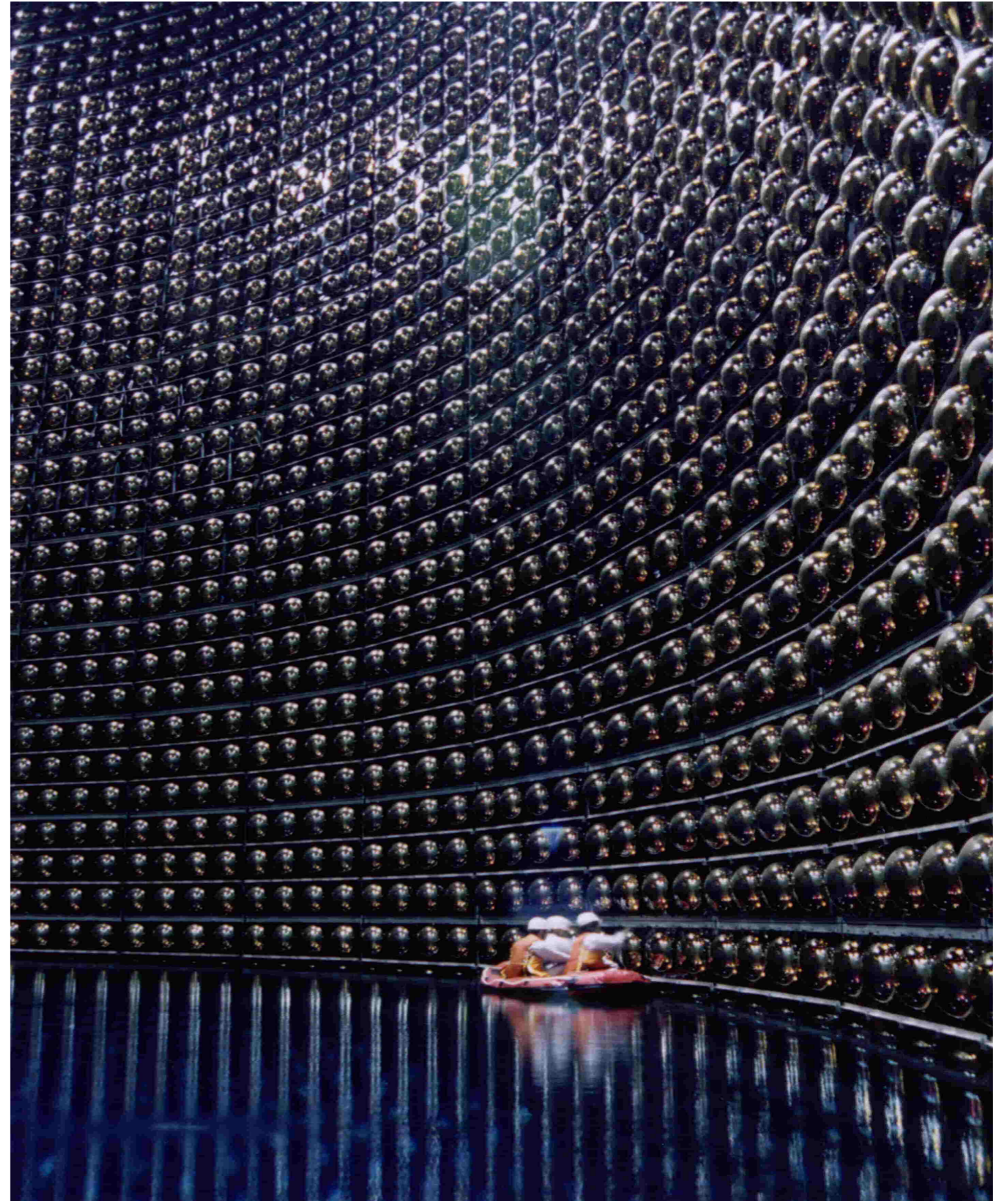
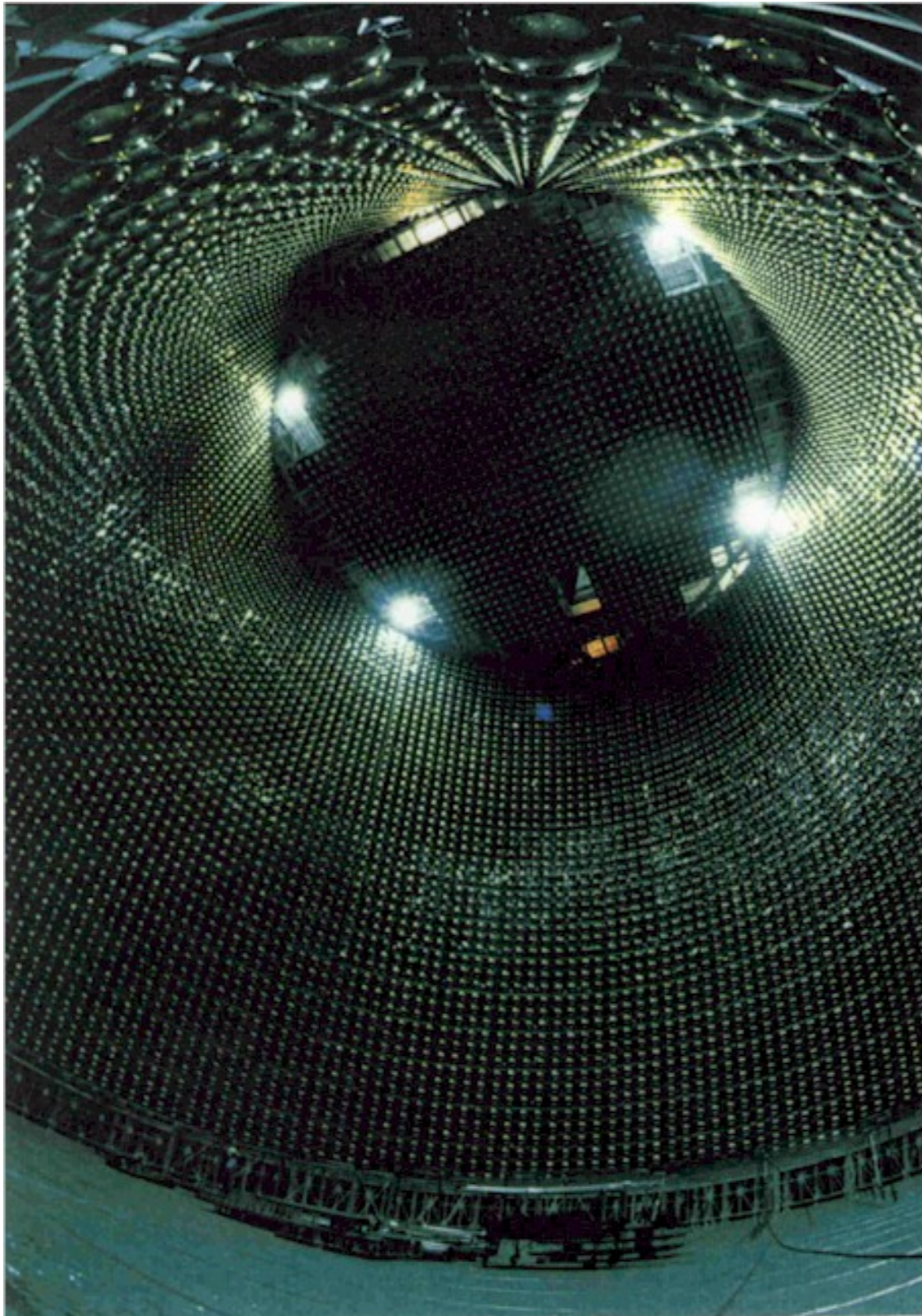


In Action

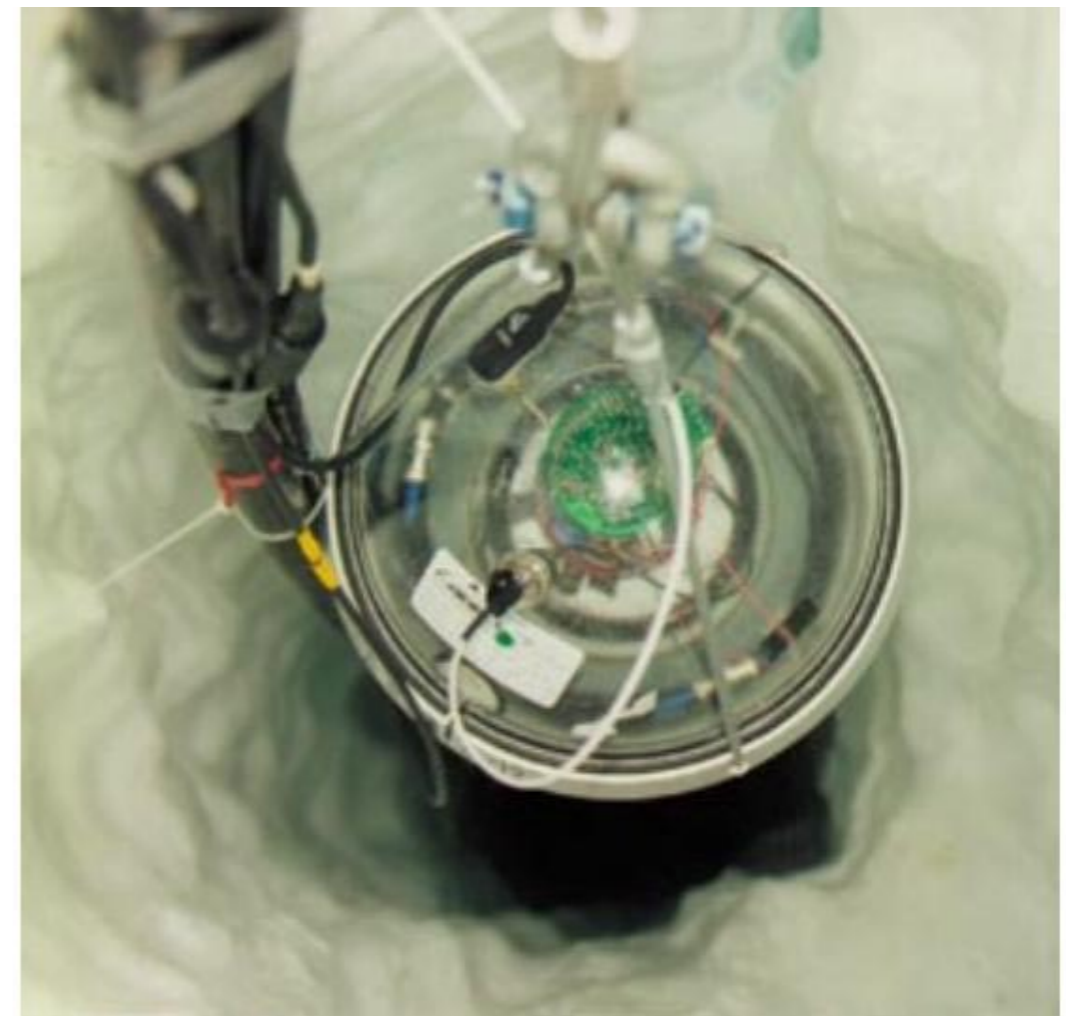
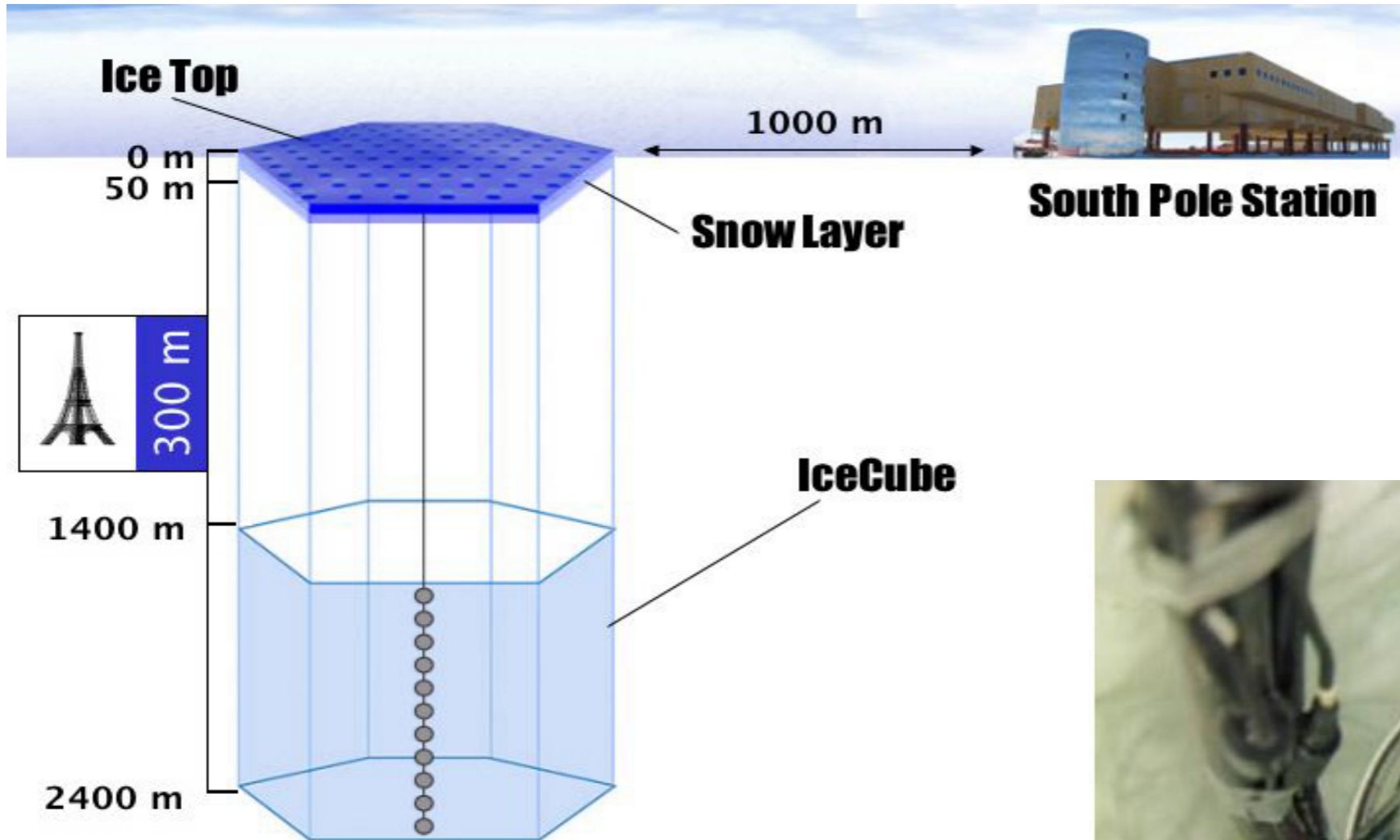
LHC Beam Pipe
27Km Long



SuperKamiokande



IceCube



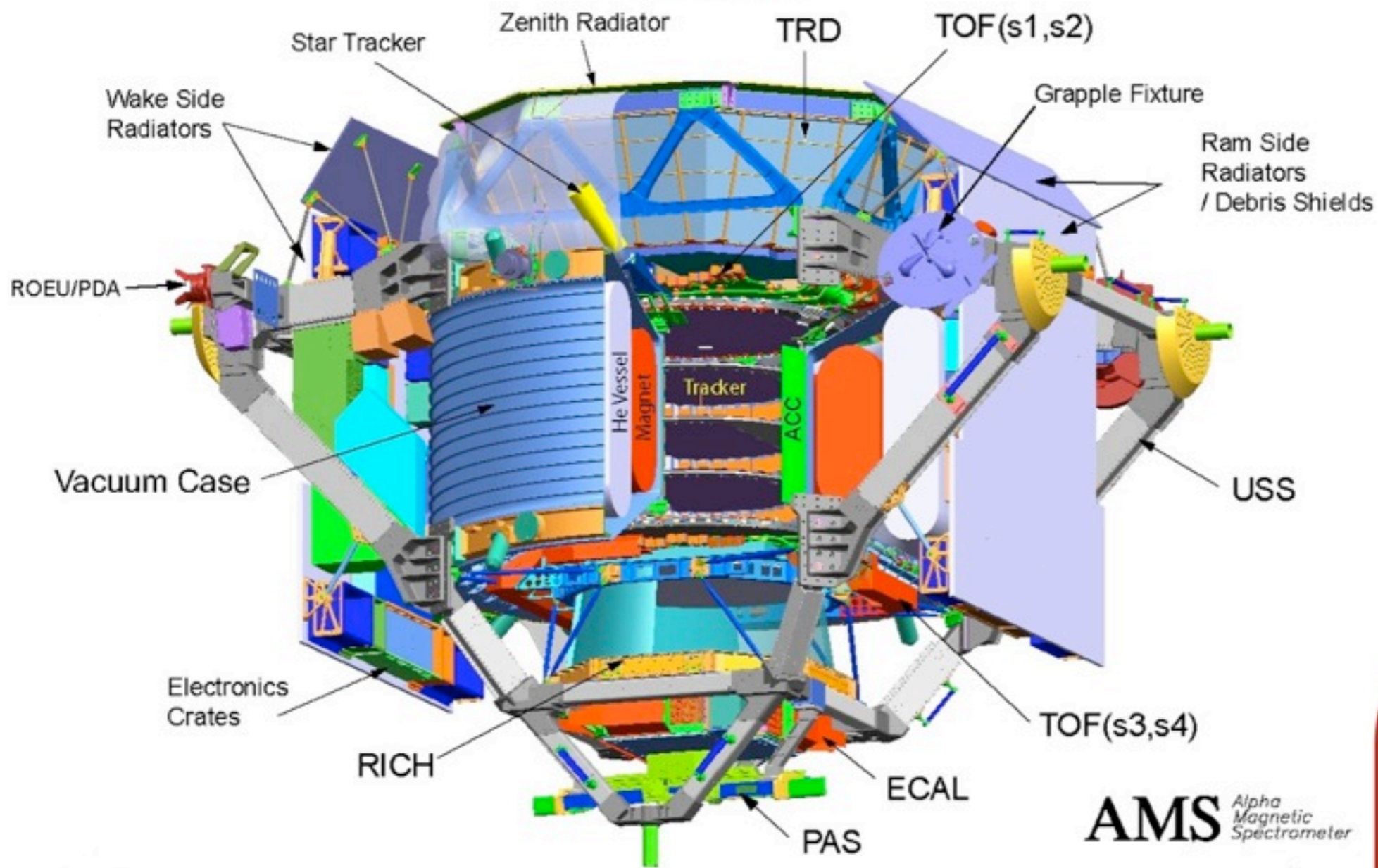
CAST



- ▶ “Light through a wall” experiment
- ▶ Uses spare LHC dipole to convert solar axions to photons

AMS

AMS 02



AMS Alpha Magnetic Spectrometer

Doing Experiments: Simulation

Decide on initial state



Build detector



Simulate performance



Record collisions



Reconstruct the events



Isolate a signal



Make measurements



Compare with theory

▶ Monte Carlo simulation

- ▶ Used for design and optimisation of detector, event selection, etc
- ▶ Also used for unfolding of detector effects
 - ▶ Though data-driven methods are usually preferable

▶ Simulation software

- ▶ De-factor standard simulation package is GEANT4
- ▶ Experiments build software on this using accurate detector descriptions
- ▶ Also have parameterised 'fast simulation' - useful for quick look at new ideas
 - ▶ GEANT sim for LHC takes ~minutes / evt

Doing Experiments: Data-Taking

Decide on initial state



Build detector



Simulate performance



Record collisions



Reconstruct the events



Isolate a signal



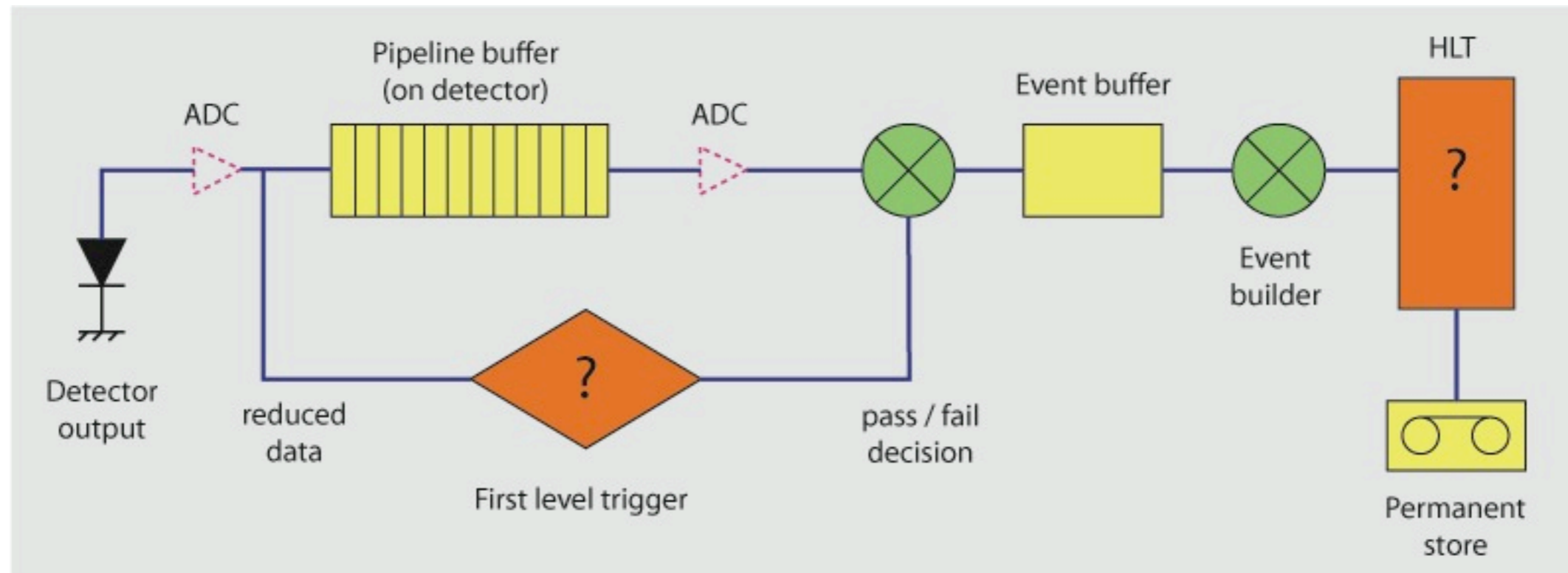
Make measurements



Compare with theory

- ▶ Transmit, digitise, record and distribute signals from detectors
- ▶ DAQ is a major challenge
 - ▶ Significant fraction of detector cost
 - ▶ High performance, reliability essential
- ▶ Data rates
 - ▶ Can be extremely high
 - ▶ e.g. LHC: $2\text{MB/evt} * 40\text{MHz}$ crossing rate = 80TByte/s or 1YB/year
 - ▶ Online event selection is required
- ▶ Processing data
 - ▶ Extremely large processing & storage is required (distributed worldwide)
 - ▶ 'Bookkeeping' is a huge task

Online Event Selection



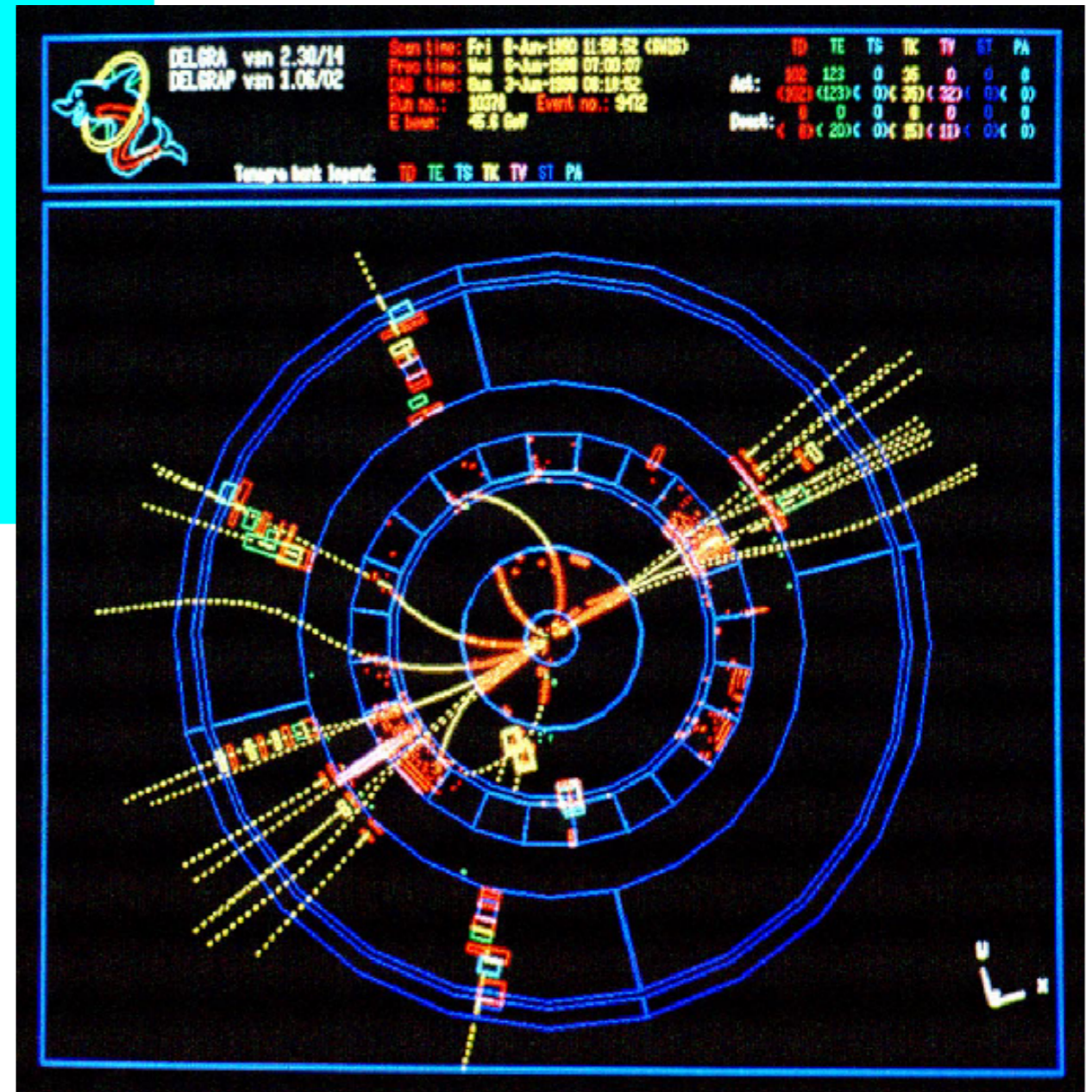
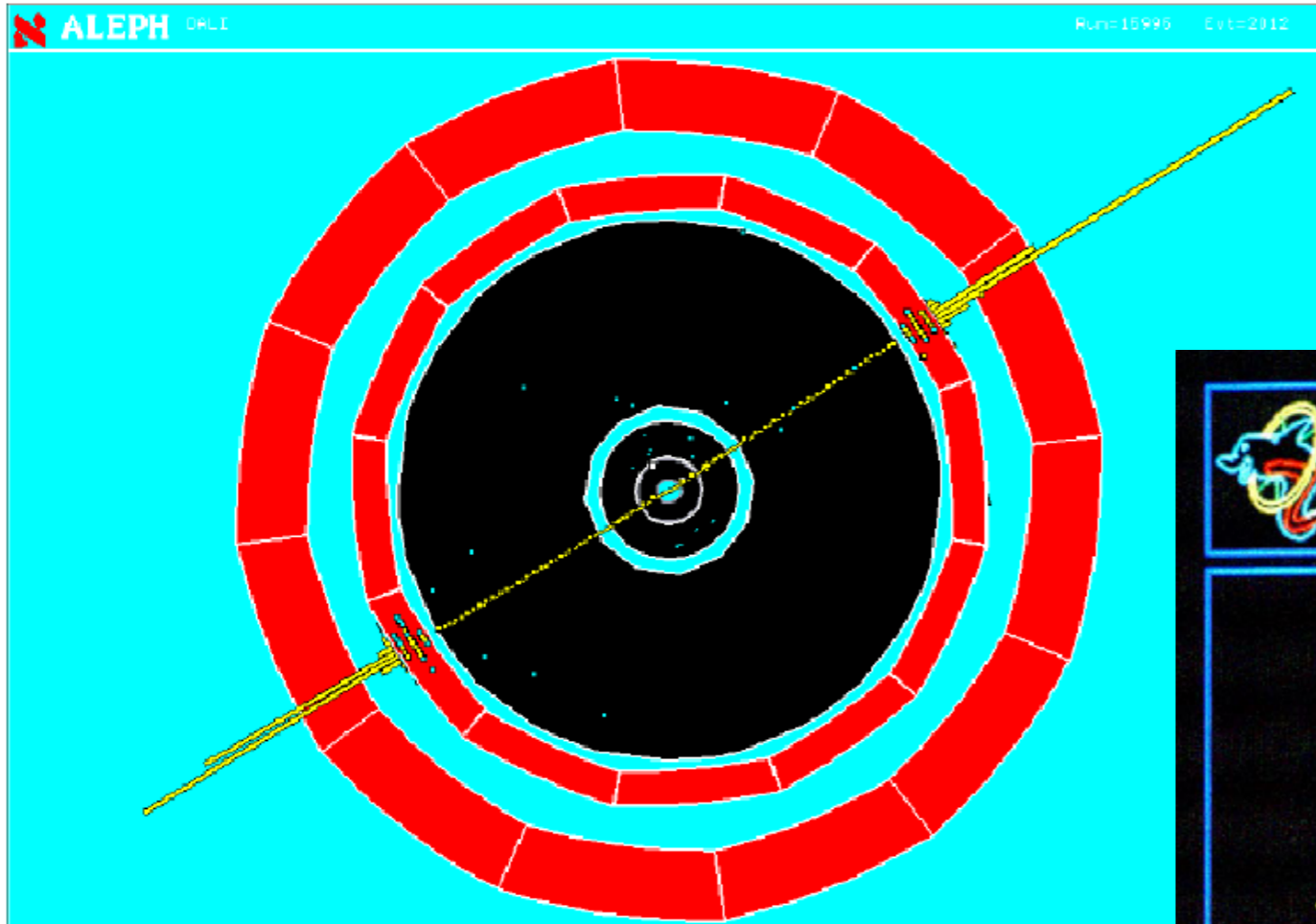
▶ Trigger system:

- ▶ Must accept all events of interest while rejecting the boring ones
- ▶ Must reduce rate of events to storage to acceptable levels
- ▶ Online decision in fixed time (few μs) – events not accepted lost for ever

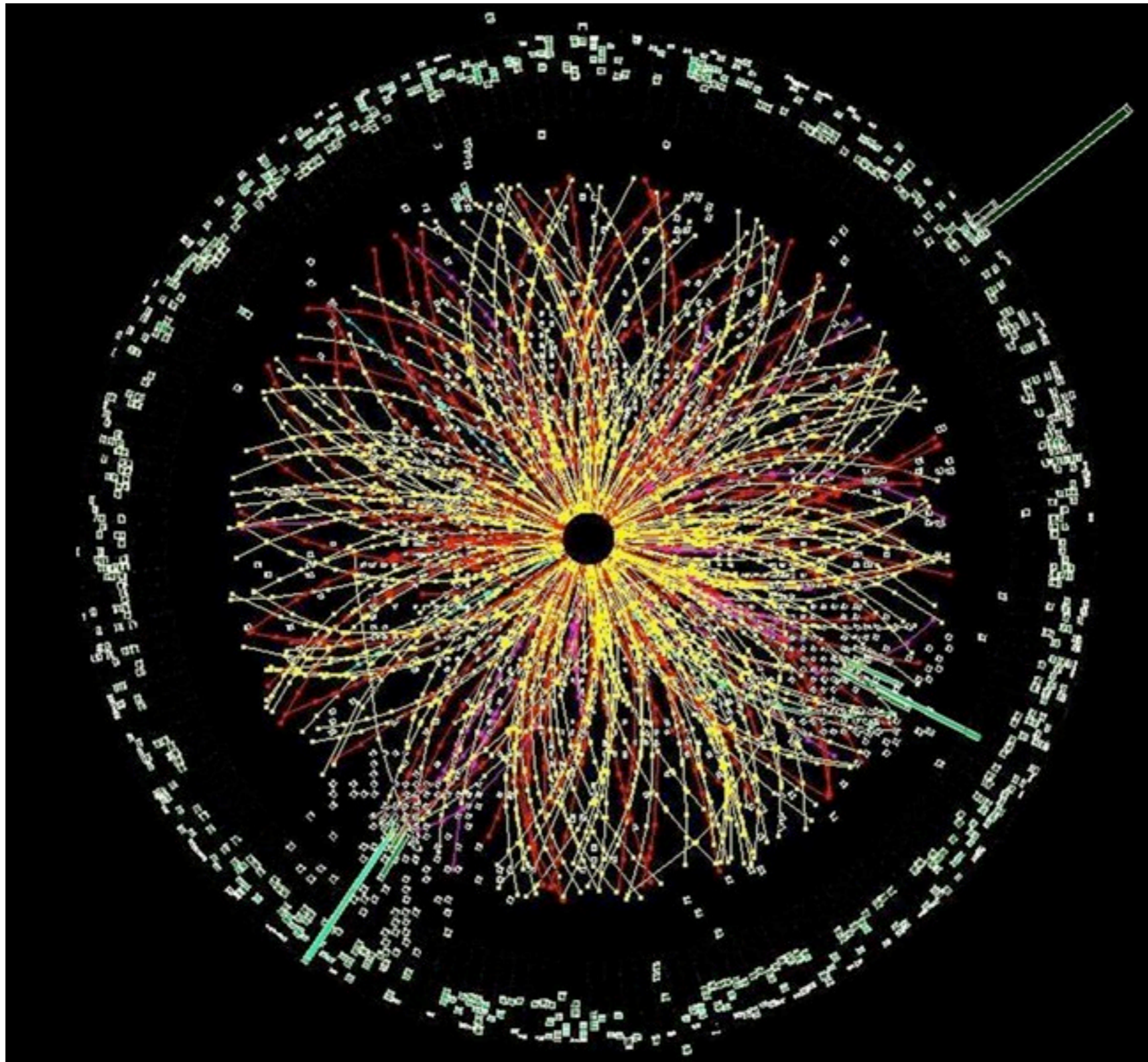
▶ Types of trigger

- ▶ First level: usually custom high-speed electronics (digital or analogue)
- ▶ Higher level: usually software on specialised or general purpose CPUs

What We 'See' - LEP

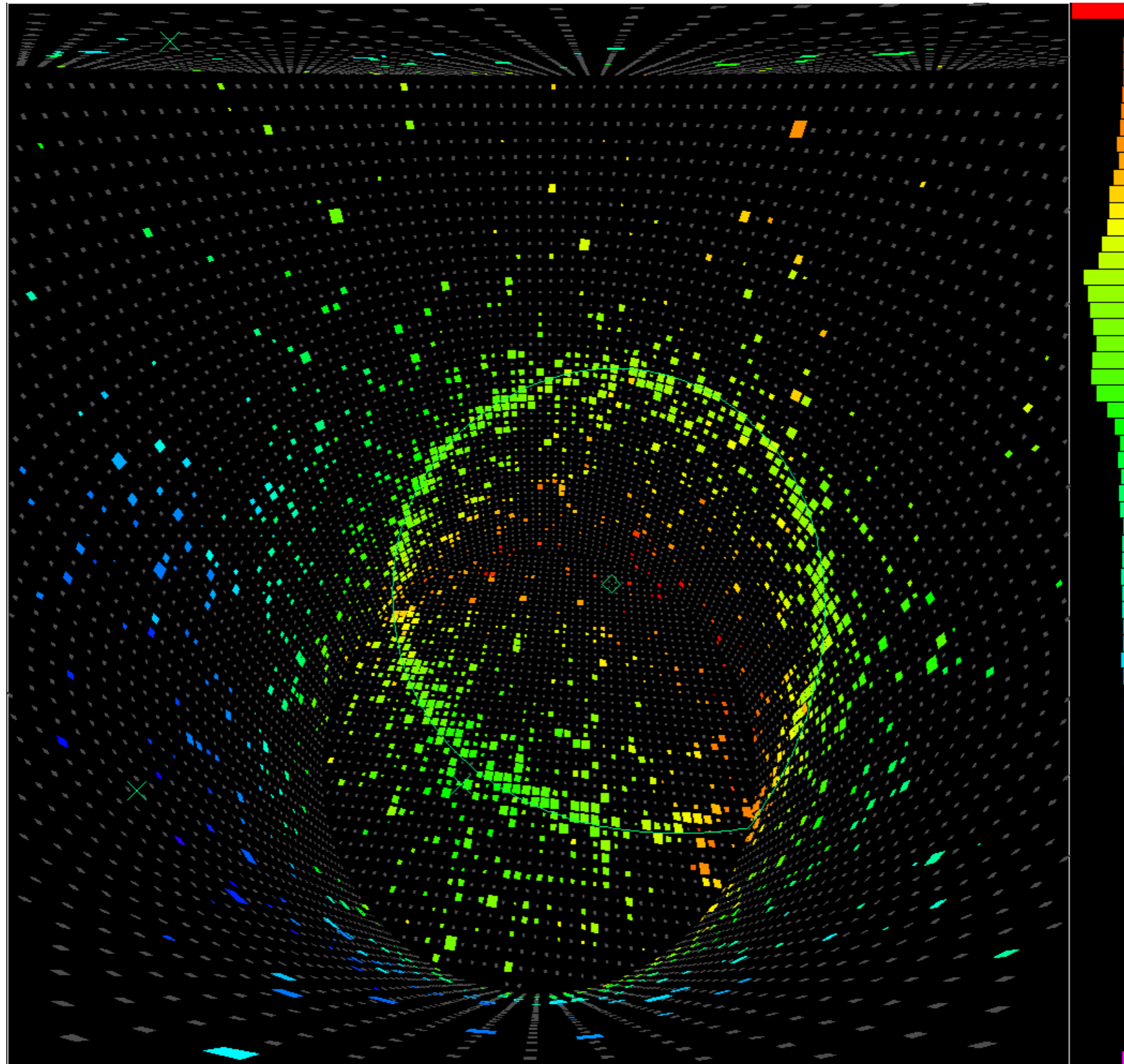


What We 'See' - LHC



Simulated crossing at $L=10^{34}$

'What We See' - SuperK



Doing Experiments: Reconstruction

Decide on initial state



Build detector



Simulate performance



Record collisions



Reconstruct the events



Isolate a signal



Make measurements



Compare with theory

▶ Pattern recognition

- ▶ Identify particle tracks / hits
- ▶ Combine information statistically to provide information on each particle

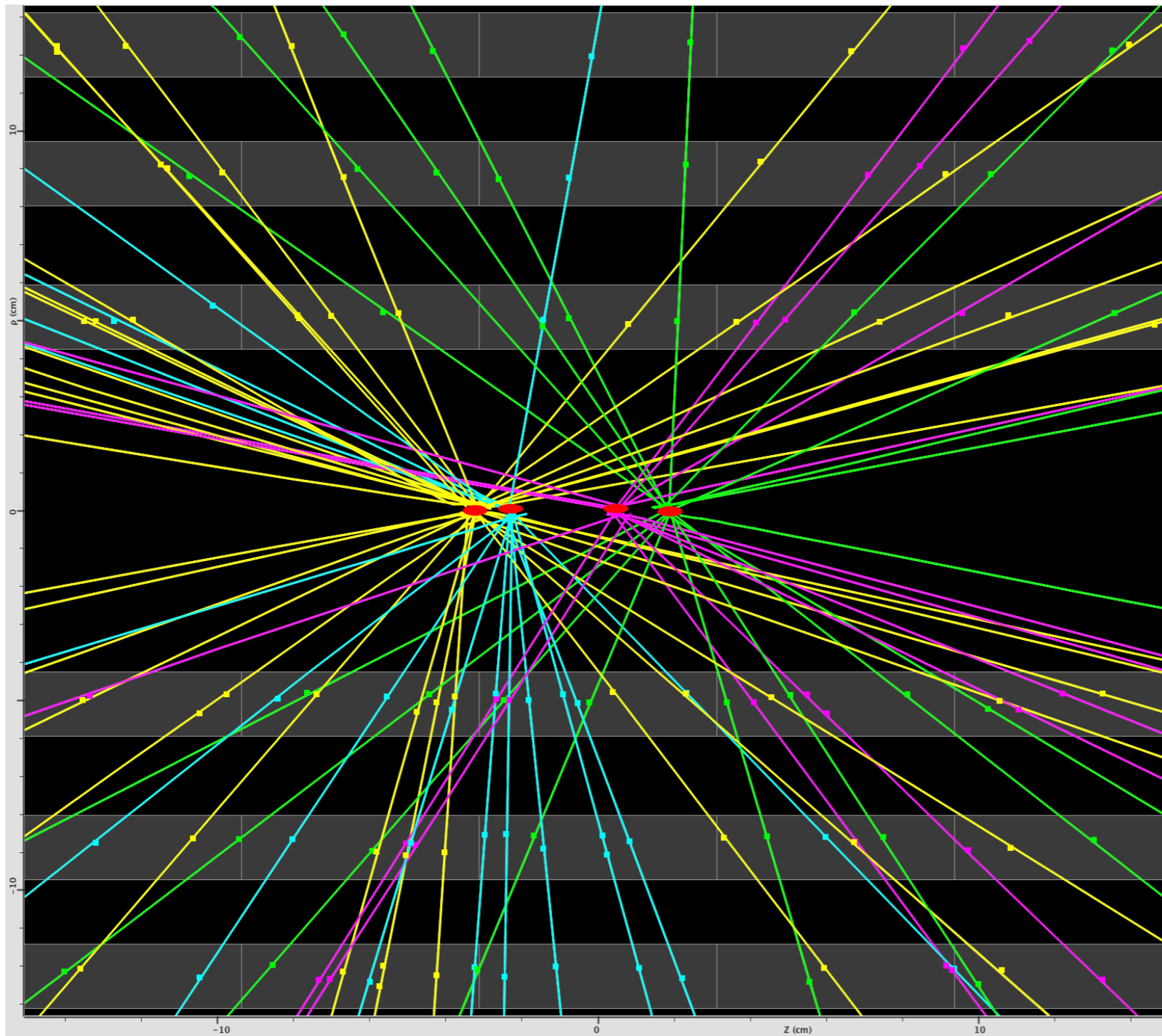
▶ Event reconstruction

- ▶ Identify primary vertex (collision pnt)
- ▶ Find secondary vertices (particle decays in flight)
- ▶ Try to identify decay topology (invariant masses, cascade decays, etc)

▶ Reconstruction software

- ▶ Often several M lines of C++
- ▶ Requires continuous tuning as conditions change in detector

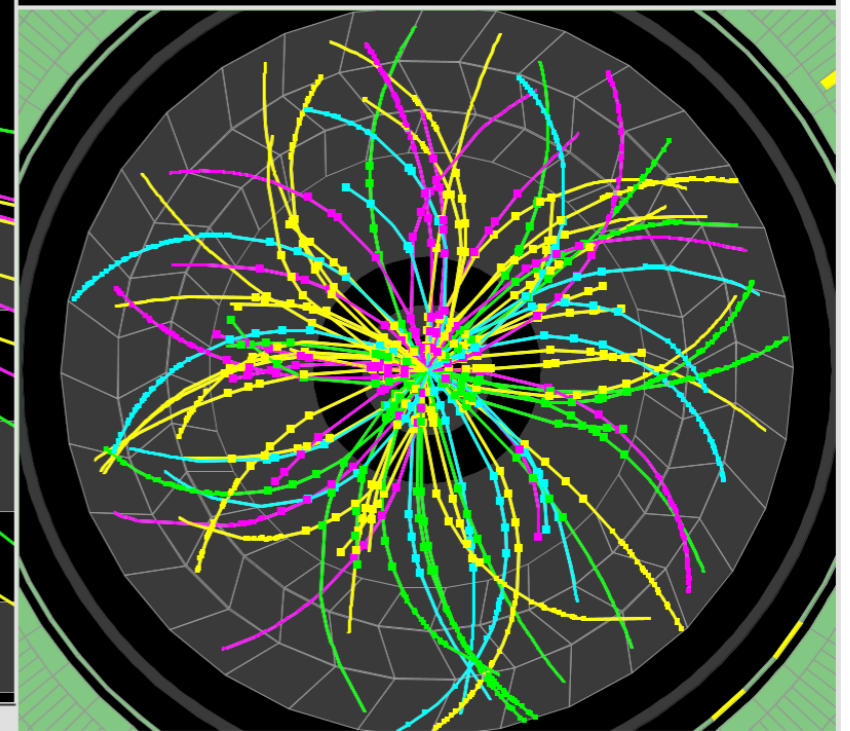
Primary Vertex Reconstruction



Run Number: 153565, Event Number: 4487360

Date: 2010-04-24 04:18:53 CEST

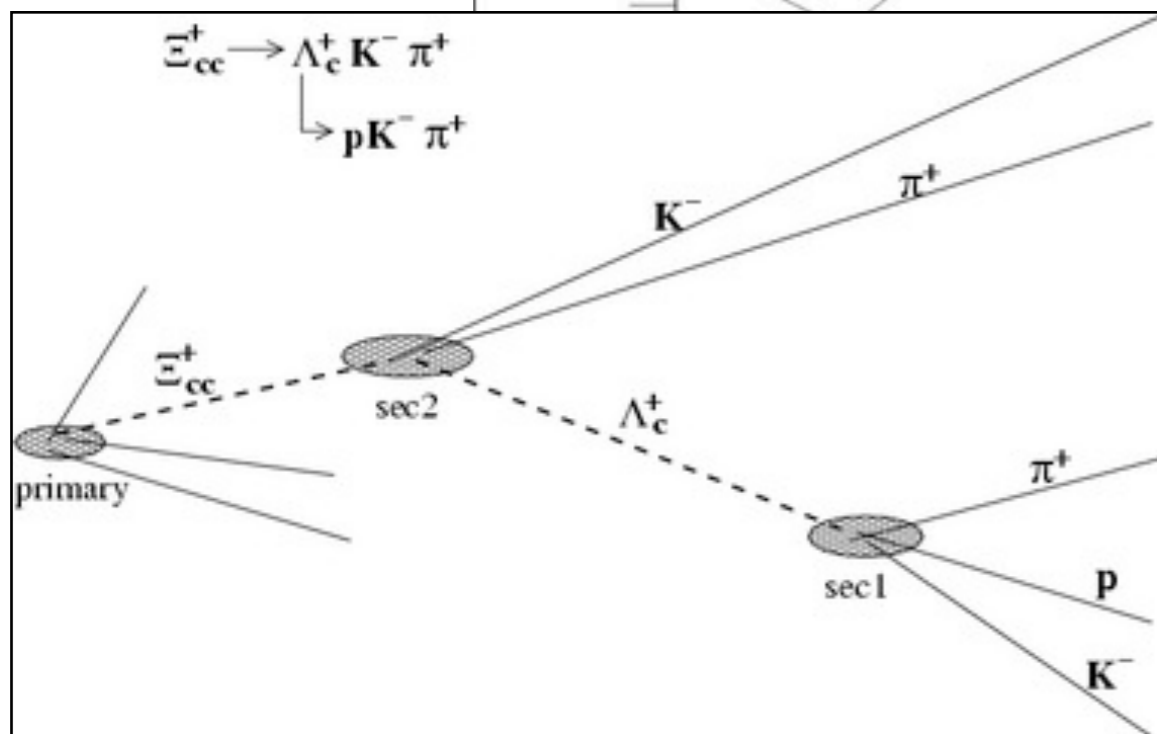
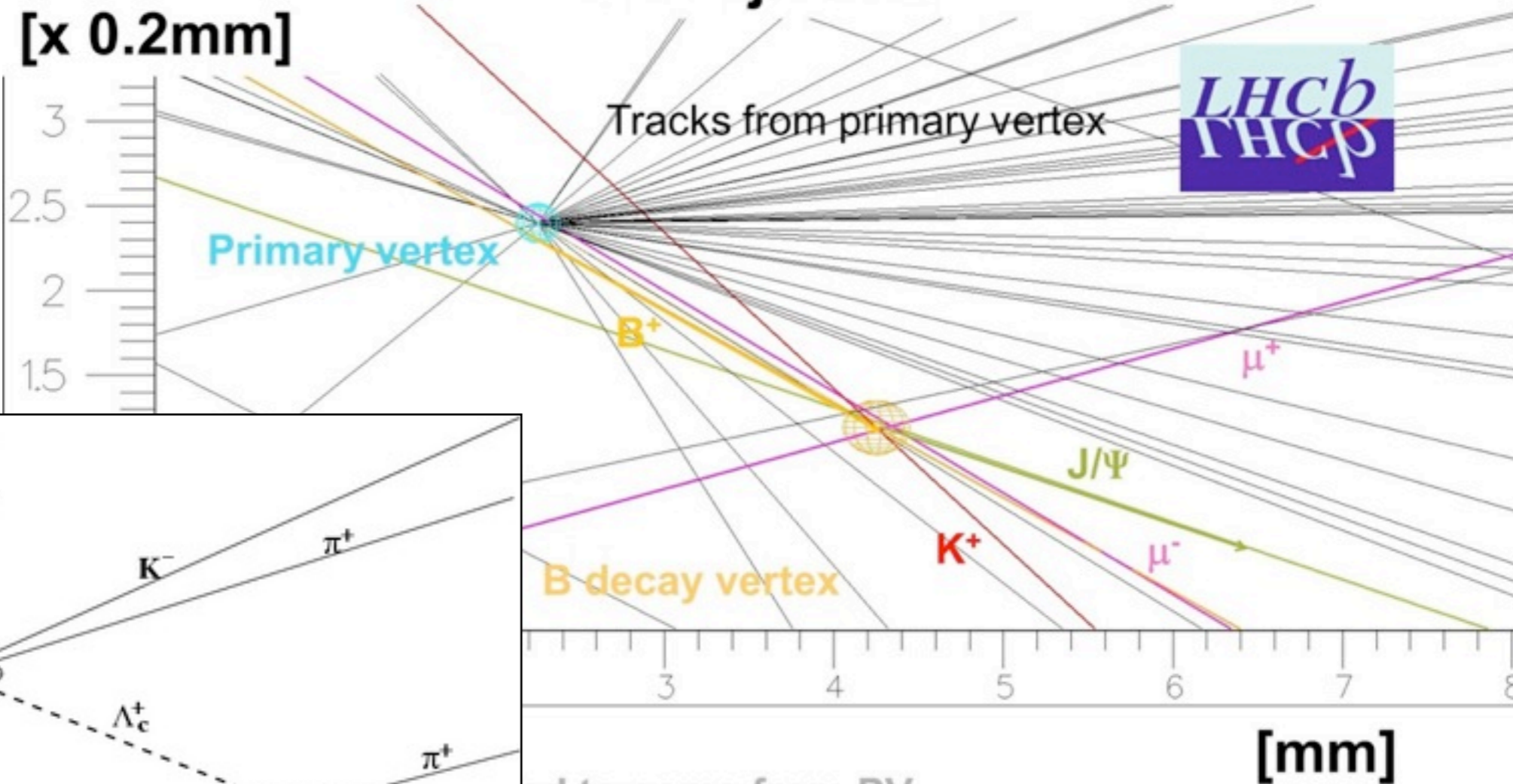
**Event with 4 Pileup Vertices
in 7 TeV Collisions**



Secondary Vertex Reconstruction



YZ Projection



Doing Experiments: Find the Signal

Decide on initial state



Build detector



Simulate performance



Record collisions



Reconstruct the events



Isolate a signal



Make measurements



Compare with theory

▶ ‘Summarise’ event

- ▶ Form invariant mass combinations
- ▶ Attempt to identify decay chain
- ▶ Extract key kinematic parameters

▶ Isolate signal

- ▶ ‘Cut and count’ traditional method
 - ▶ Make ‘cuts’ in multidimensional parameter space to enhance signal over background
- ▶ More complex methods now used
 - ▶ Neural net, decision trees, etc
- ▶ No signal is background free
 - ▶ Understanding background within the selection is of utmost importance

▶ Usually a long iterative process

- ▶ What experimental PhD students spend their time doing...

Doing Experiments: Analysis

Decide on initial state



Build detector



Simulate performance



Record collisions



Reconstruct the events



Isolate a signal



Make measurements



Compare with theory

▶ At last!

- ▶ Analyse distributions of whatever you are trying to measure...
- ▶ Correct for background contamination
- ▶ Write the paper.

▶ Never so simple

- ▶ Is the result significant? How significant? Does it mean *anything*?
- ▶ Need to carefully assess statistical and systematic errors
- ▶ Complex multivariate statistics now commonplace in HEP.

▶ Usually turns out that...

- ▶ The result is a statistical limit on observation of some event class

Statistics: Advice for the Unwary

- ▶ Statistics in experimental HEP
 - ▶ An essential tool in producing information from the data
 - ▶ Typically not well understood
 - ▶ By most readers of experimental papers, and many writers
- ▶ A quiet revolution in the last ~15 years
 - ▶ Statistics now generally done ‘properly’ (by HEP standards)
 - ▶ A small industry of experts has grown up.
 - ▶ But... one still hears of ‘2-sigma exclusions’ and ‘3-sigma observations’
- ▶ How to interpret experimental limits
 - ▶ Quite often, you simply can’t without additional information
 - ▶ And certainly not from plots shown at conferences
 - ▶ Many assumptions are ‘in the small print’
 - ▶ In particular, take great care when:
 - ▶ Comparing a result to the predictions of a model
 - ▶ Comparing or combining results from different experiments or runs (correlated errors, different assumptions, etc)

Wise Comment

Wise Comment

“If your result needs a statistician, you should design a better experiment”

- Rutherford

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“If your result needs a statistician, you should design a better experiment”

- Rutherford

“We haven't got the money, so we've got to think!”

- Rutherford

Wise Comment

“If your result needs a statistician, you should design a better experiment”

- Rutherford

“We haven't got the money, so we've got to think!”

- Rutherford

“Don't let me catch anyone talking about the Universe in my department”

- Rutherford

Comparison with Theory

Decide on initial state



Build detector



Simulate performance



Record collisions



Reconstruct the events



Isolate a signal



Make measurements



Compare with theory

▶ Typical methodology

- ▶ Use theoretical model to predict event yield of a given type
- ▶ Requires assumptions about PDFs, etc, as well as hard collision model
- ▶ Fold in detector resolutions, efficiencies
- ▶ Estimate compatibility with data

▶ Event generators

- ▶ The lingua franca between theory and experiment
- ▶ If you want a model tested, make sure there is a generator implementation
- ▶ A good knowledge of theory uncertainties will be required
 - ▶ In the limit of decent statistics

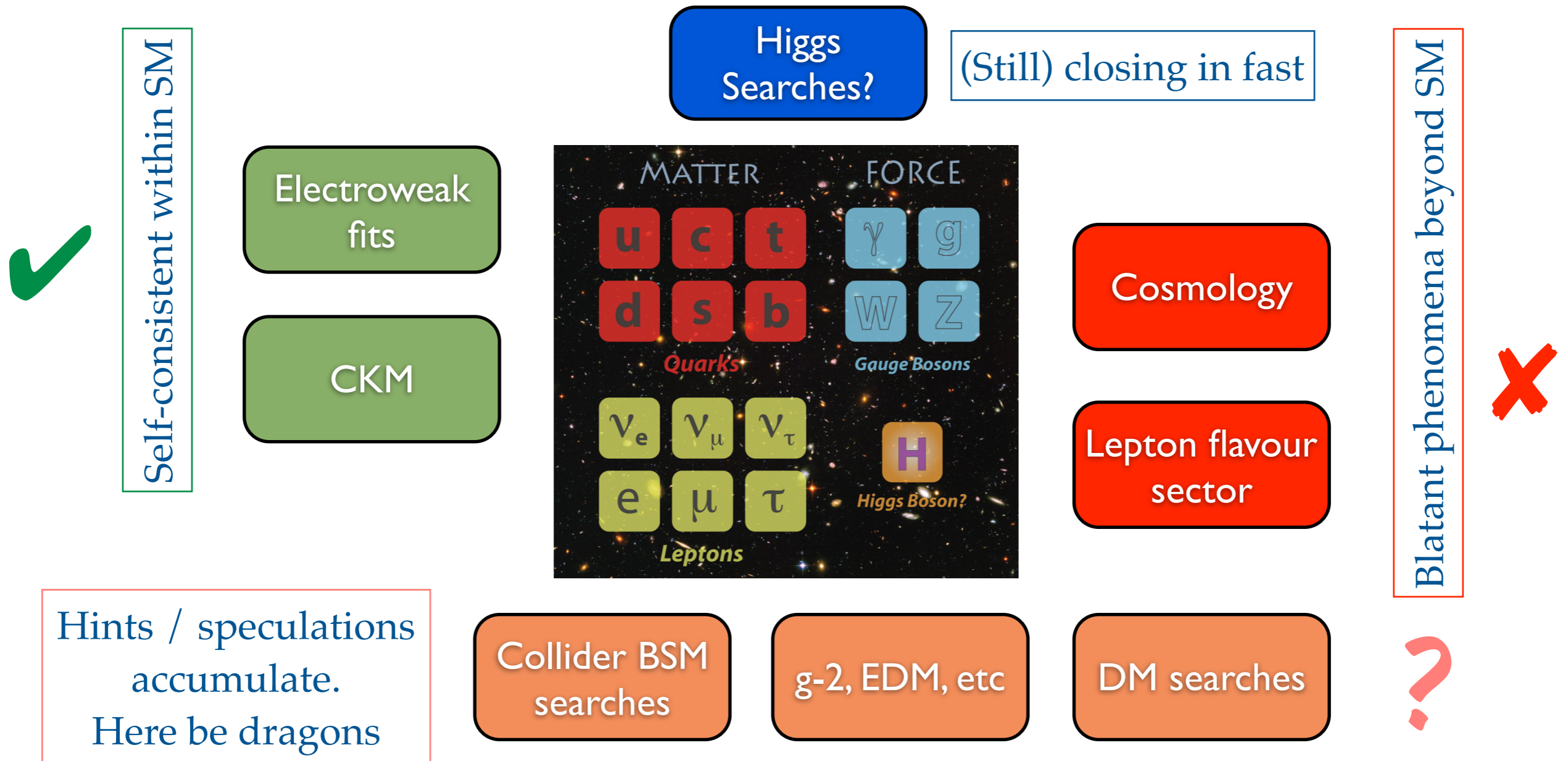
Your Mission

Should you choose to accept it

- ▶ Complete confirmation of the Standard Model
 - ▶ i.e. find the Higgs boson – or whatever else does the job
 - ▶ Go beyond the standard model
 - ▶ Understand hierarchy problem, i.e. SUSY – or whatever else does the job
 - ▶ Search for additional structure above 1TeV
 - ▶ Gauge extensions, 4th generation, compositeness, leptoquarks, etc
 - ▶ Relate all of this to cosmology
 - ▶ Antimatter asym., CMWB, dark matter, dark energy, proton lifetime, etc
 - ▶ It is unlikely to happen in this order
 - ▶ No plan survives contact with enemy
-

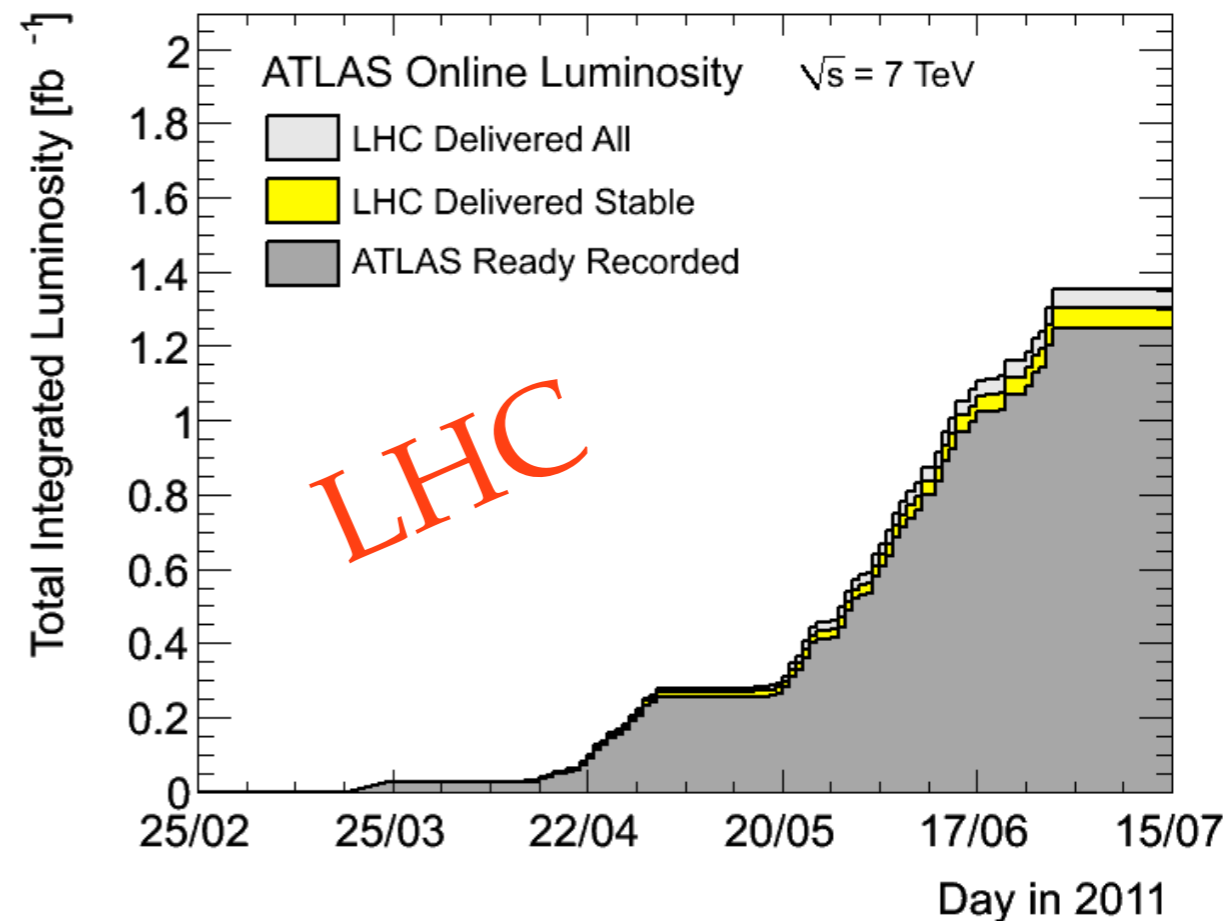
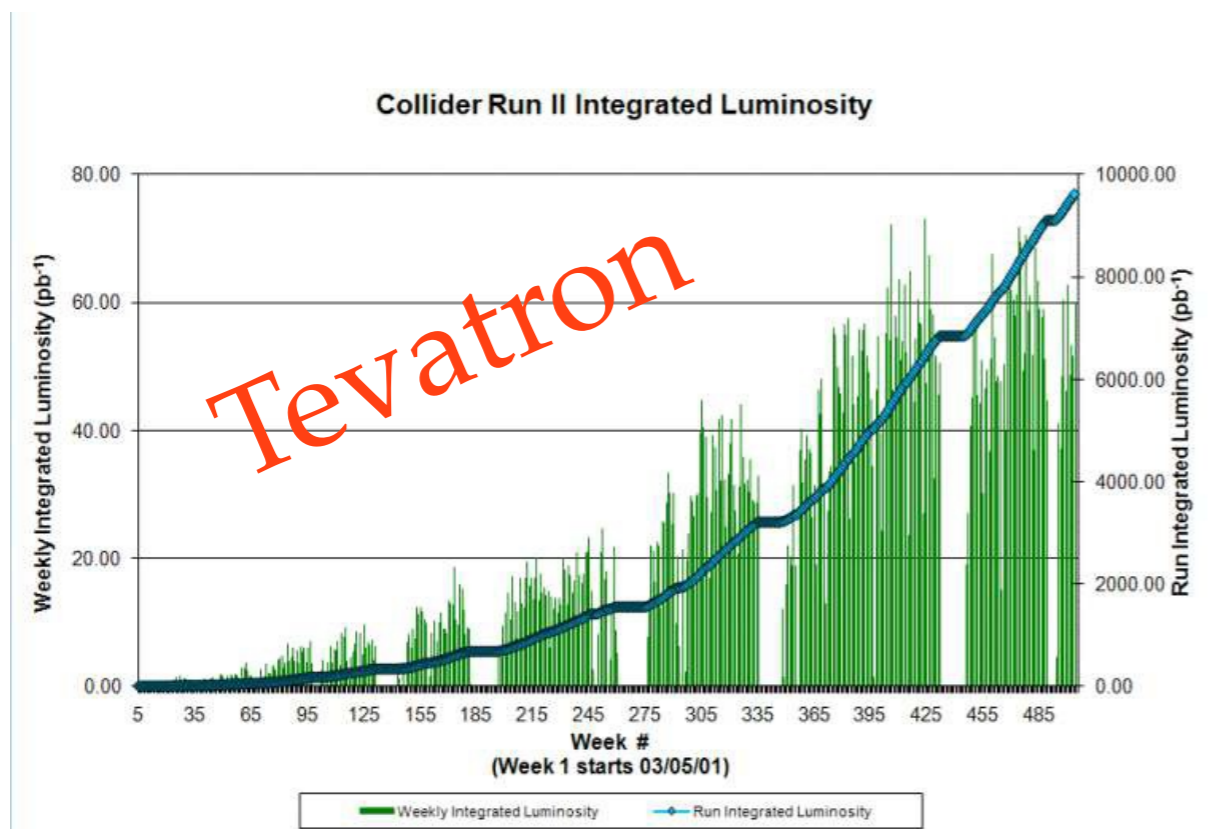
This slide will self-destruct in ten seconds

Constraining the Standard Model



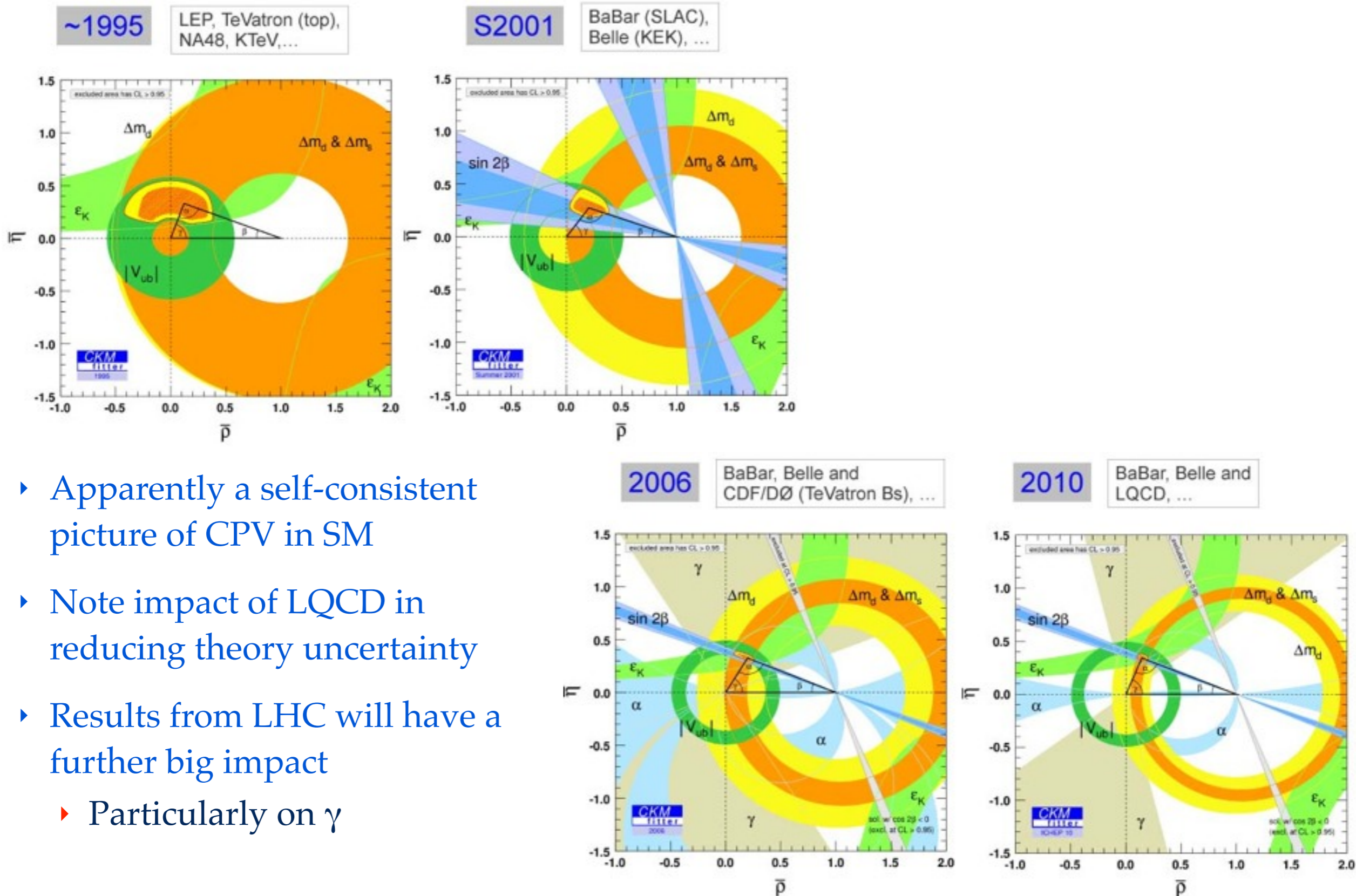
- ▶ Plus, a wealth of clear theoretical arguments for BSM physics
- ▶ For a full review, see proceedings of EPS2011, LP2011, etc

Tevatron vs LHC vs ...



- ▶ Asymmetries accessible
- ▶ Cleaner environment
- ▶ Very well understood detectors
- ▶ Many other complementary & competitive facilities
 - ▶ Non-accelerator searches; 'intensity frontier'; neutrino beams, etc
- ▶ Symmetric environment
- ▶ Increasingly dirty environment
- ▶ Higher energy – decisive advantage

Quark Flavour Sector



- ▶ Apparently a self-consistent picture of CPV in SM
- ▶ Note impact of LQCD in reducing theory uncertainty
- ▶ Results from LHC will have a further big impact
 - ▶ Particularly on γ

Lepton Flavour Sector

Mixing angles

Atmospheric (SK)
Accelerators (K2K, Minos)

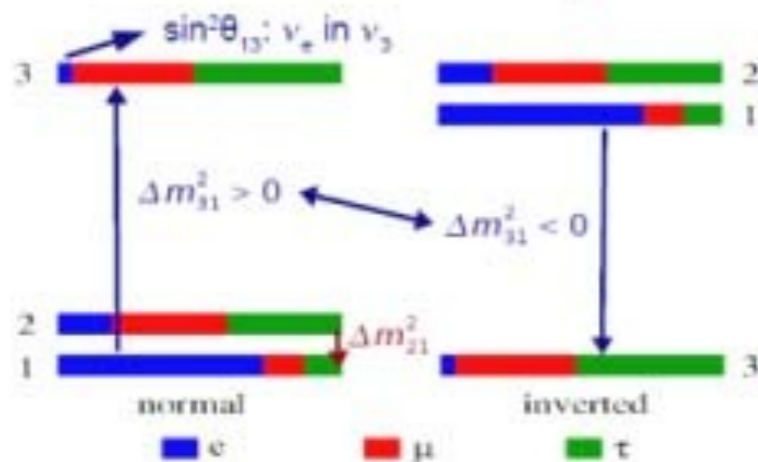
Reactors (CHOOZ)
Accelerators (JPARC)

Solar (SNO, SK)
Reactors (KamLAND)

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos\theta_{23} & \sin\theta_{23} \\ 0 & -\sin\theta_{23} & \cos\theta_{23} \end{pmatrix} \times \begin{pmatrix} \cos\theta_{13} & 0 & e^{-i\delta_{CP}} \sin\theta_{13} \\ 0 & 1 & 0 \\ -e^{i\delta_{CP}} \sin\theta_{13} & 0 & \cos\theta_{13} \end{pmatrix} \times \begin{pmatrix} \cos\theta_{12} & \sin\theta_{12} & 0 \\ -\sin\theta_{12} & \cos\theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \times \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{i\alpha/2} & 0 \\ 0 & 0 & e^{i\alpha/2+i\beta} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$\tan^2\theta_{23} = 1.0 \quad 0.3$
 $\sin^2 2\theta_{13} < 0.16$
 $\tan^2\theta_{12} = 0.47 \quad 0.05$
 α, β : CP Majorana phase
 δ_{CP} : CP Dirac phase

Mass hierarchy



$$\Delta m_{sol}^2 = \Delta m_{12}^2 = (7.58 \pm 0.21) 10^{-5} \text{ eV}^2$$

$$\Delta m_{atm}^2 = \Delta m_{31}^2 = (2.2 \pm 0.2) 10^{-3} \text{ eV}^2$$

Absolute neutrino mass

Beta decay : $|m_\nu| = \sum |U_{ei}| m_i < 2.6 \text{ eV (90 \% CL)}$

Double beta : $\langle m_{ee} \rangle = |\sum U_{ei}^2 m_i| < 0.3 - 0.7 \text{ eV (95 \% CL)}$

Cosmology : $m_\nu = m_1 + m_2 + m_3 < 0.5 - 1 \text{ eV (95 \% CL)}$

Lepton Flavour Sector

Mixing angles

Atmospheric (SK)
Accelerators (K2K, Minos)

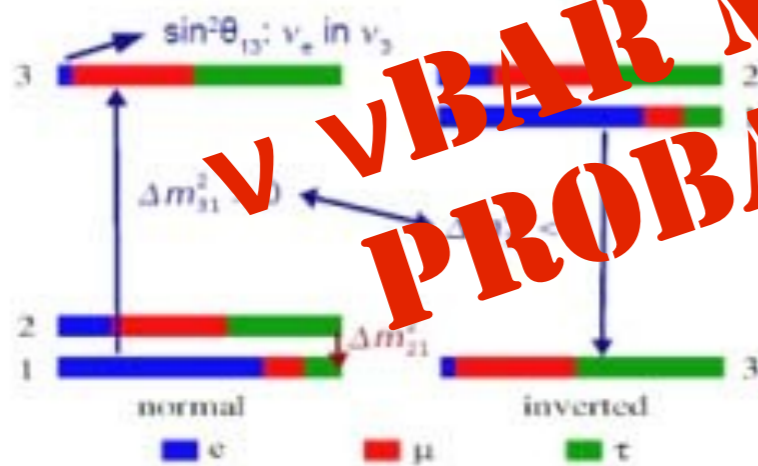
Reactors (CHOOZ)
Accelerators (JPARC)

Solar (SNO, SK)
Reactors (KamLAND)

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos\theta_{23} & \sin\theta_{23} \\ 0 & -\sin\theta_{23} & \cos\theta_{23} \end{pmatrix} \times \begin{pmatrix} \cos\theta_{13} & 0 & e^{-i\delta_{CP}} \sin\theta_{13} \\ 0 & 1 & 0 \\ -e^{i\delta_{CP}} \sin\theta_{13} & 0 & \cos\theta_{13} \end{pmatrix} \times \begin{pmatrix} \cos\theta_{12} & \sin\theta_{12} & 0 \\ -\sin\theta_{12} & \cos\theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \times \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{i\alpha/2} & 0 \\ 0 & 0 & e^{-i\alpha/2} \end{pmatrix} \times \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$\tan^2\theta_{23} = 1.0$ 0.3 $\sin^2\theta_{13} < 0.16$ $\tan^2\theta_{12} = 0.47$ 0.95 α : Majorana phase
 δ_{CP} : CP Dirac phase

Mass hierarchy



$$\Delta m_{sol}^2 = \Delta m_{12}^2 = (7.58 \pm 0.21) 10^{-5} \text{ eV}^2$$

$$\Delta m_{atm}^2 = \Delta m_{31}^2 = (2.2 \pm 0.2) 10^{-3} \text{ eV}^2$$

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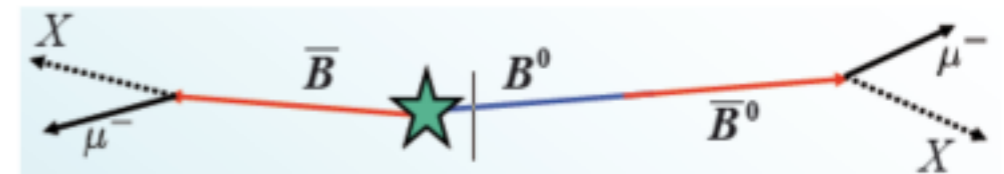
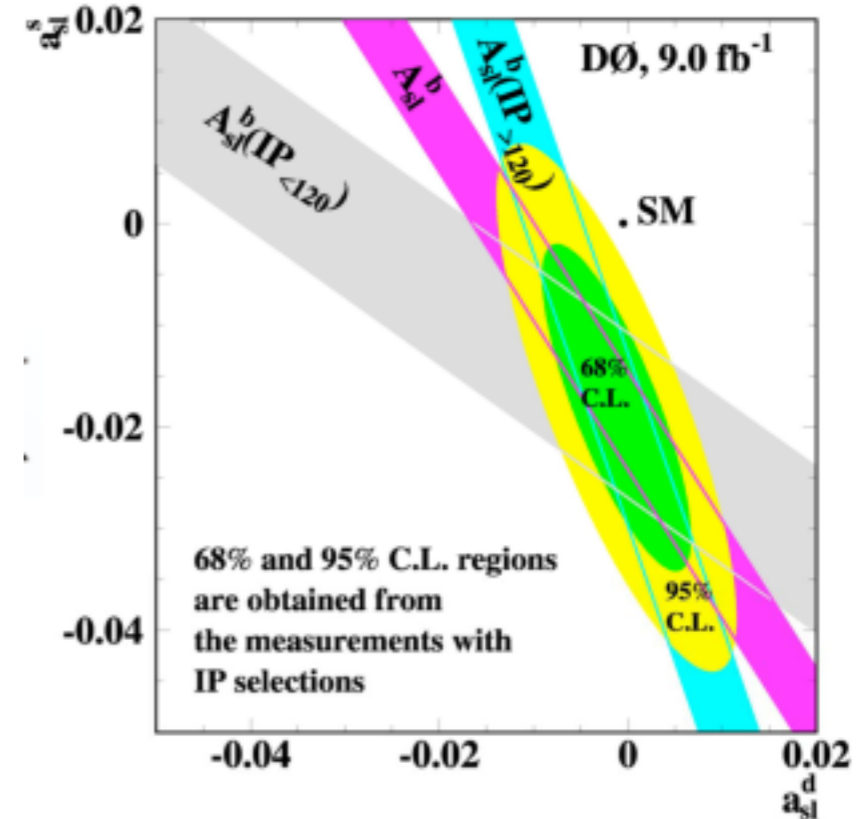
Cosmology : $m_\nu = m_1 + m_2 + m_3 < 0.5 - 1 \text{ eV (95 \% CL)}$

Tevatron New Physics Searches

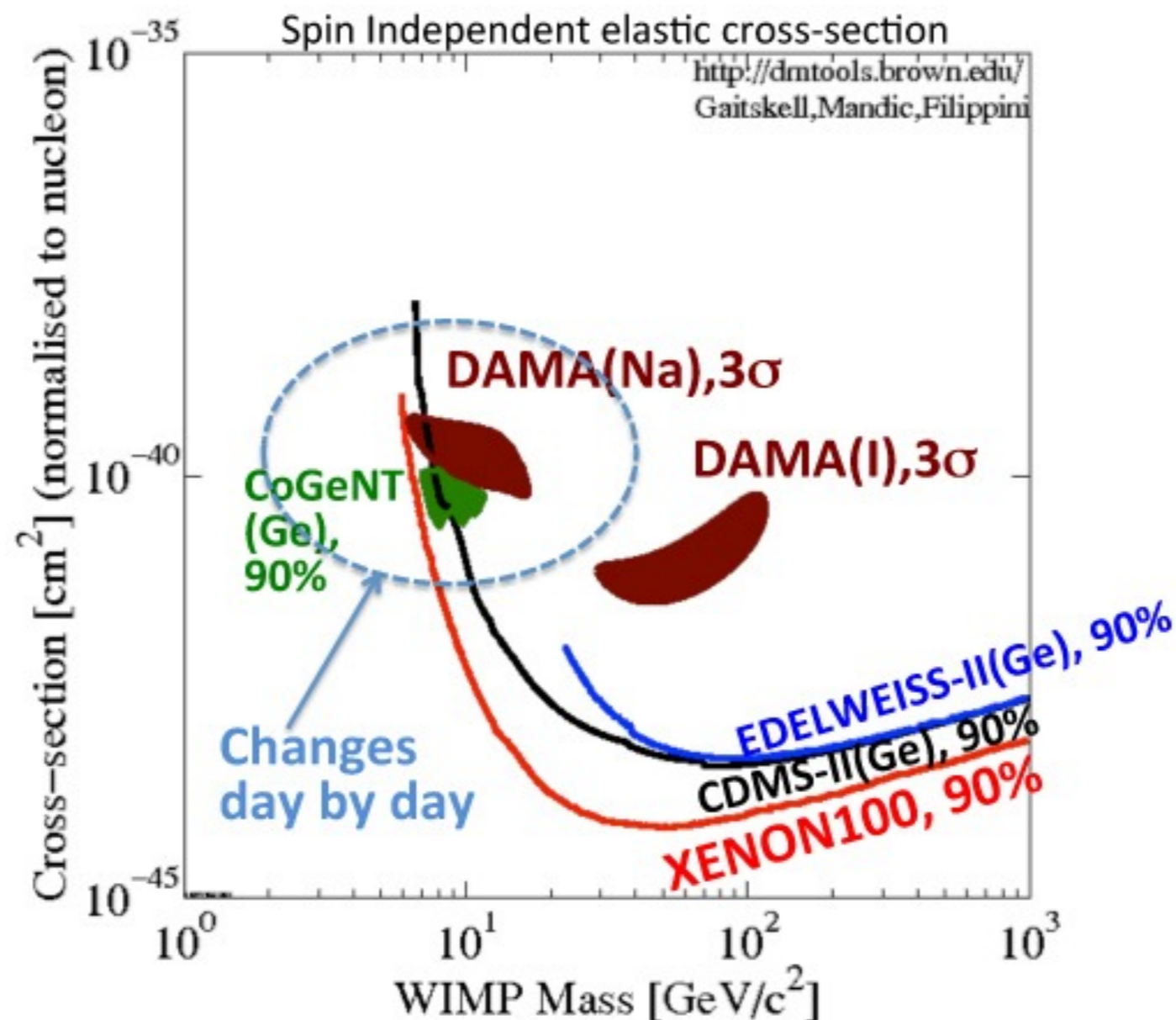
- ▶ No evidence for BSM physics in:
 - ▶ Structure / deviations in jet spectrum
 - ▶ Structure / deviations in dilepton / diphoton spectrum
 - ▶ tt resonances / anomalous top σ
 - ▶ 4th generation searches
 - ▶ Leptoquarks etc
 - ▶ Hidden valley and novel signatures
 - ▶ Many other studies
- ▶ Hints?
 - ▶ W+jets spectrum?
 - ▶ Looks to have gone away in cross-checks
 - ▶ Dimuon asymmetry (3.9σ from SM)?
 - ▶ t tbar FB asymmetry?
 - ▶ Bs \rightarrow mu mu?



arXiv: 1106.6308

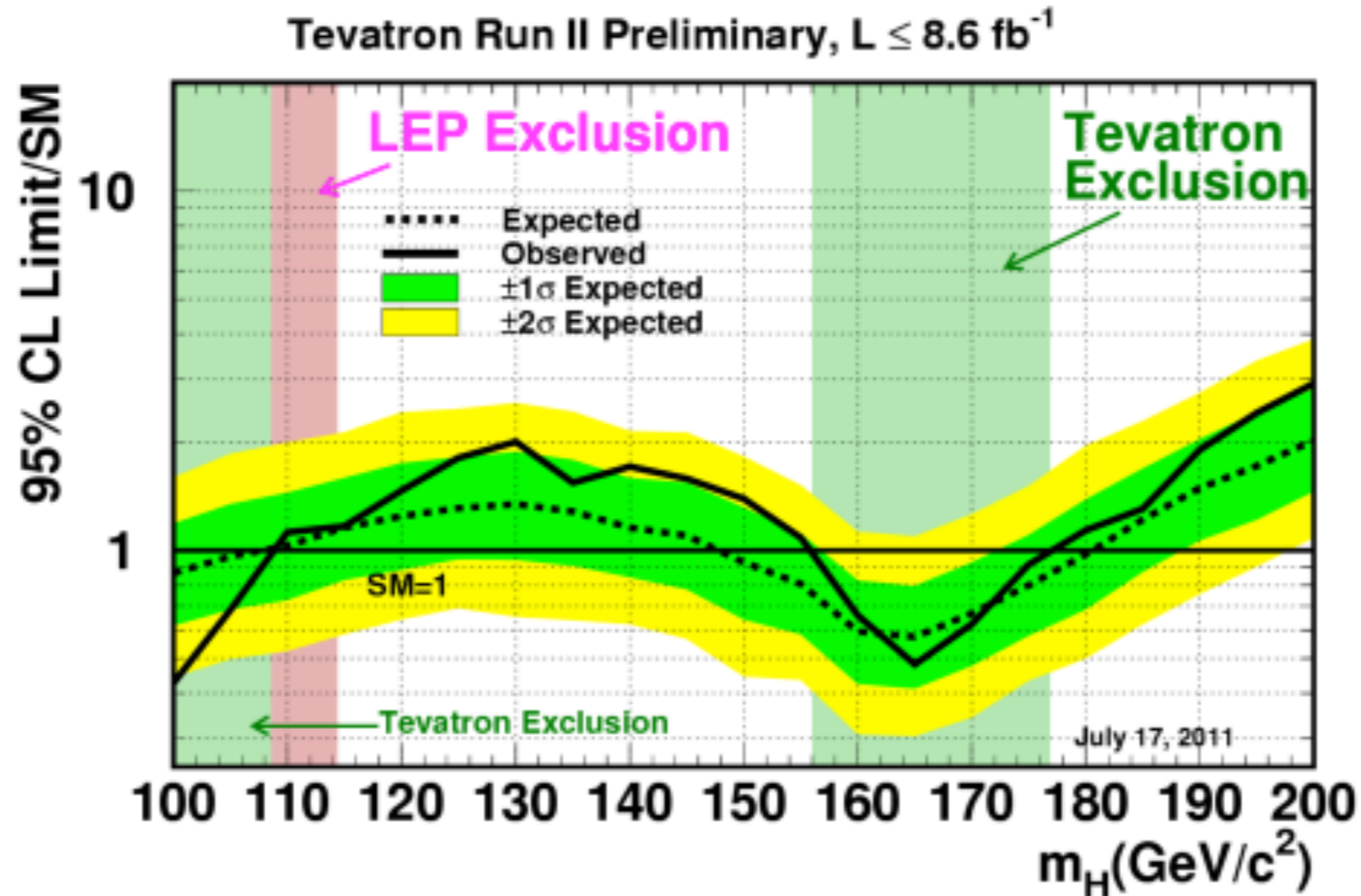


Direct DM Searches



- ▶ Substantial confusion & controversy over DM results
 - ▶ *Two* experiments observe annual modulation – interpretation challenging
 - ▶ More results and cross-checks urgently needed

Tevatron / LEP Higgs Limits



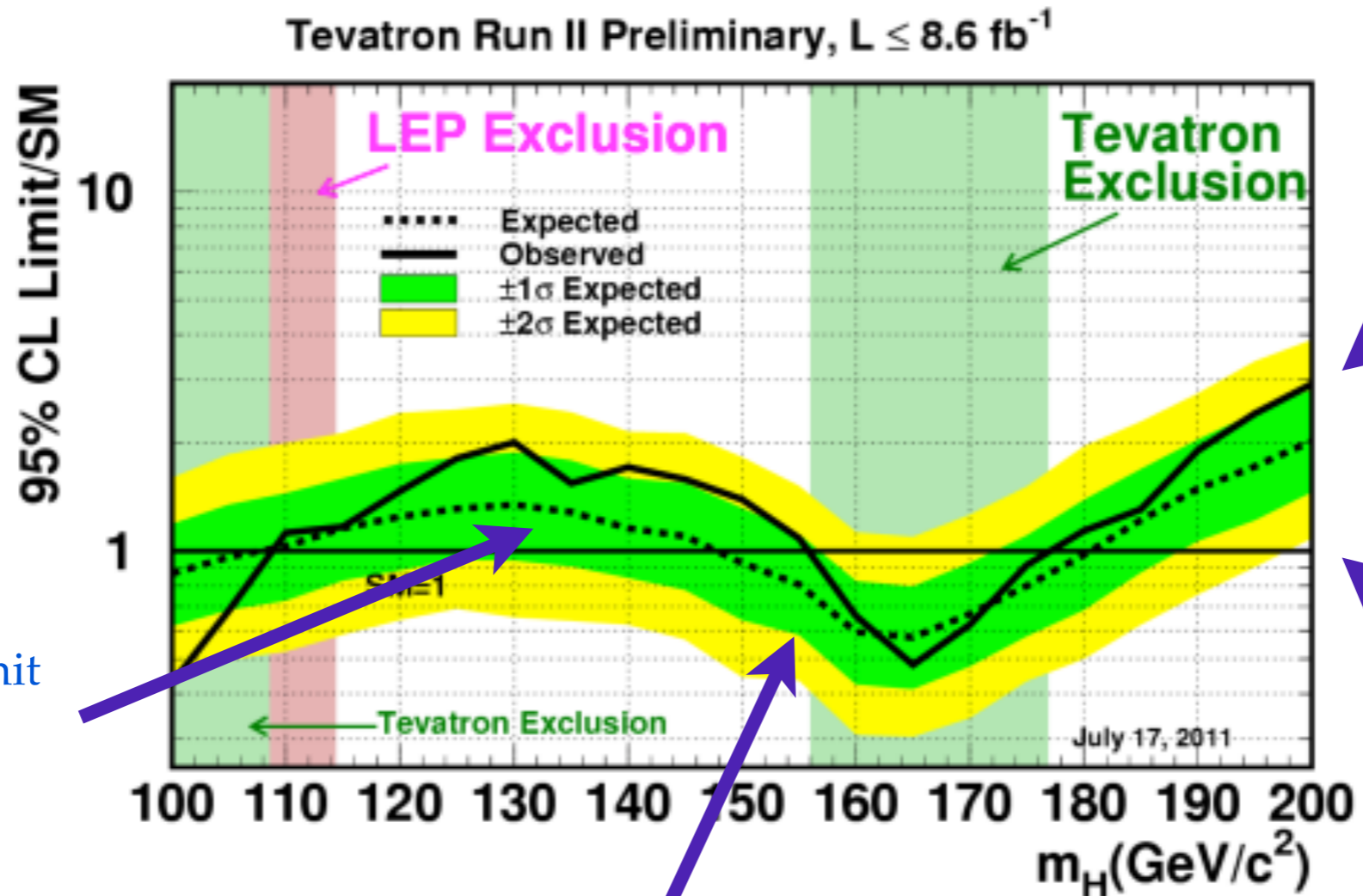
Observed Exclusion : 100-109 and 156-177 GeV/c^2

Expected Exclusion : 100-108 and 148-181 GeV/c^2

- ▶ Not taking into account various 'ways out'
 - ▶ *Higgsless* models; *invisible* higgs; *buried* higgs; Higgs with *phobias*, etc etc

Limit Plots: Spotter's Guide

Upper x-section
limit relative to SM



Observed limit
from data

Median expected limit
based on many
background-only
pseudoexperiments

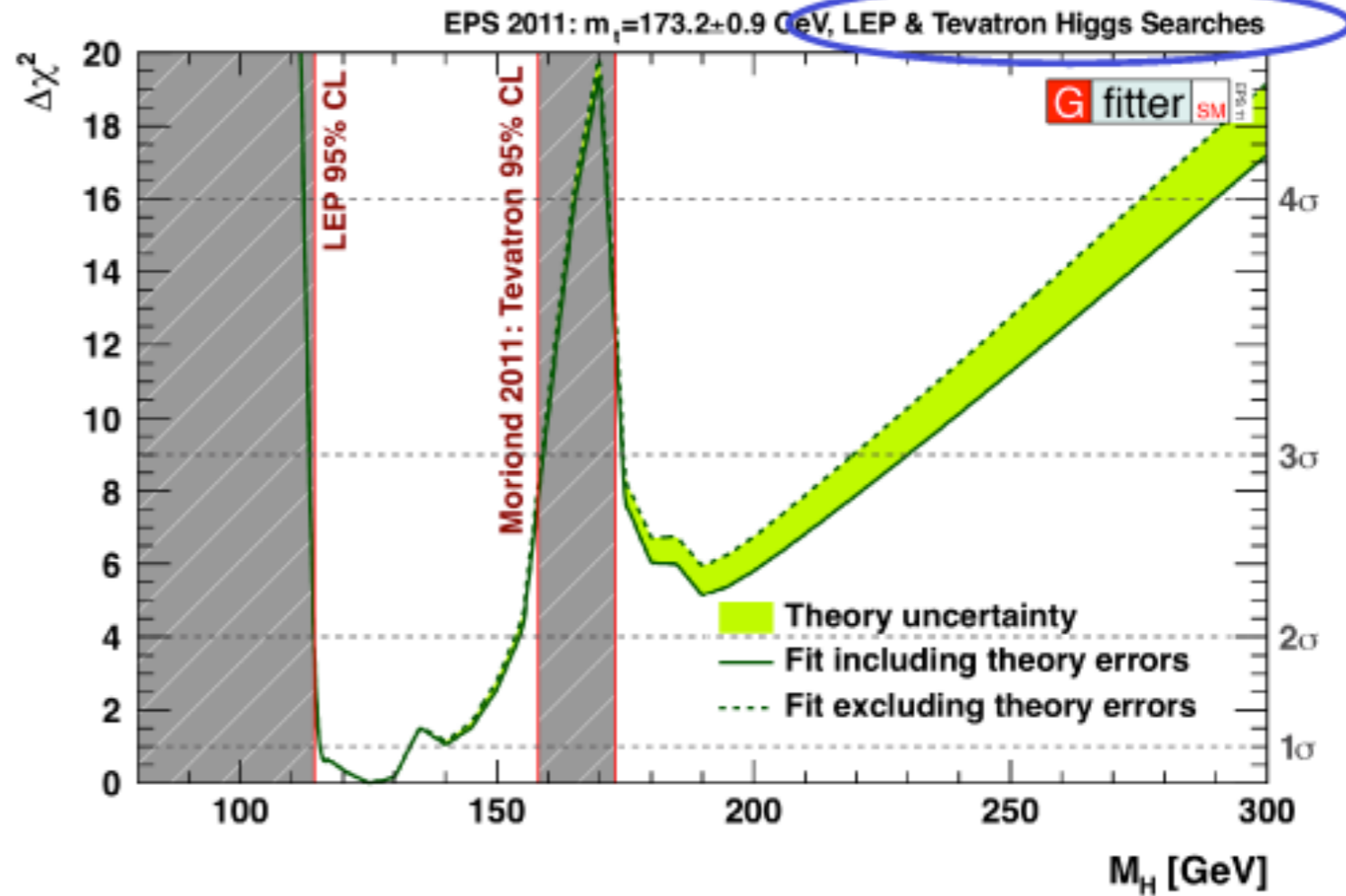
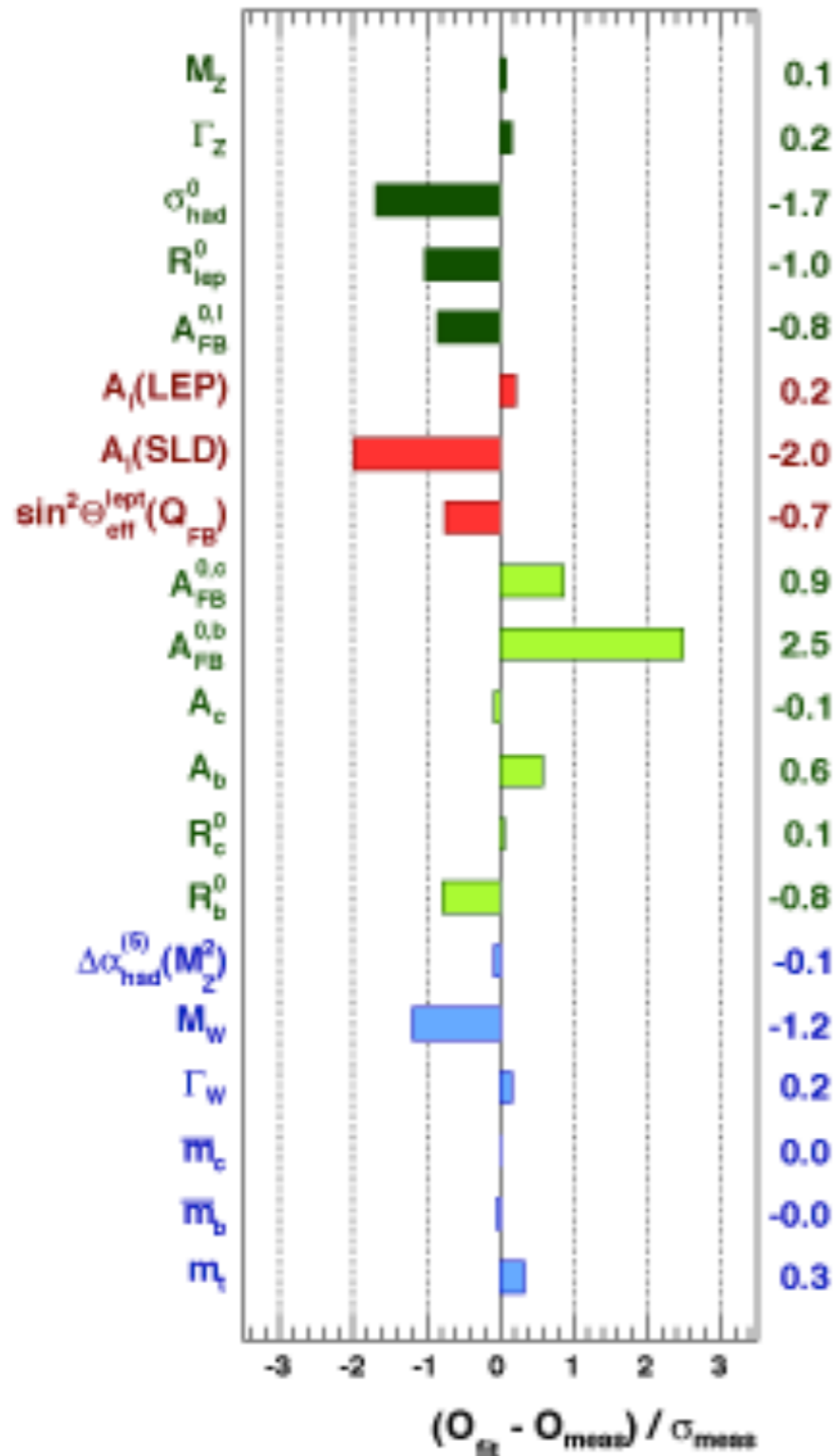
Standard Model
prediction

1σ (67%) and 2σ (95%)
bands on expected limit

Limit calculated at
many different mass
hypotheses
- not independent!

- ▶ Statistics is important: remember, only $O(10)$ events expected
 - ▶ Beware: expected limit band refers to *statistical* fluctuations only
 - ▶ These plots widely misinterpreted, esp. for searches (look elsewhere effect)

Electroweak Global Fits



Whither the Standard Model?

Possibly more than a flesh wound



But Not Dead Yet

What Next?

What Next?



What Next?



“We’re going to need a bigger boat”

Welcome to the Energy Frontier

▶ Something interesting (probably) happens at $\sim 1\text{TeV}$

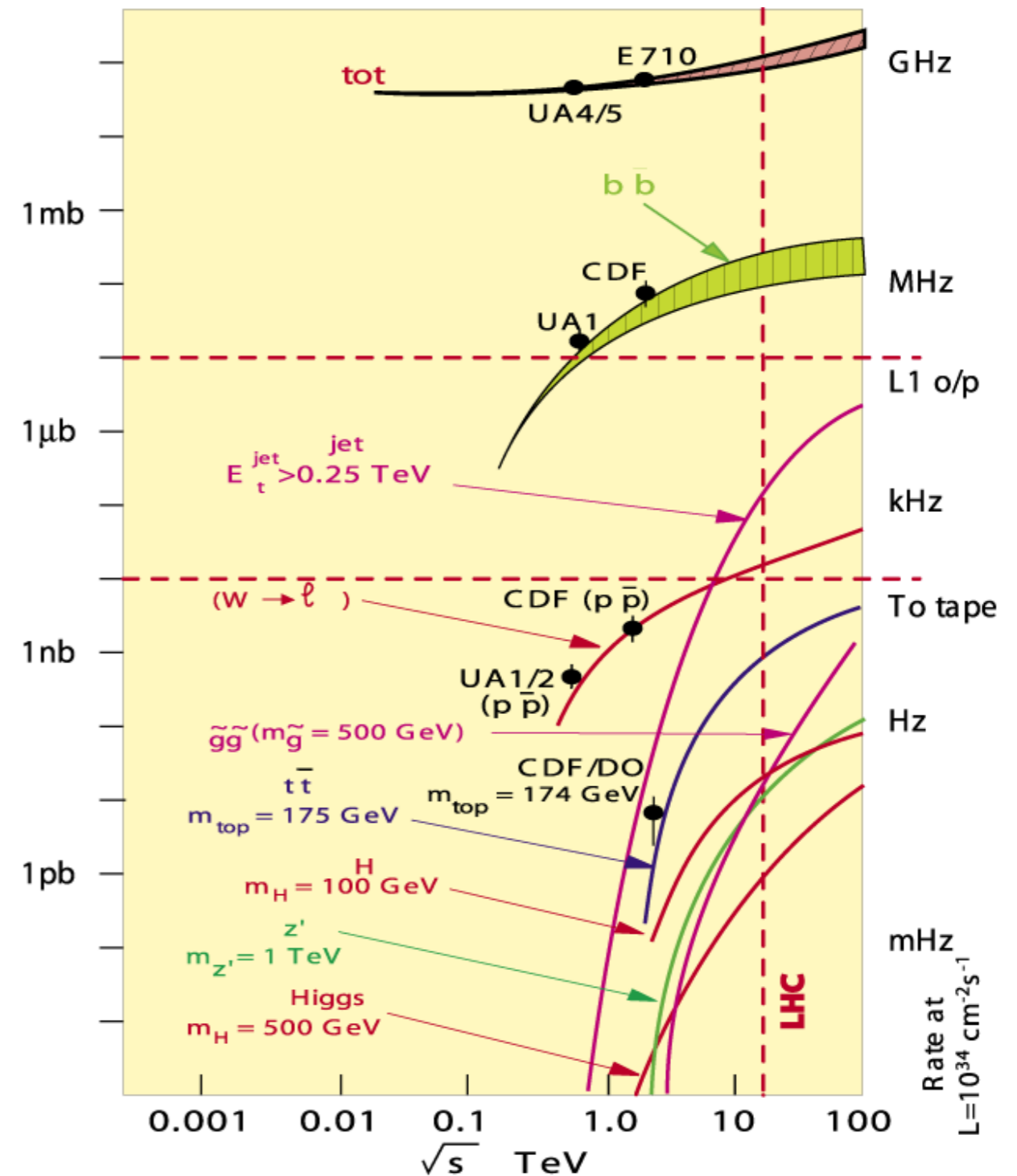
- ▶ Build a collider with significant luminosity above $\sqrt{s}=1\text{TeV}$
- ▶ See what happens ...
- ▶ ... be prepared for almost anything

▶ Looking for rare processes

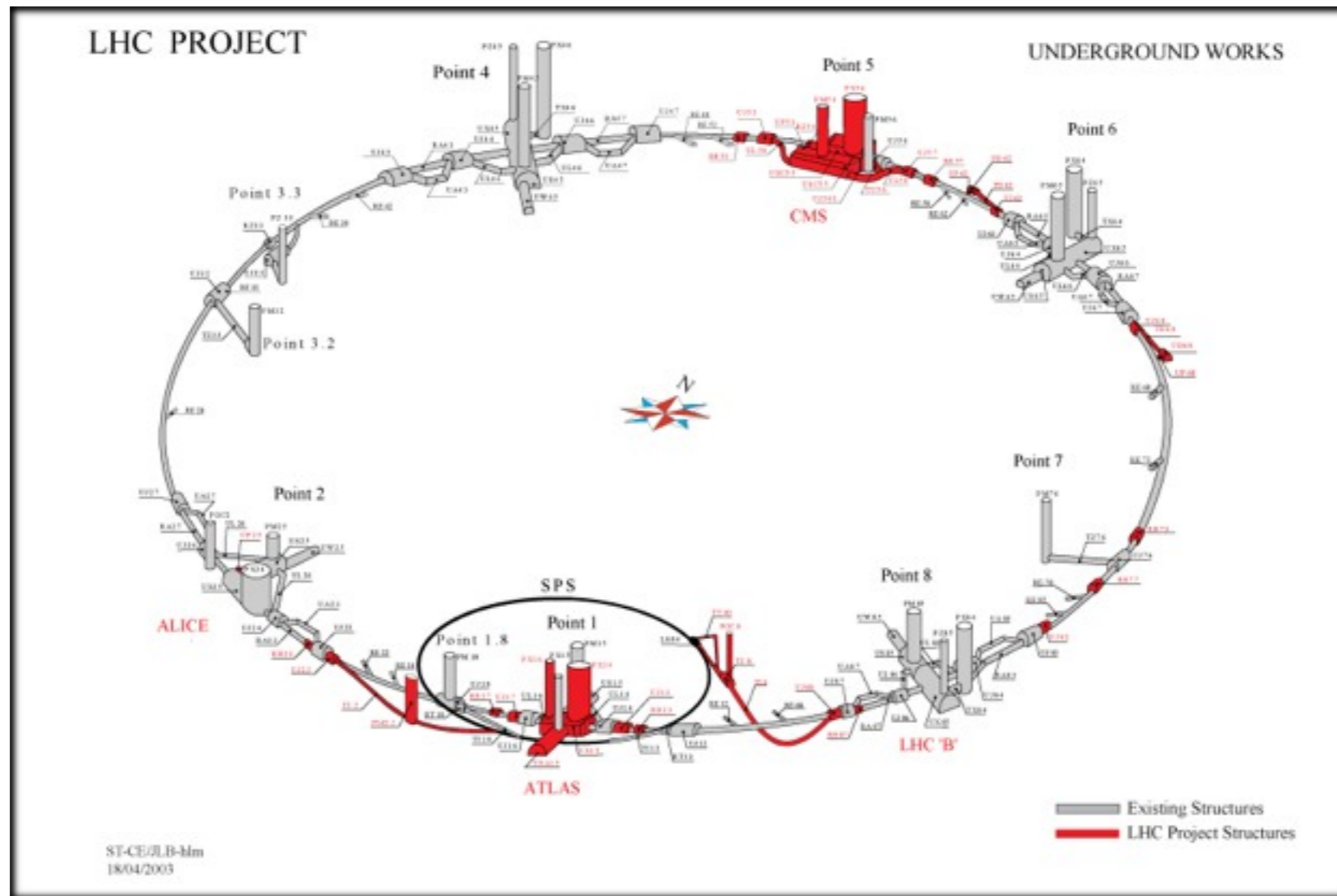
- ▶ Need very high luminosity
- ▶ pp collider is the only choice
 - ▶ Better described as a g-g collider

▶ Experimental challenges

- ▶ QCD – there's a lot of it about
- ▶ Triggering and data handling
- ▶ High energy final states
 - ▶ Strong B field, big detectors
- ▶ Radiation dose & longevity



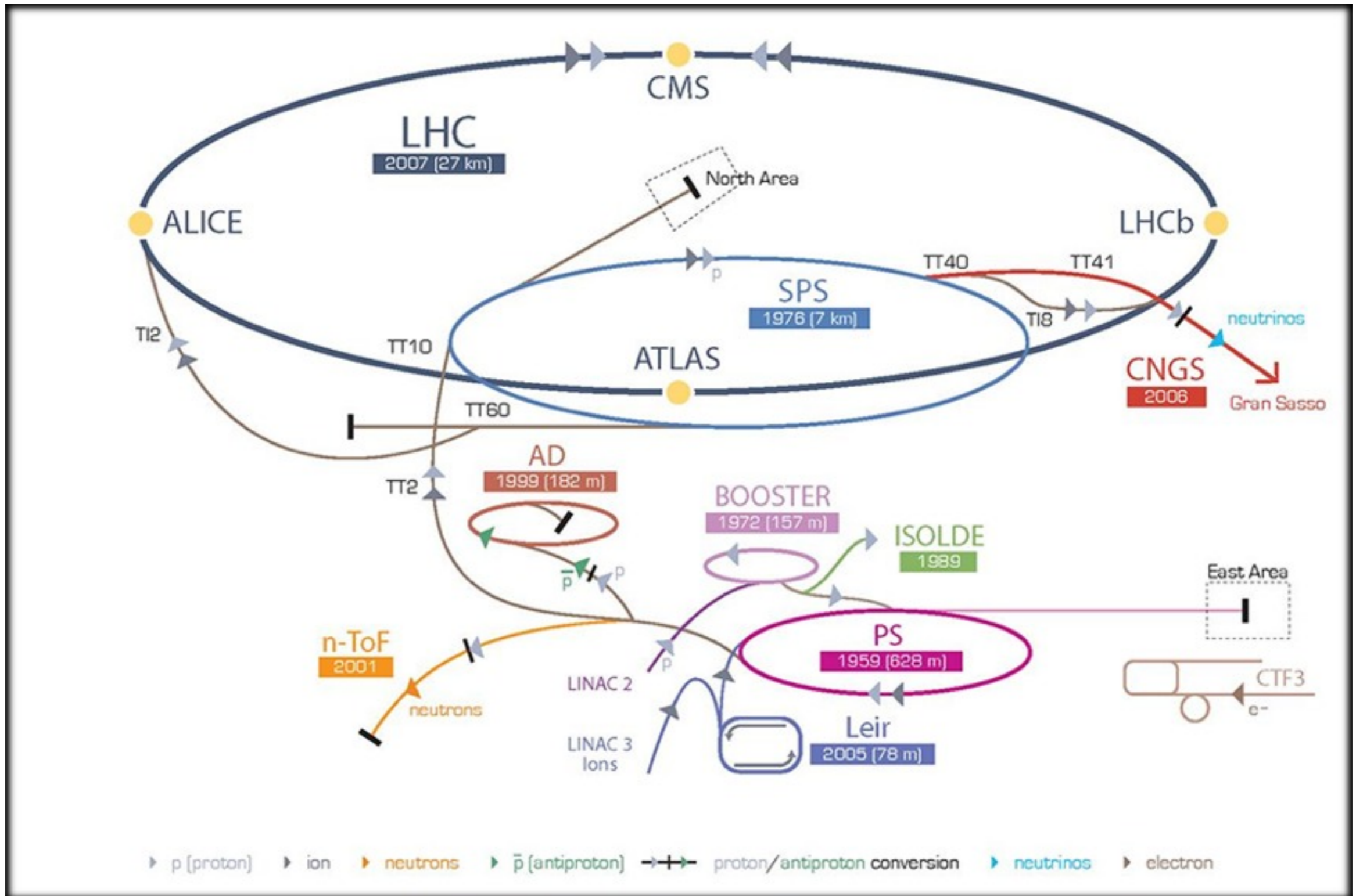
The Machine



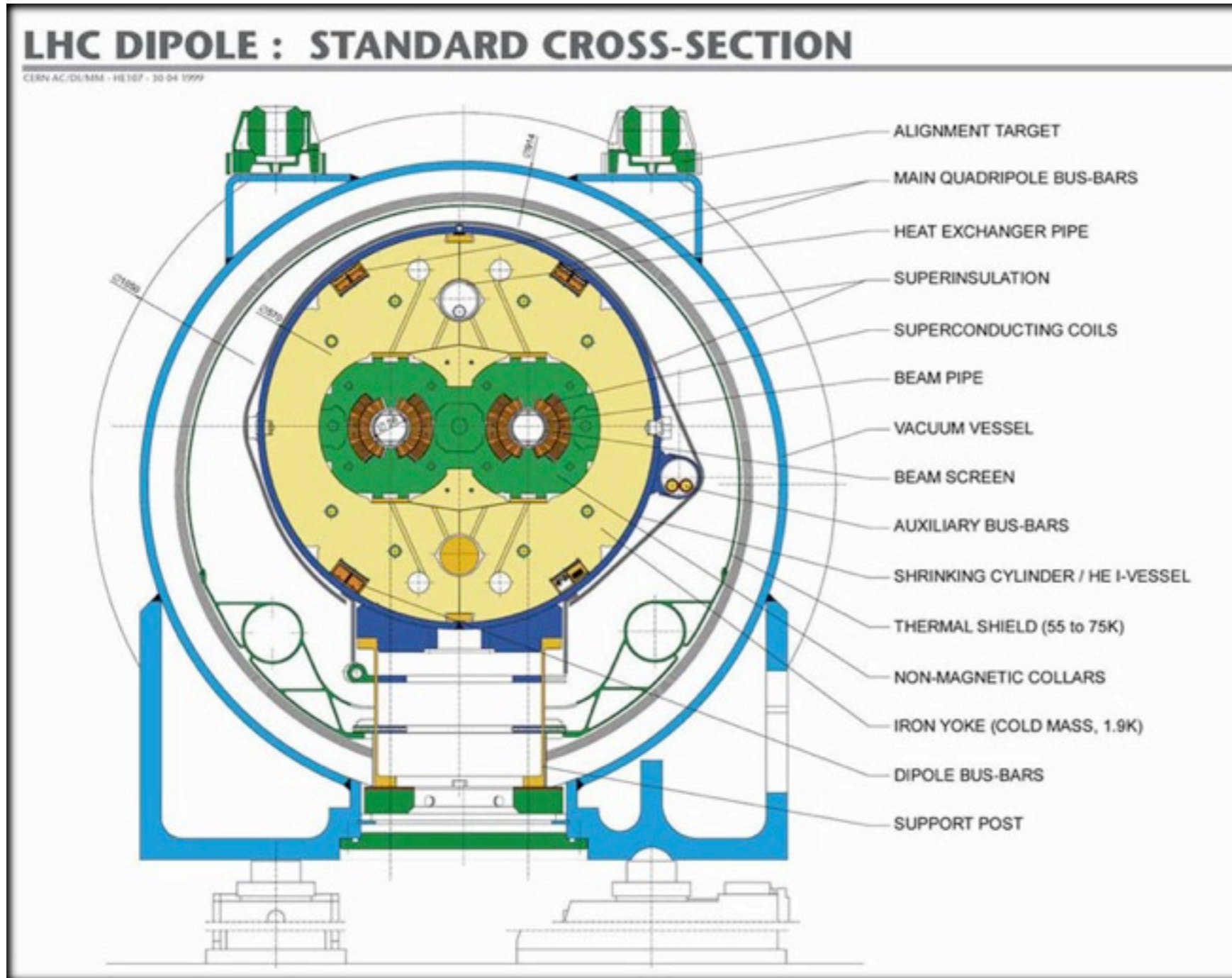
▶ Design parameters

- ▶ Maximum possible energy in already-constructed LEP ring ($7 + 7$ TeV)
- ▶ Highest possible luminosity within cost ($10^{34} \text{cm}^{-2}\text{s}^{-1}$, $100 / \text{fb} / \text{yr}$)
- ▶ Cost: $\sim 2 \text{GCHF}$ (same as LEP); ~ 12 years construction

CERN Accelerator Complex



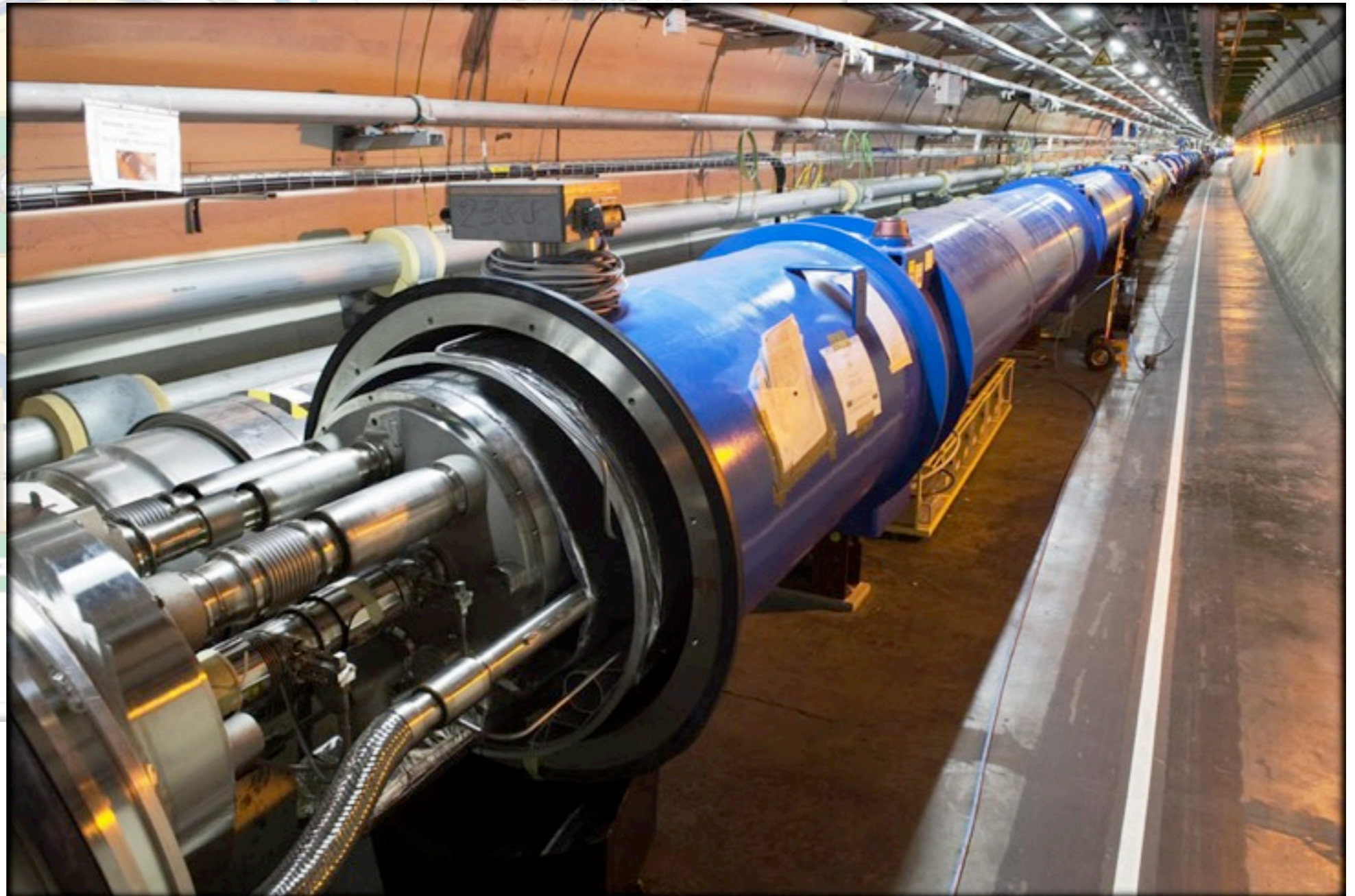
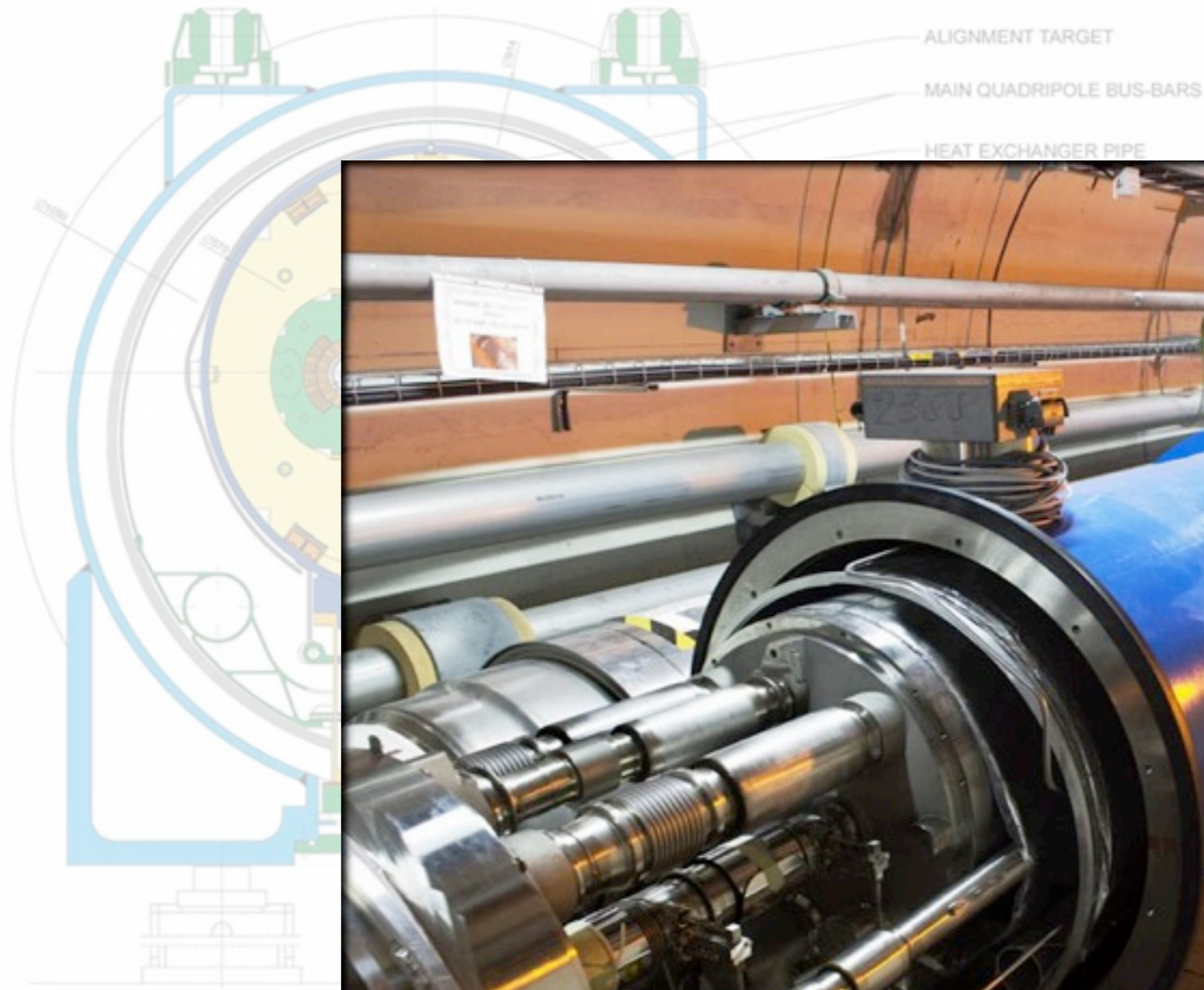
Superconducting Magnets



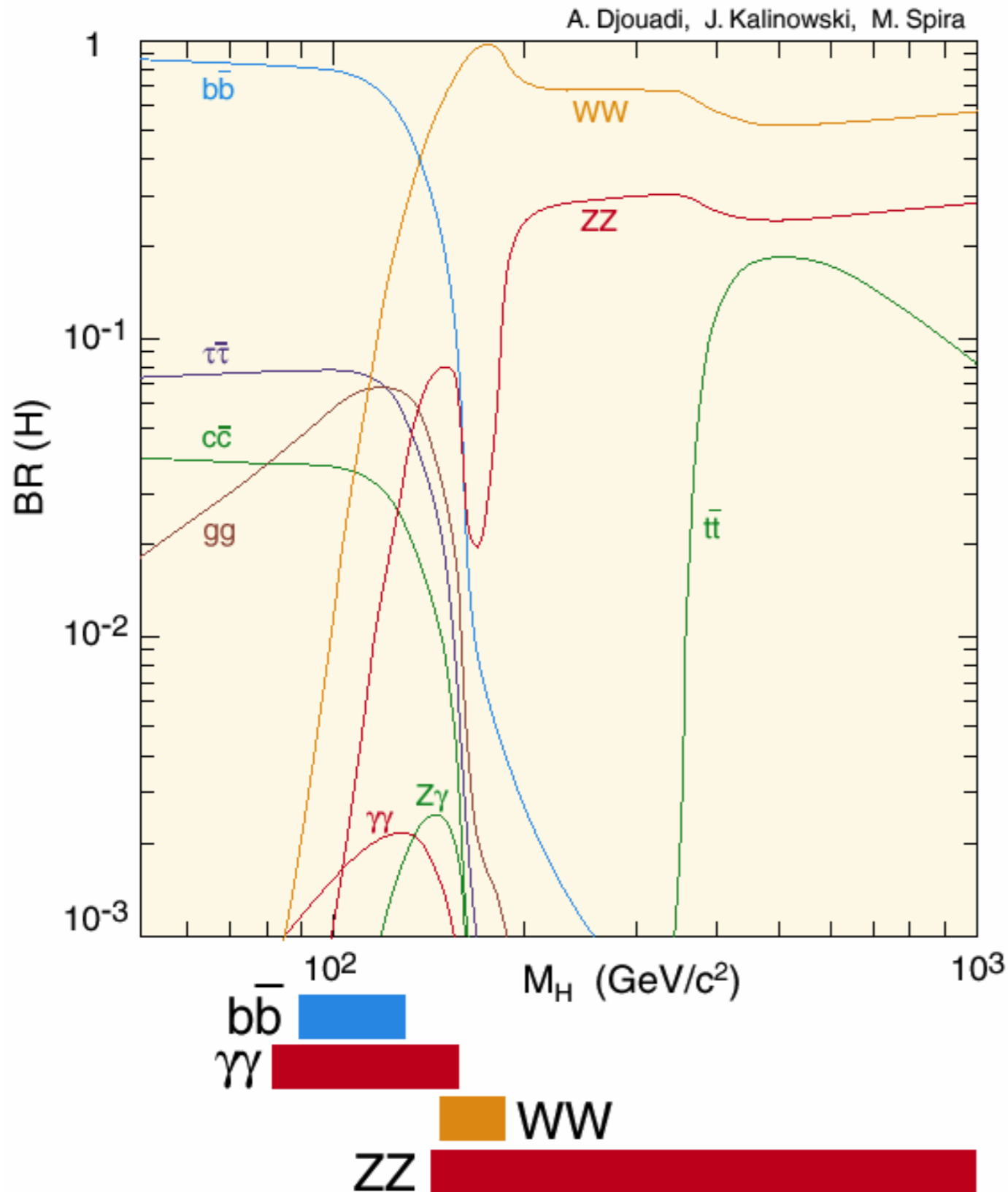
Superconducting Magnets

LHC DIPOLE : STANDARD CROSS-SECTION

CERN AC.DU.MM - H1187 - 30.04.1999



Designing the GPDs



- ▶ Design for most challenging case

- ▶ e.g. light SM H $\rightarrow \gamma\gamma$
- ▶ Fortunate decision, it seems

- ▶ Points to note:

- ▶ H $\rightarrow b\bar{b}$ is experimentally inaccessible at LHC
 - ▶ Modulo tricks with exclusive production?
 - ▶ Or is it?
- ▶ H $\rightarrow \gamma\gamma$ has very low BR
 - ▶ And there are plenty of high p_t photons from other sources
- ▶ SM Higgs width at ~ 100 GeV is narrow
 - ▶ This helps considerably

GPDs: Main Features

- ▶ Robust muon systems
 - ▶ Background estimates had large uncertainties in 1990s
 - ▶ If all else fails... nobody ever got fired for finding muons
 - ▶ Heavy shielding can cut down most sources of background
- ▶ Exceptional electromagnetic calorimetry (e, γ)
 - ▶ Motivated by narrowness of H \rightarrow $\gamma\gamma$ peak
 - ▶ Also helps with leptonic decays of heavy states
- ▶ High performance, highly redundant, tracking
 - ▶ Requires extremely strong magnetic field (4T / 2T for CMS / ATLAS)
 - ▶ Many layers of radiation-resistant silicon
 - ▶ Multi-layer pixel detectors for track finding with large occupancy
- ▶ Two GPDs at the LHC
 - ▶ Identical physics goals, different experimental strategies
 - ▶ So far, both appear to have comparable performance

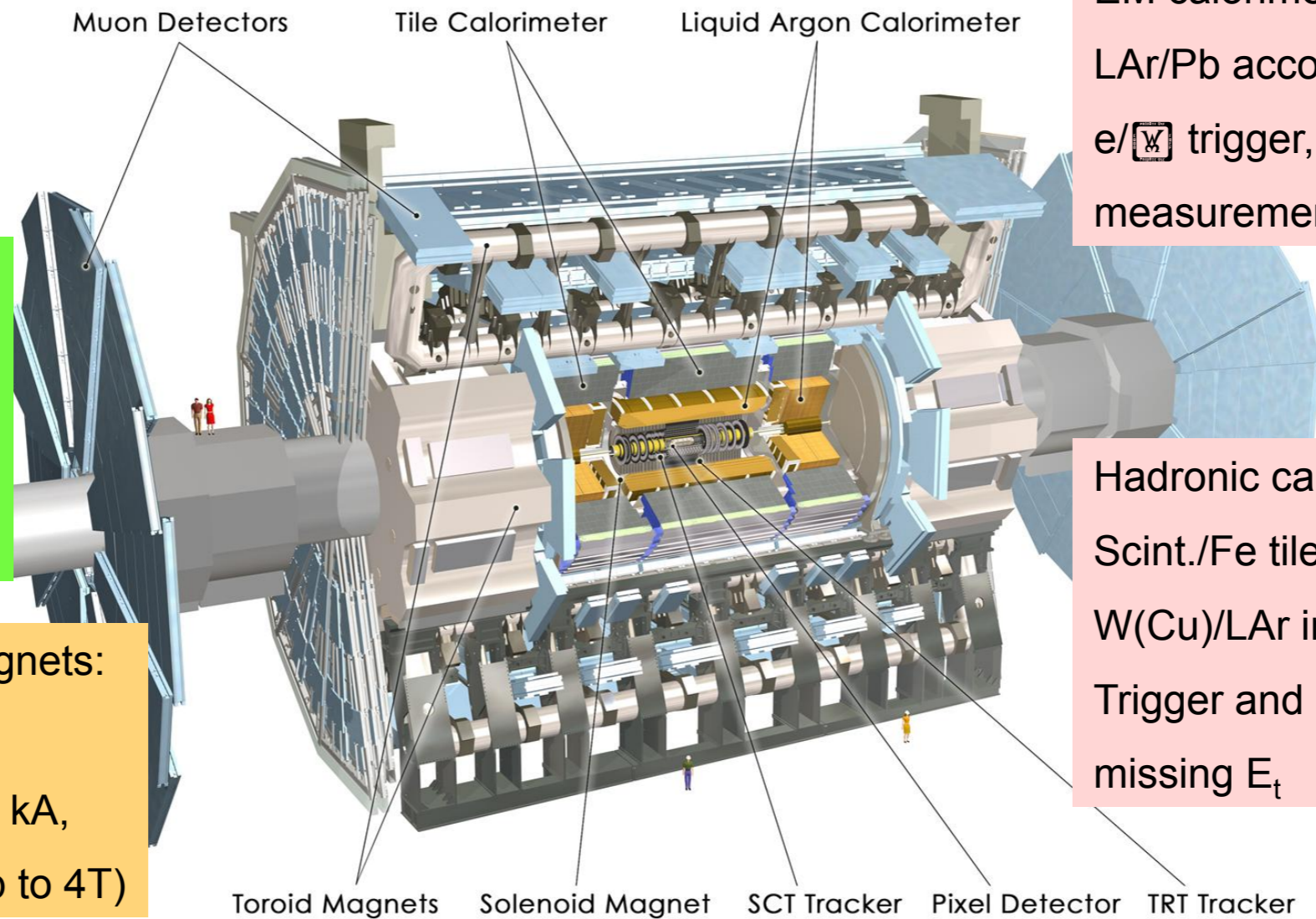
ATLAS

Standalone muon spectrometer ($\eta < 2.7$), 3 layers gas based muon chambers,
muon trigger and muon momentum determination

- 46m length,
- 25m diameter,
- 7000t weight

- 3 level trigger,
- Collision rate 40MHz,
- LV1 accepts up to 75kHz,
- recorded ~300 Hz

Super-conducting magnets:
2T Solenoid (7.6kA) ,
3 Air Core Toroids (22 kA,
peak field strength up to 4T)



EM calorimeter:
LAr/Pb according structure
 e/γ trigger, identification +
measurement

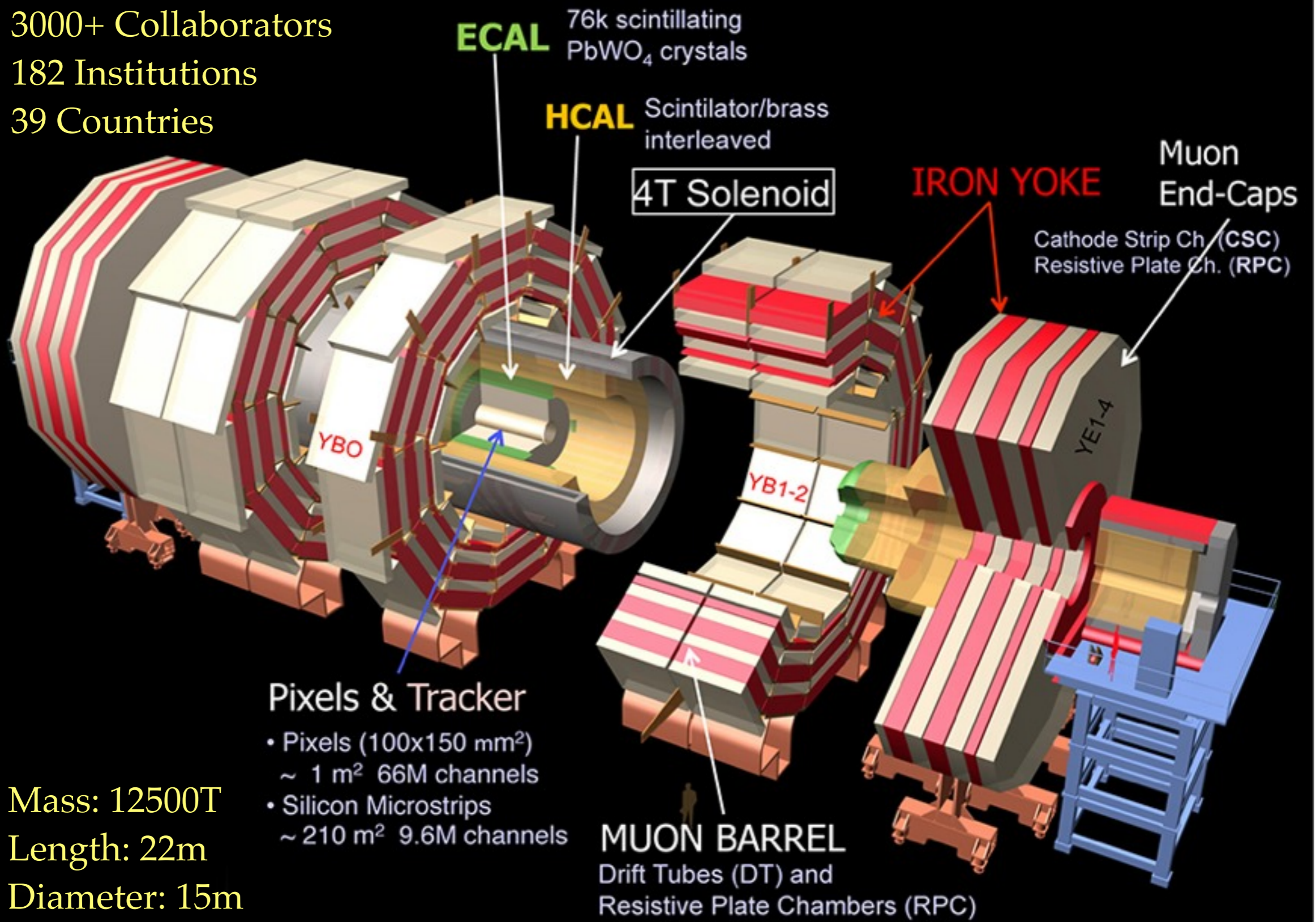
Hadronic calorimeter:
Scint./Fe tiles in the central,
W(Cu)/LAr in fwd region
Trigger and measure jets +
missing E_t

Inner Detector: $\sim 10^8$ Si Pixels, $6 \cdot 10^6$ Si Strips, Transition Radiation Tracker (TRT) – Xe-filled straw tubes interleaved with PP/PE foil for Cherenkov light: precise vertexing, tracking, e/γ separation

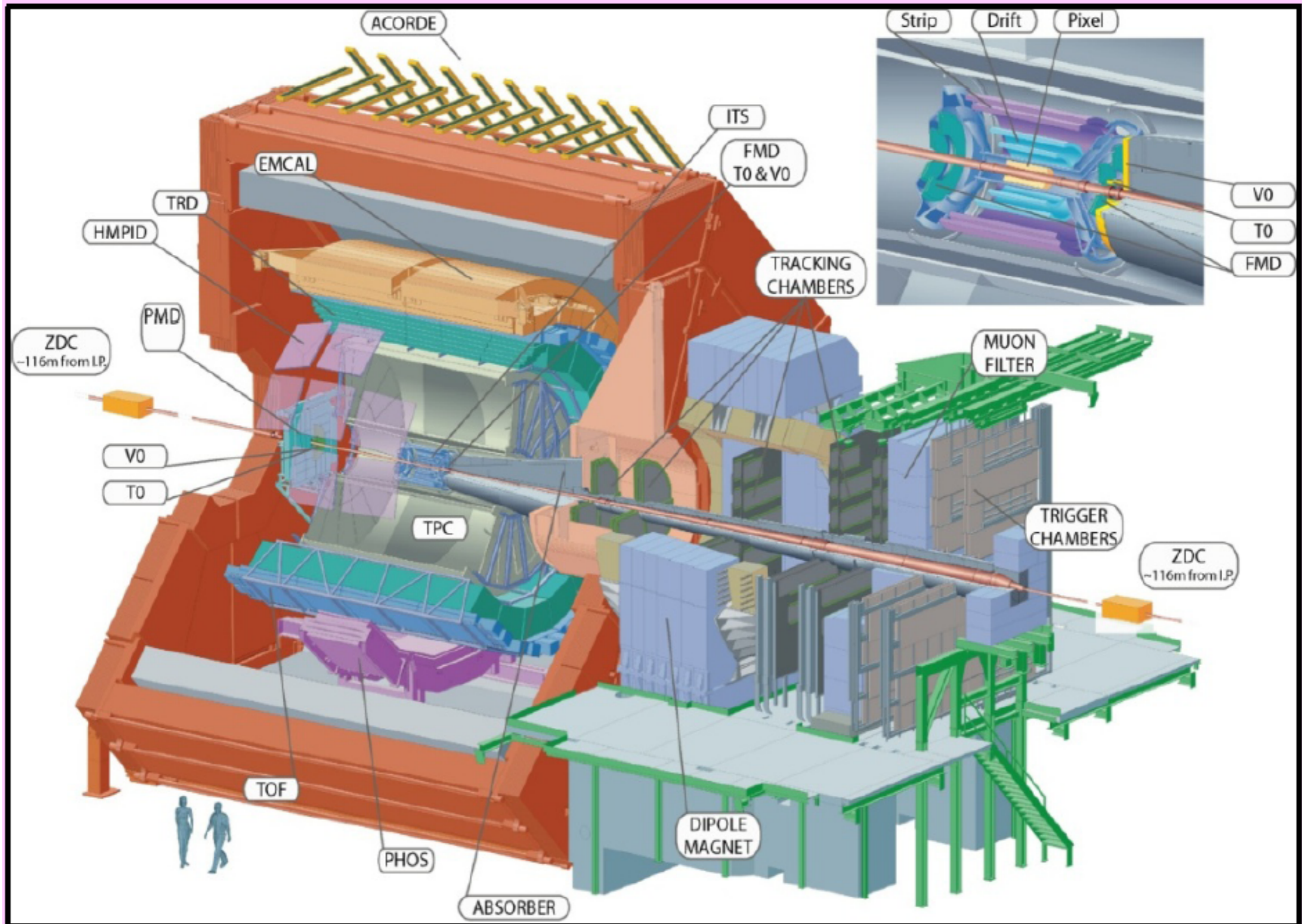
CMS

3000+ Collaborators
182 Institutions
39 Countries

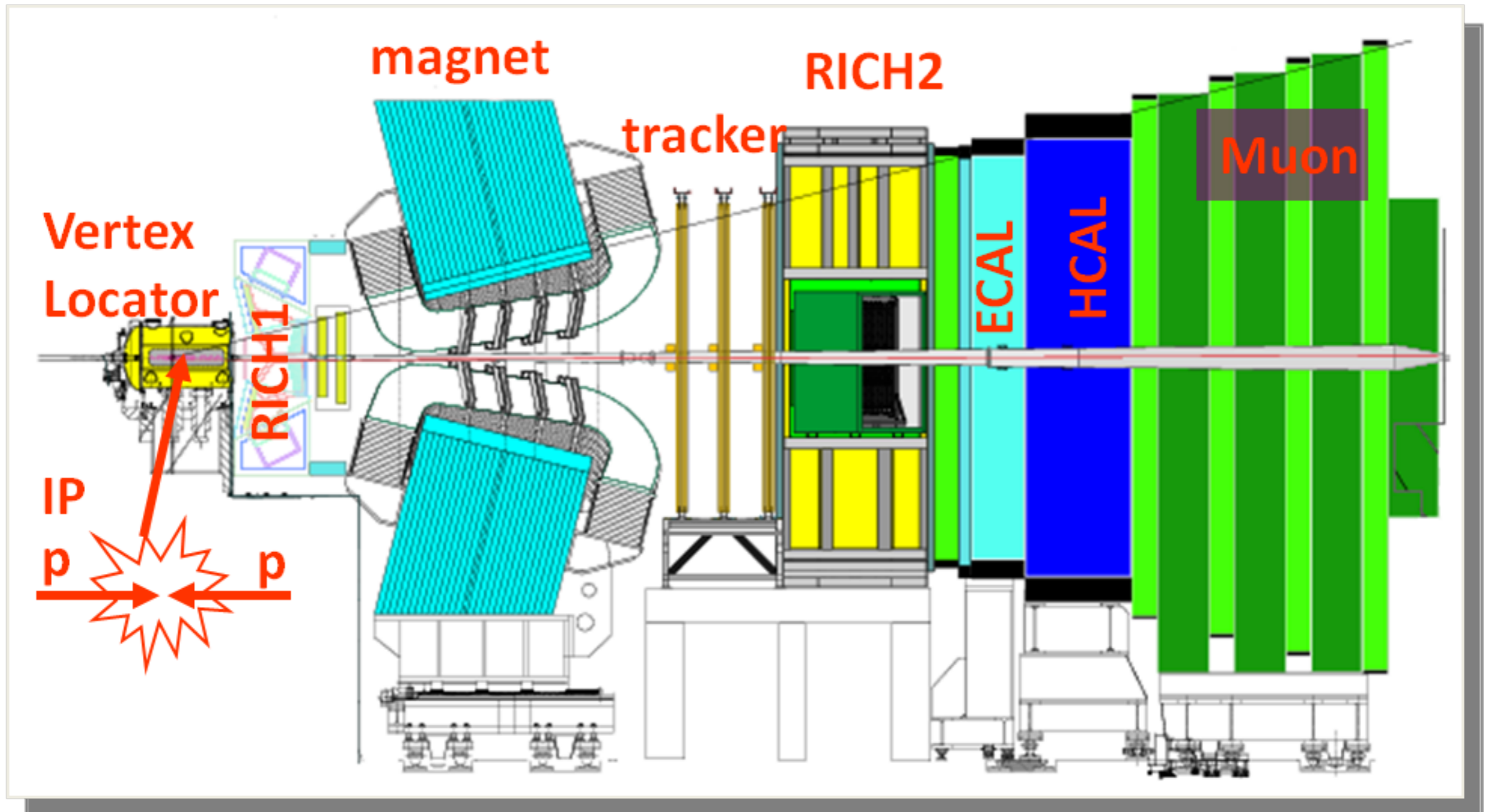
Mass: 12500T
Length: 22m
Diameter: 15m



ALICE (Heavy Ions)



LHCb (b Physics)

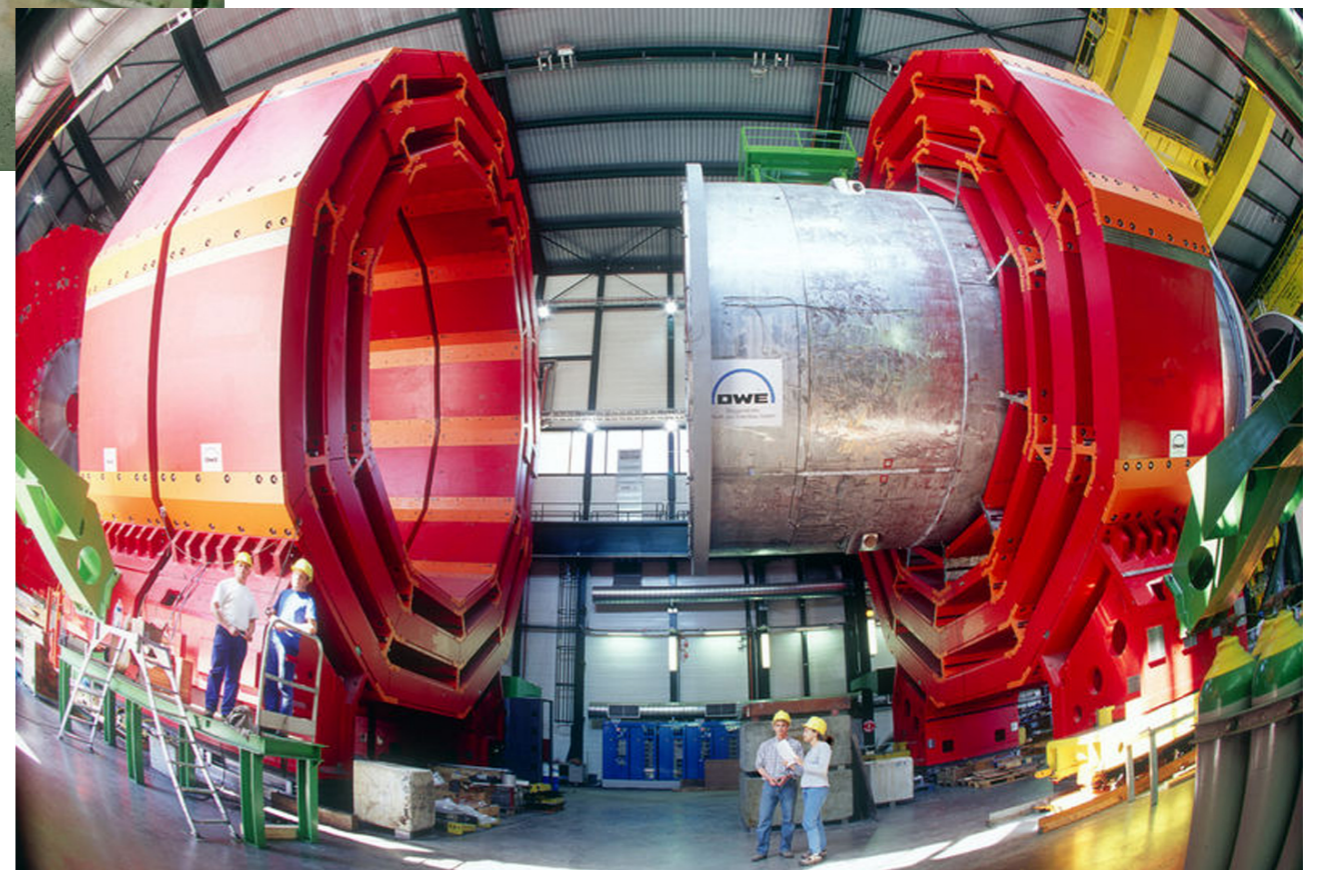


Detectors: Magnets

ATLAS toroid in transit



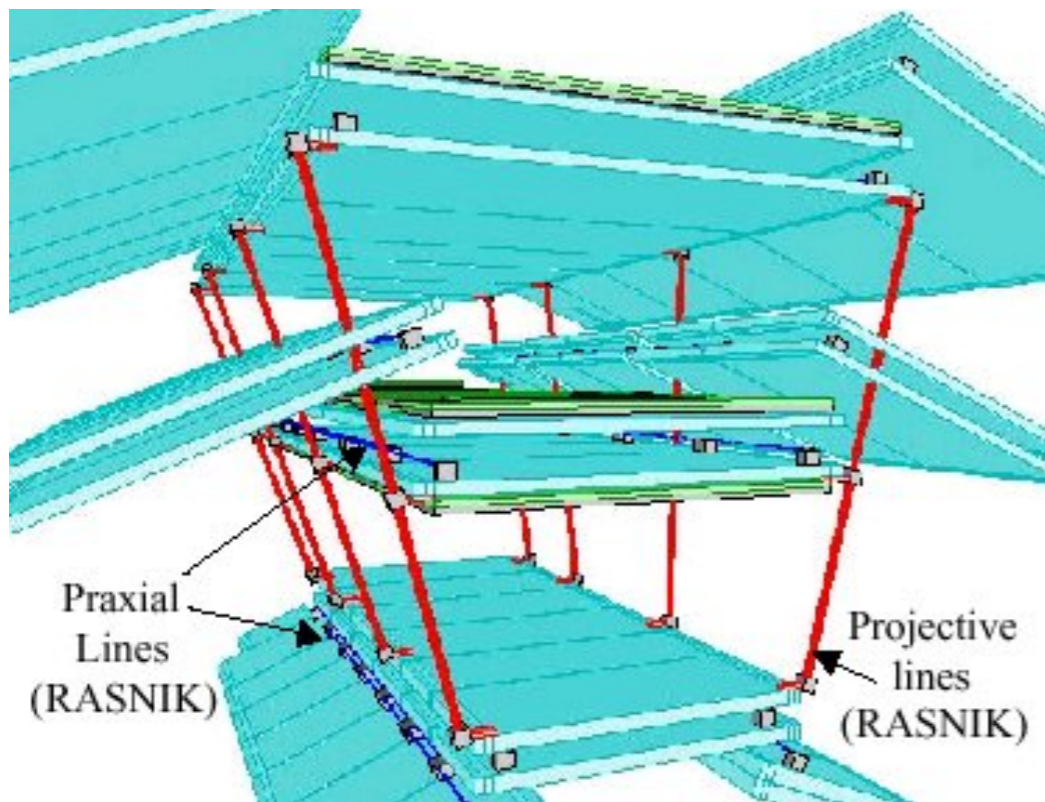
CMS yoke and solenoid vessel



▶ Key challenges

- ▶ Logistics and cost
- ▶ Safety
- ▶ Longevity

Detectors: Muons



Optical alignment of muons / tracking

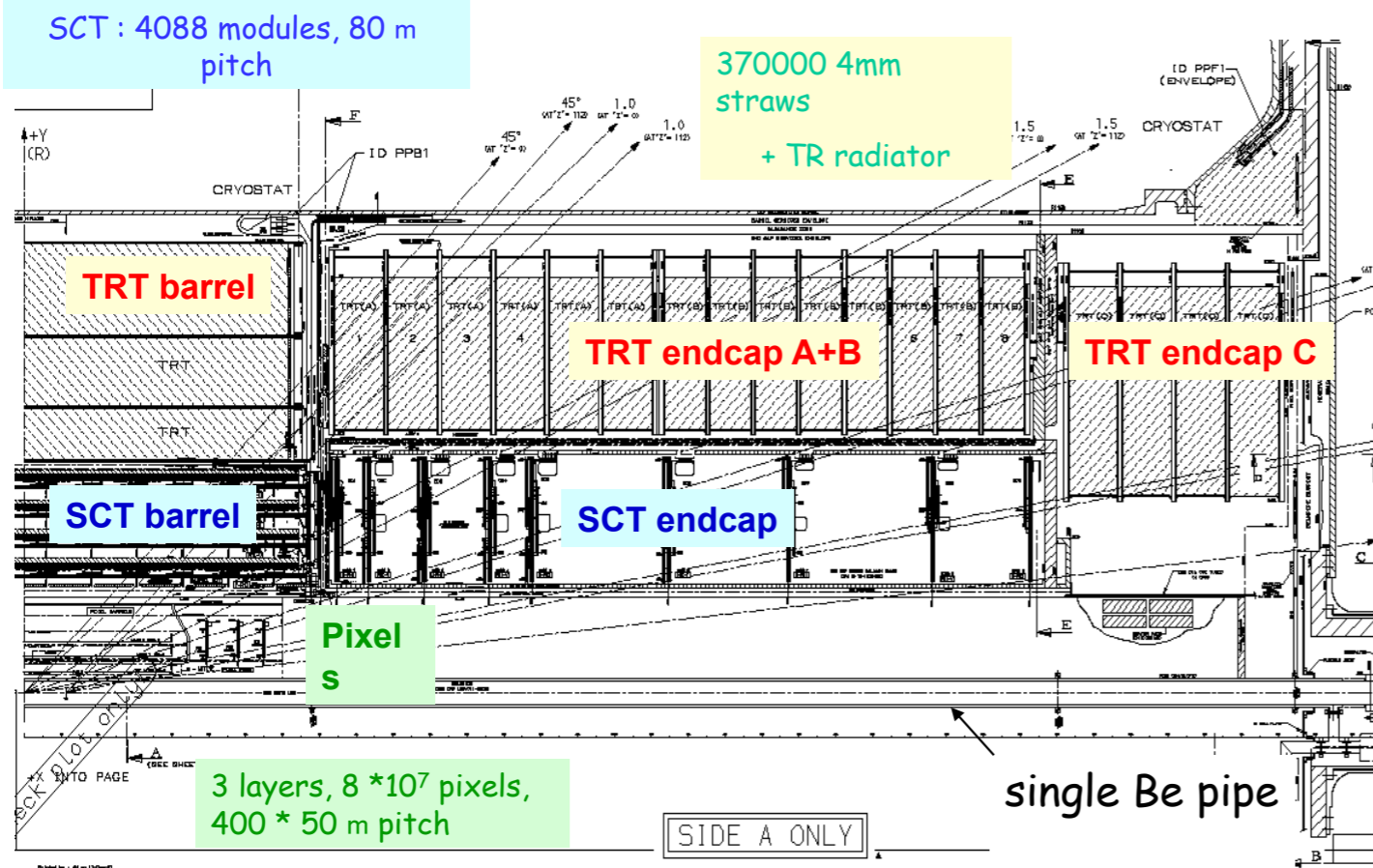
▶ Key challenges

- ▶ Construction and scale
- ▶ Alignment
- ▶ Stable operation in aging



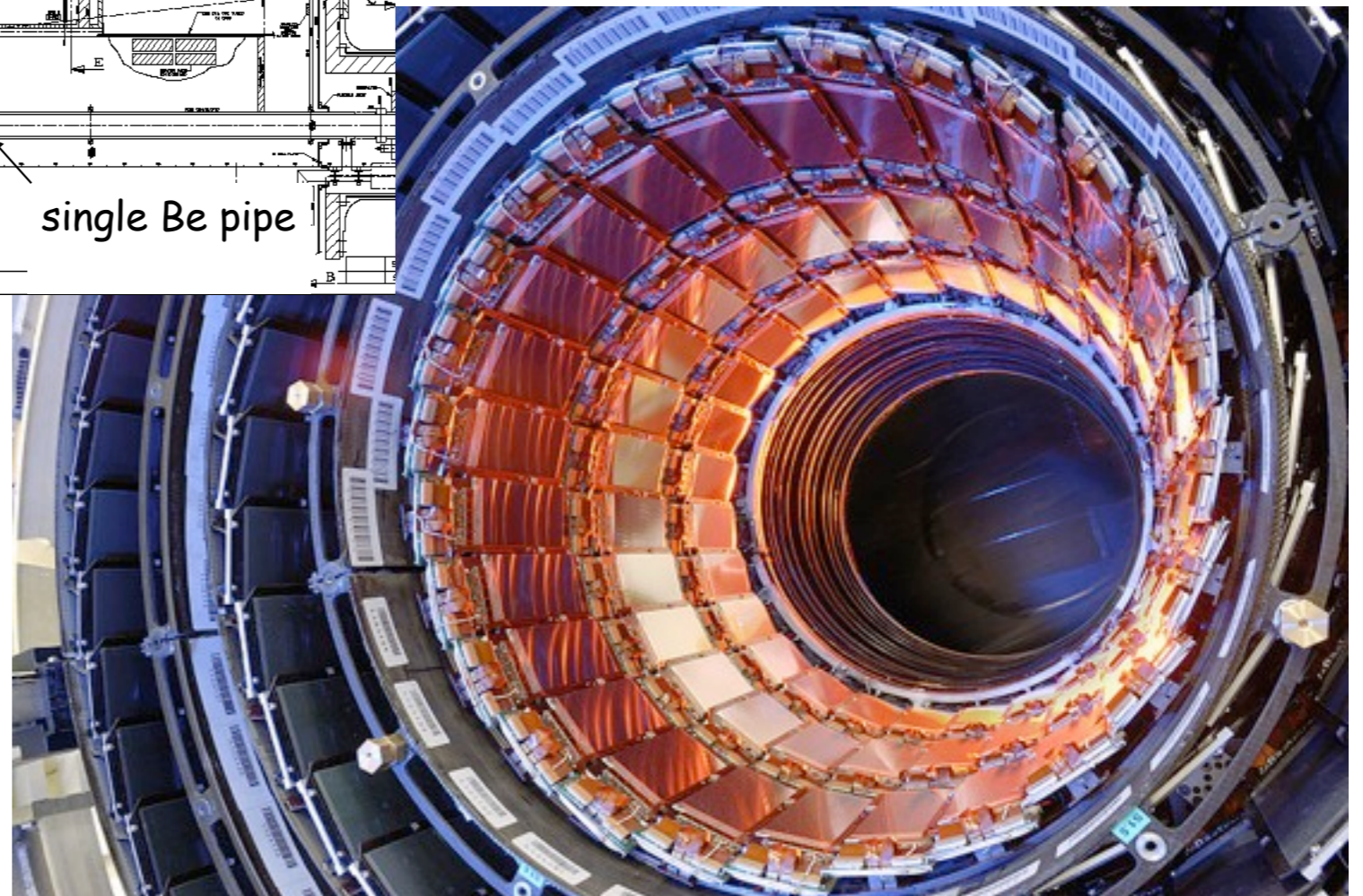
CMS endcap muon chambers

Detector Highlights: Tracking



ATLAS Inner detector layout

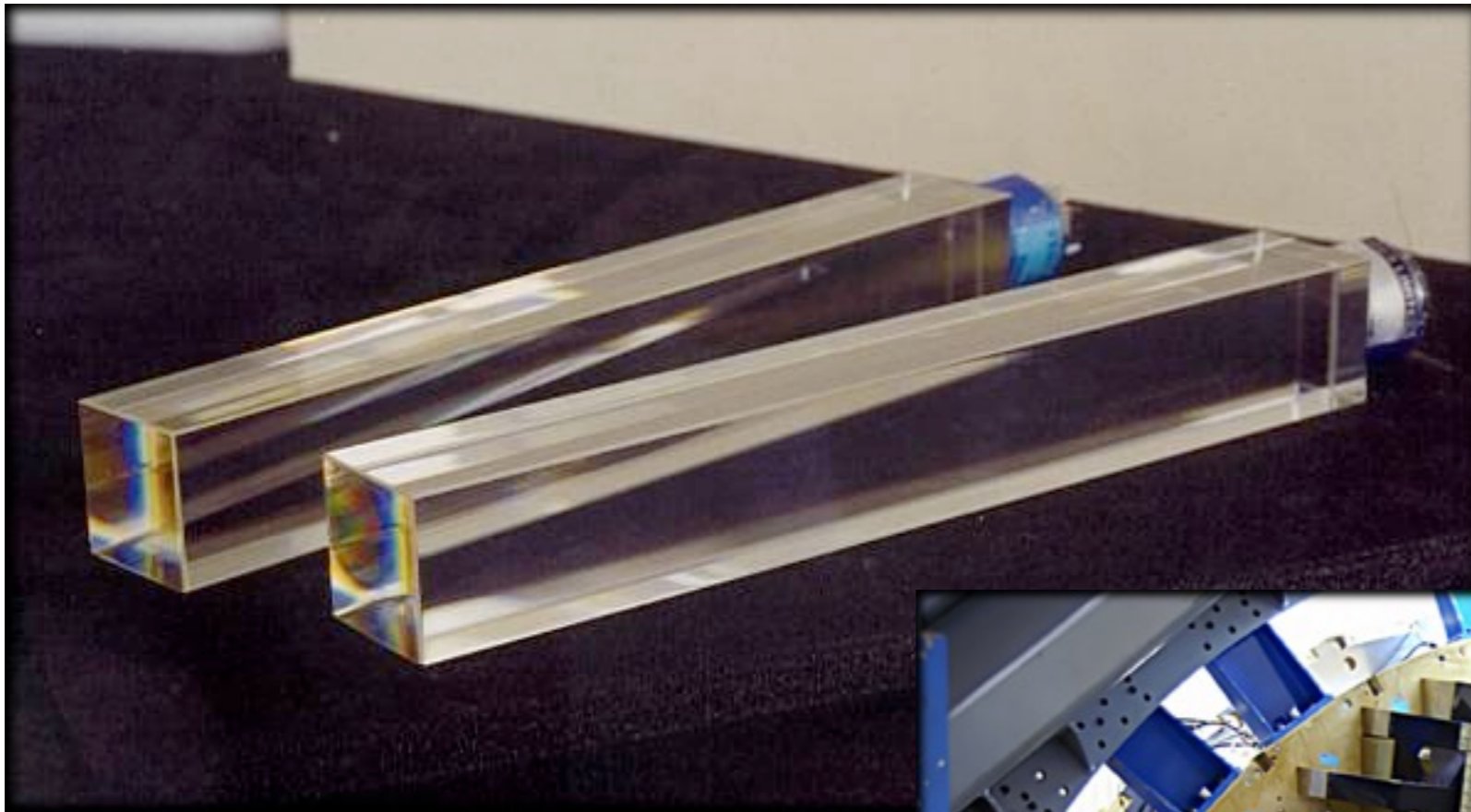
CMS Tracker under construction



Key challenges

- ▶ Power, cooling, cabling
- ▶ Alignment
- ▶ Radiation tolerance
- ▶ Beam safety

Detector Highlights: Calorimetry

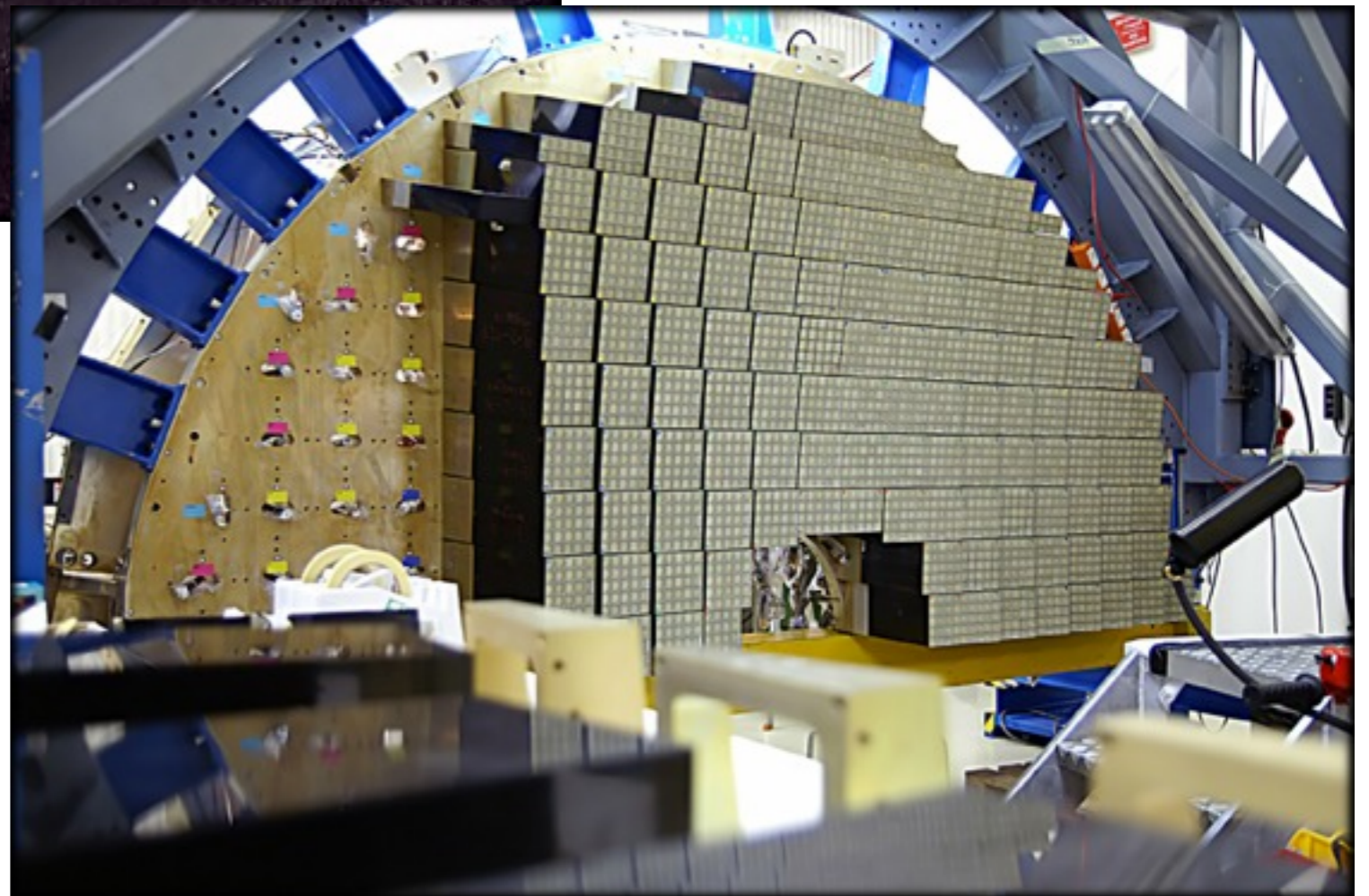


CMS PbWO₄ crystals:
effectively, transparent metal
(3kg each)

CMS endcap ECAL
assembly by robot

▶ Key challenges

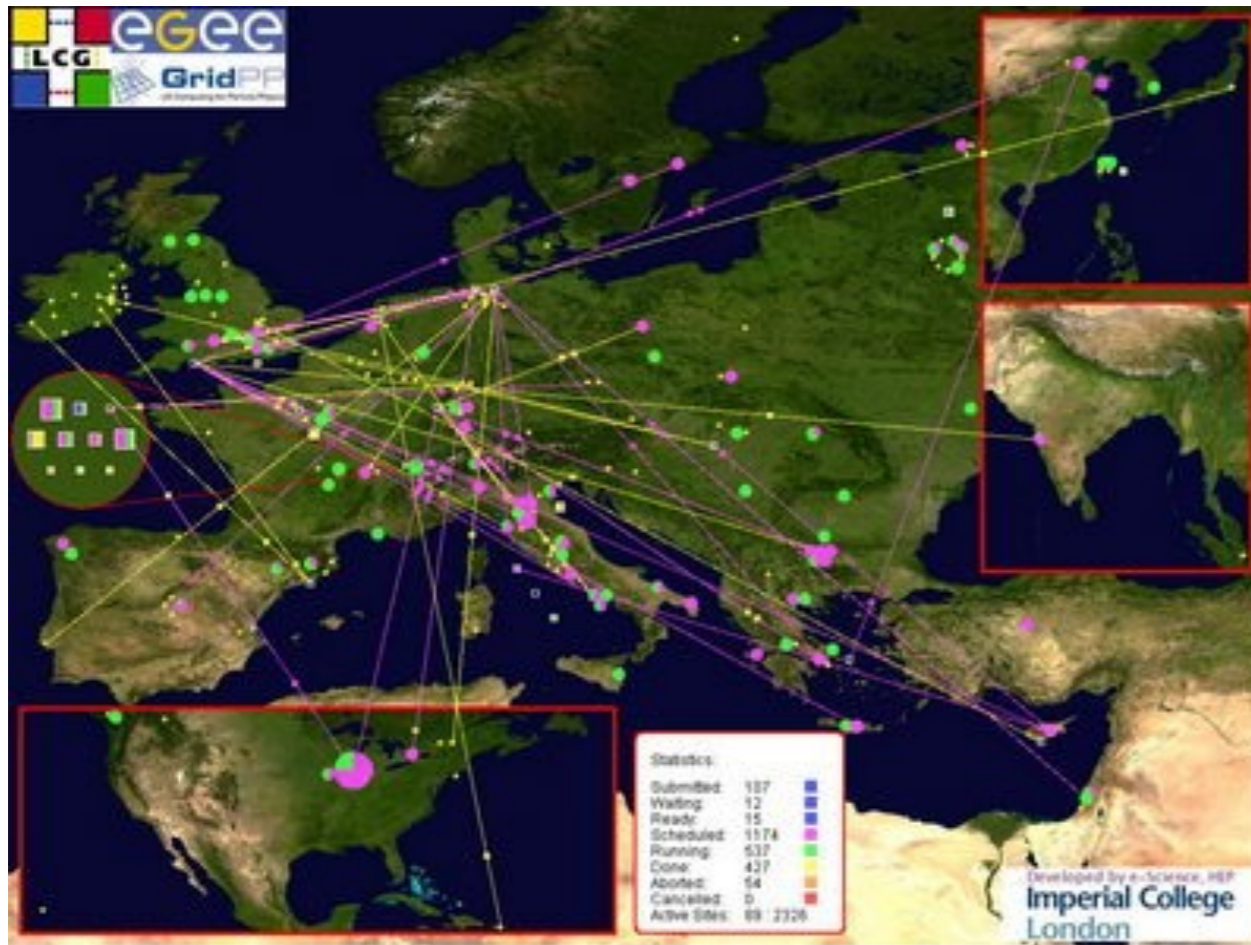
- ▶ Calibration (0.5% required)
- ▶ Stability
- ▶ Energy scale determination
- ▶ Maintenance and disposal



Level-I Trigger Strategy

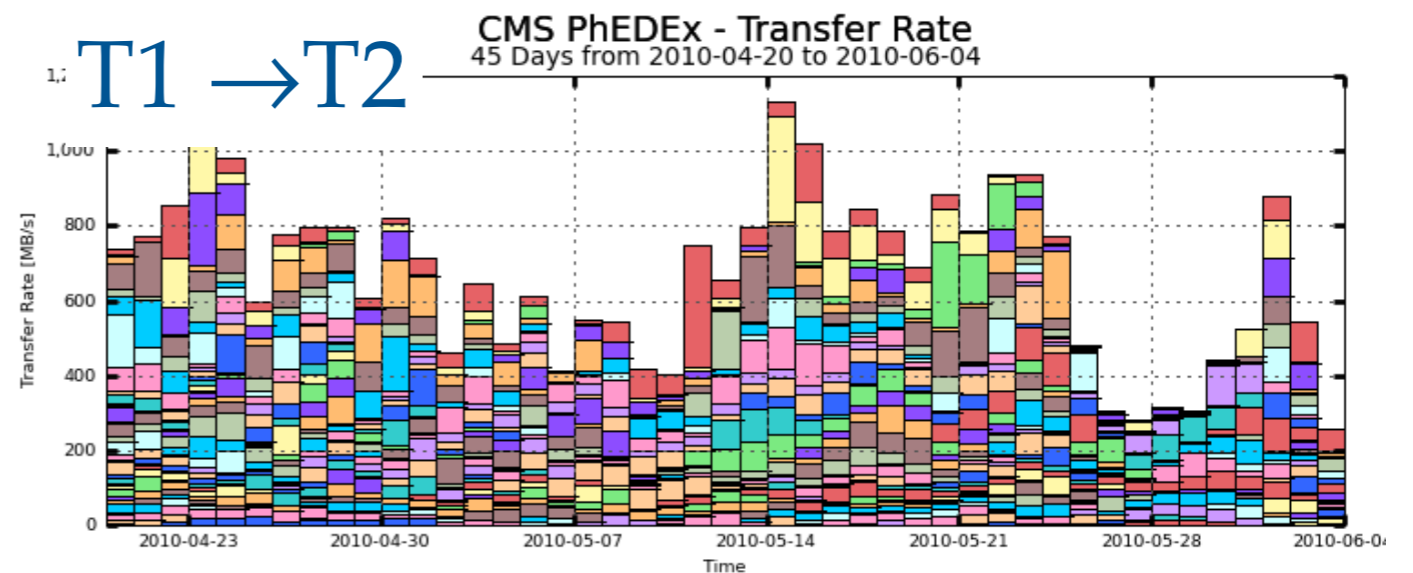
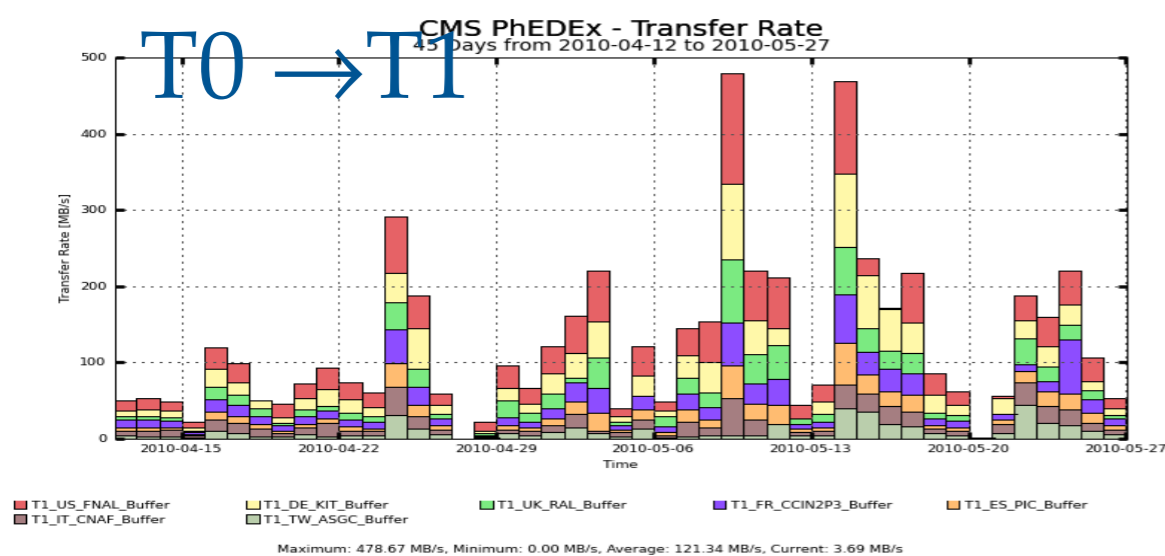
- ▶ Driven by LHC physics conditions
 - ▶ Heavy decays against “soft” QCD b/g; intermediate W / Z; H- \rightarrow $\gamma\gamma$
 - ▶ -> Identify high-pt leptons* and photons (*including τ)
 - ▶ Low pt thresholds motivated by efficiency for W / Z / light Higgs
- ▶ Trigger combinations
 - ▶ $>20\text{GeV}$ limit on single-lepton thresholds due to quark decay + π^0 b/g
 - ▶ Can use lower thresholds for objects in combination (e.g. dileptons)
 - ▶ -> Find trigger objects locally, combine and cut only at last stage
- ▶ Large uncertainties in background (and perhaps signal)
 - ▶ Flexibility and control of rate are both vital
 - ▶ -> All trigger thresholds and conditions must be programmable
 - ▶ Trigger architecture is fixed, but this is a function of detector geometry
- ▶ Must have high and well-understood efficiency
 - ▶ -> Need to include overlapping and minbias triggers to measure ϵ

Data Handling

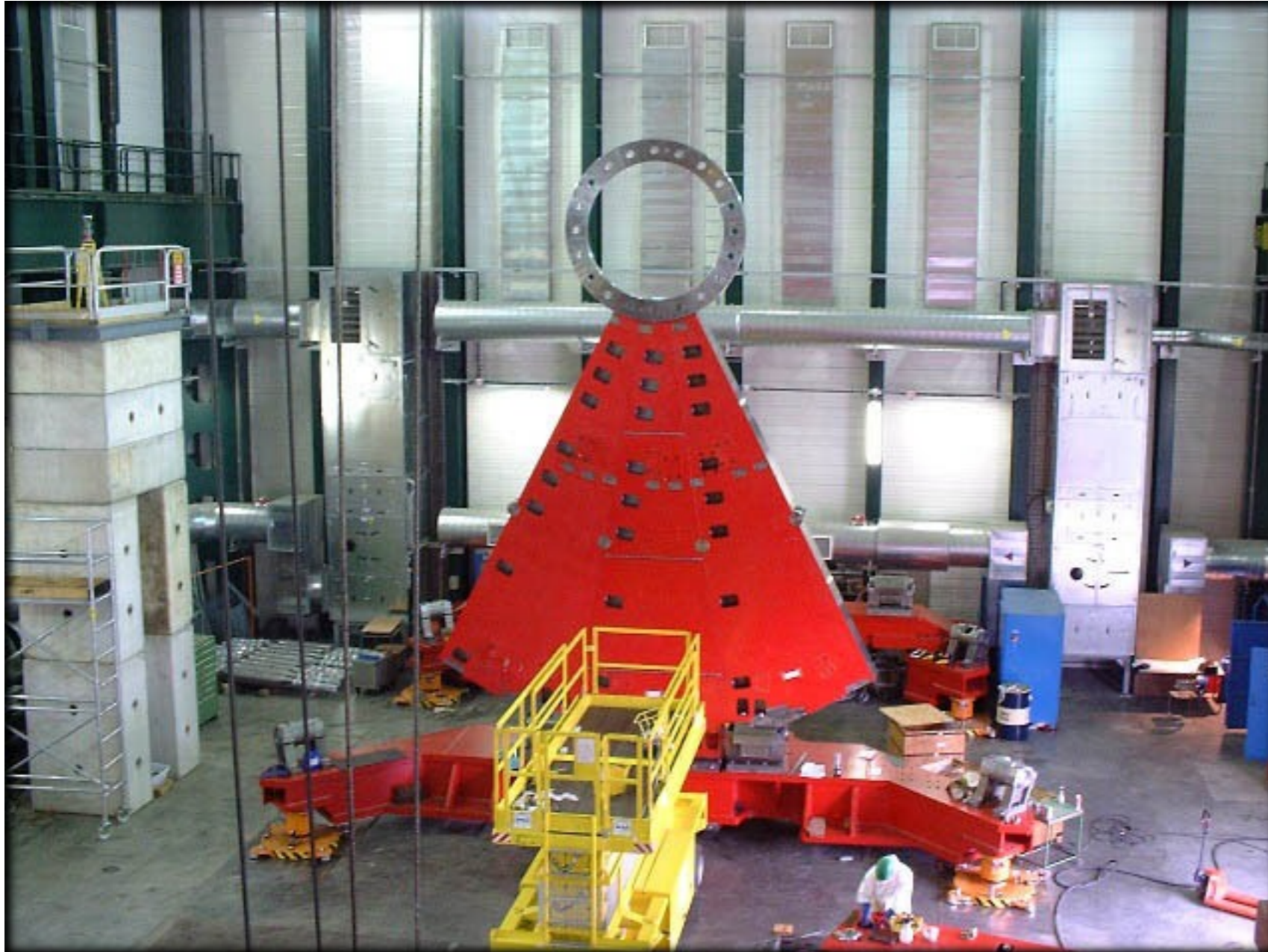


- ▶ Grid computing system
 - ▶ Built expressly for LHC data handling
 - ▶ World's largest distributed system
 - ▶ Dataflows > 400Gb/s partly on private optics
 - ▶ 100,000+ CPUs online

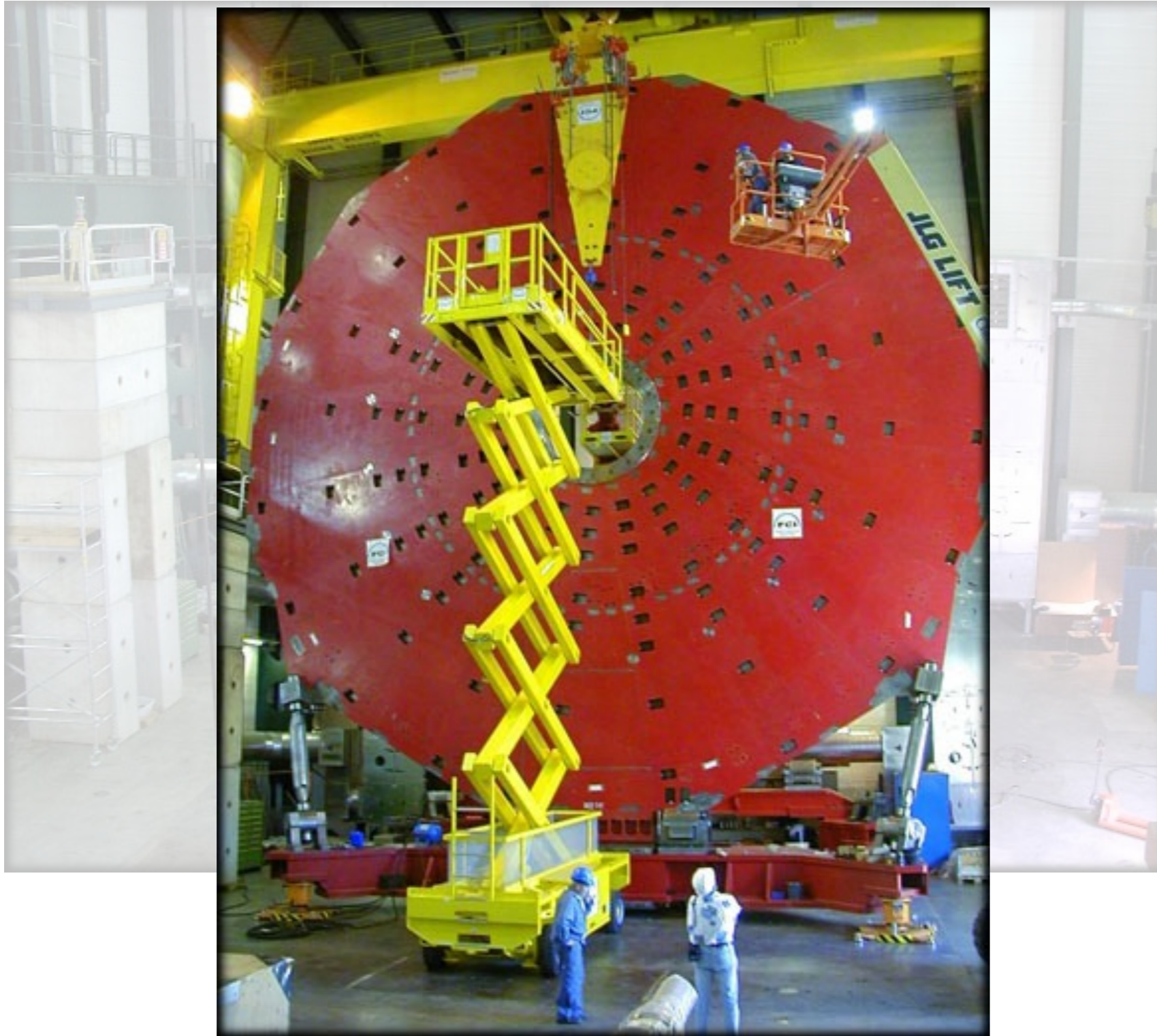
Data transfers worldwide



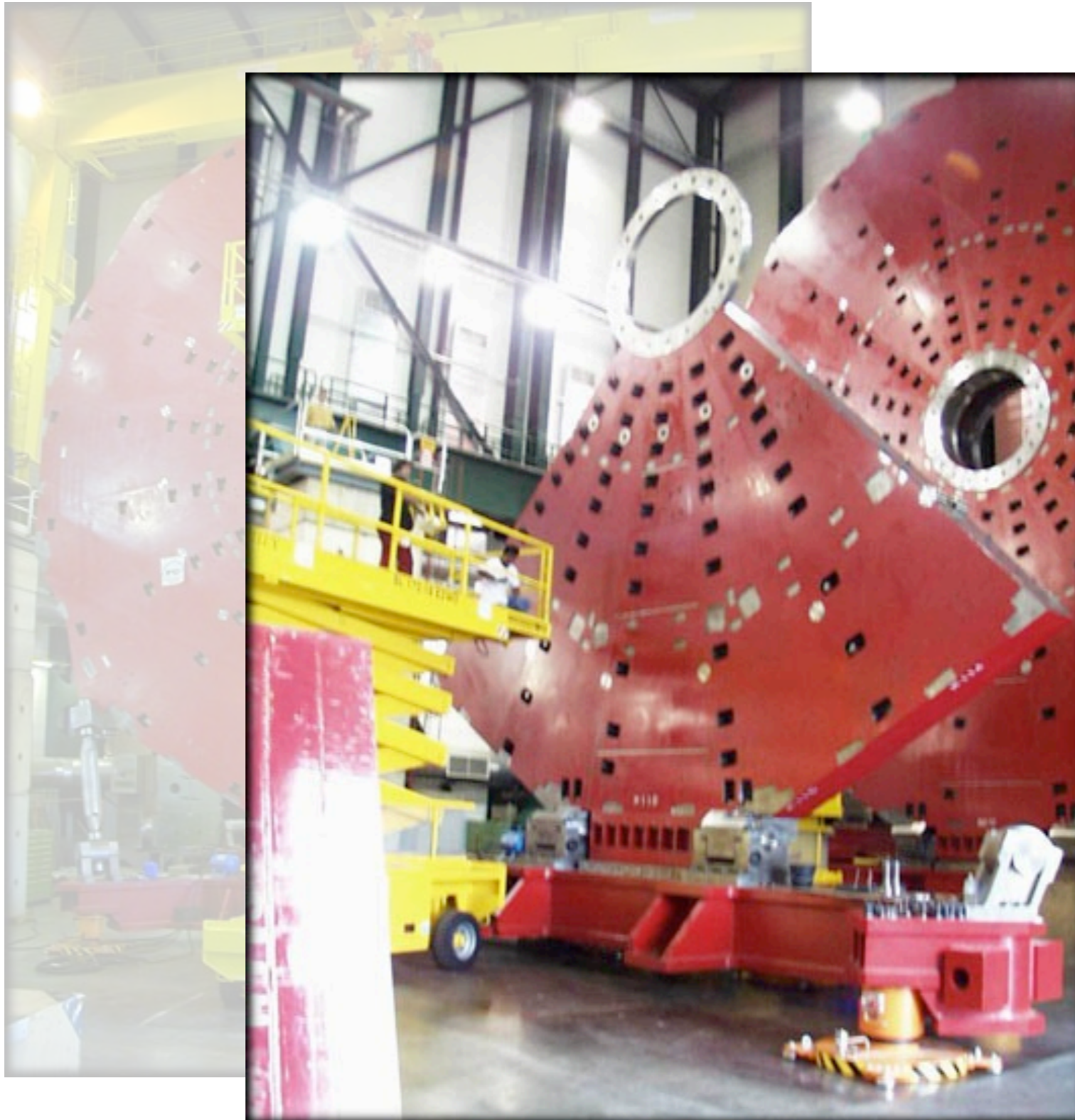
Construction of CMS



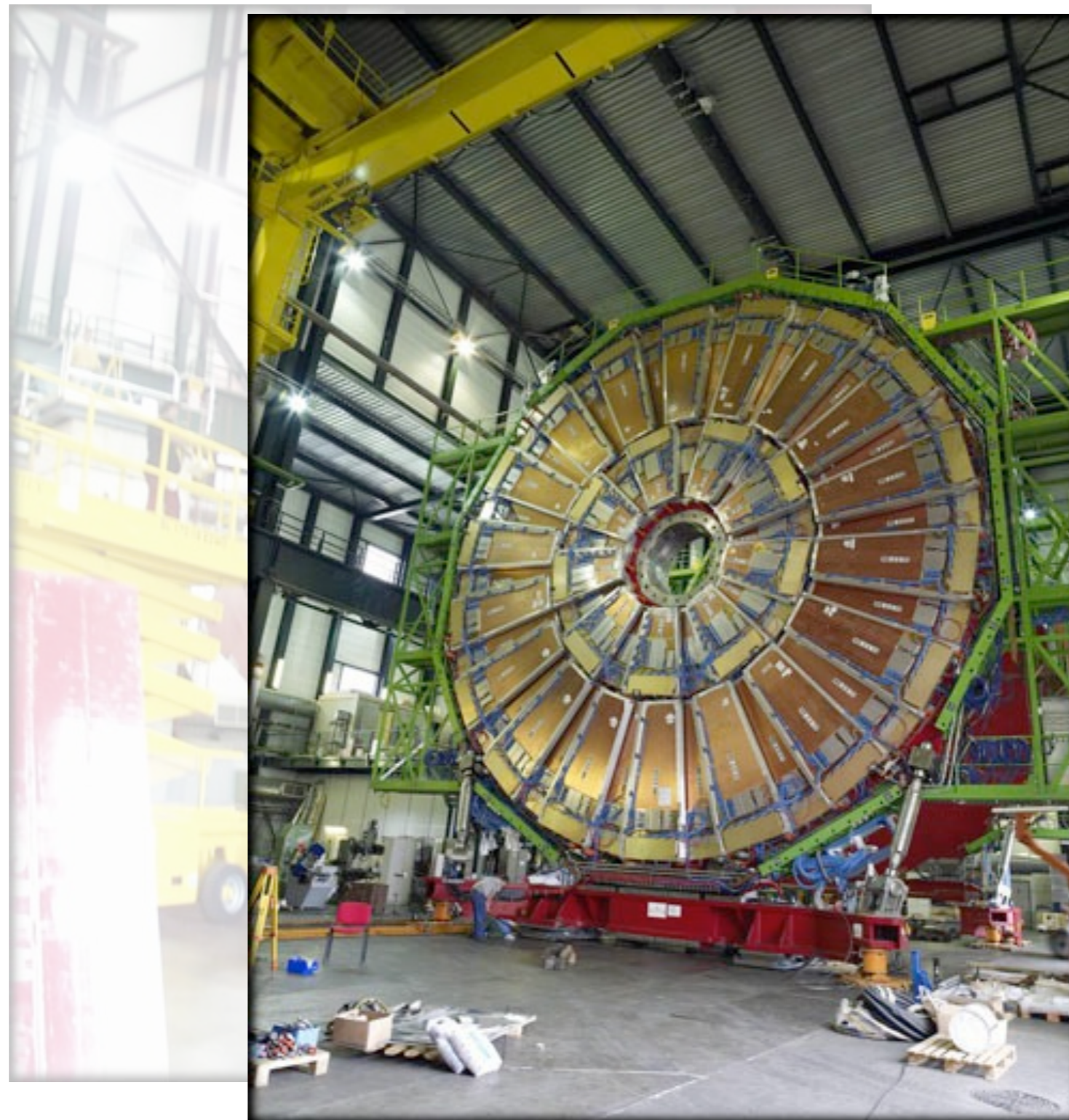
Construction of CMS



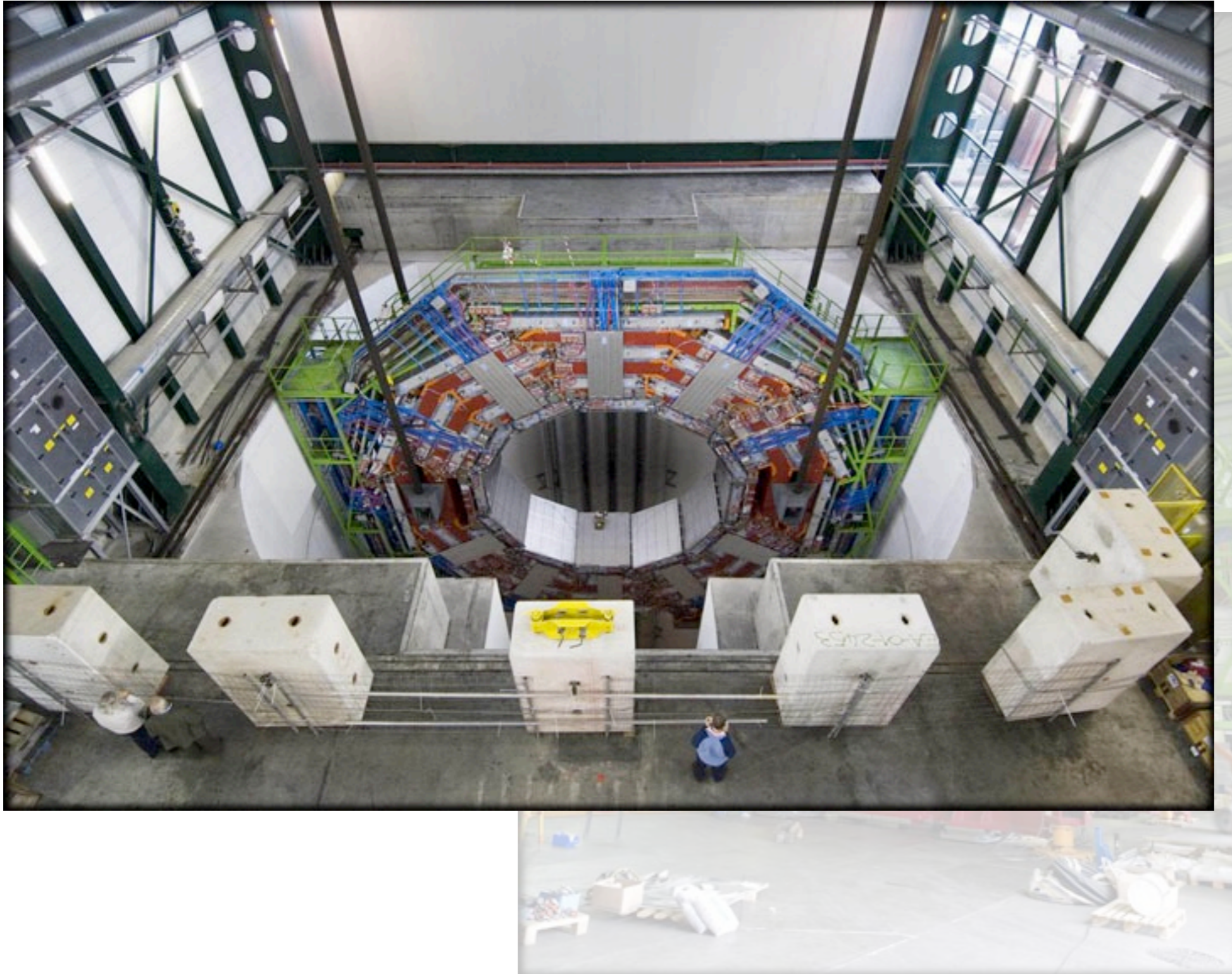
Construction of CMS



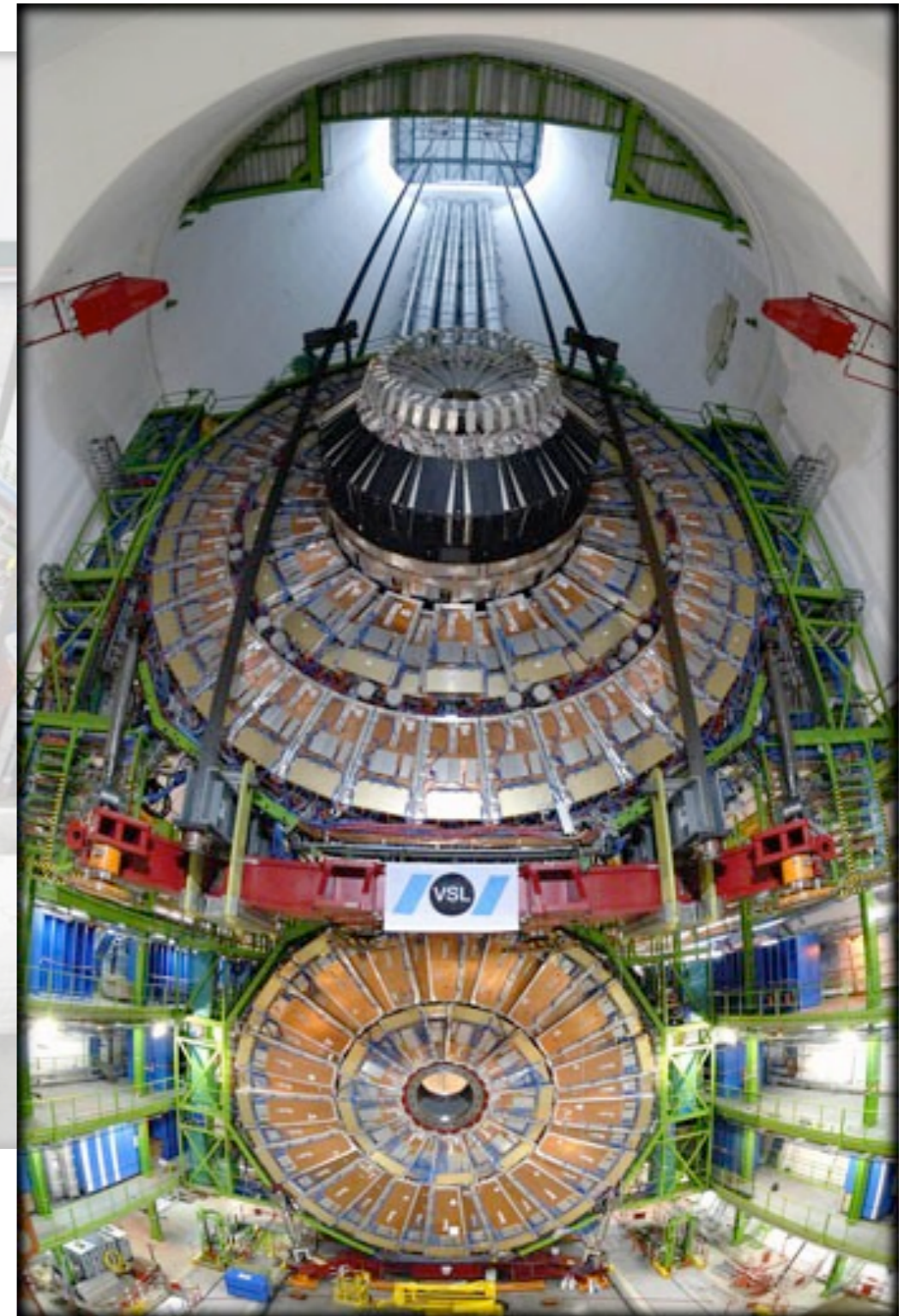
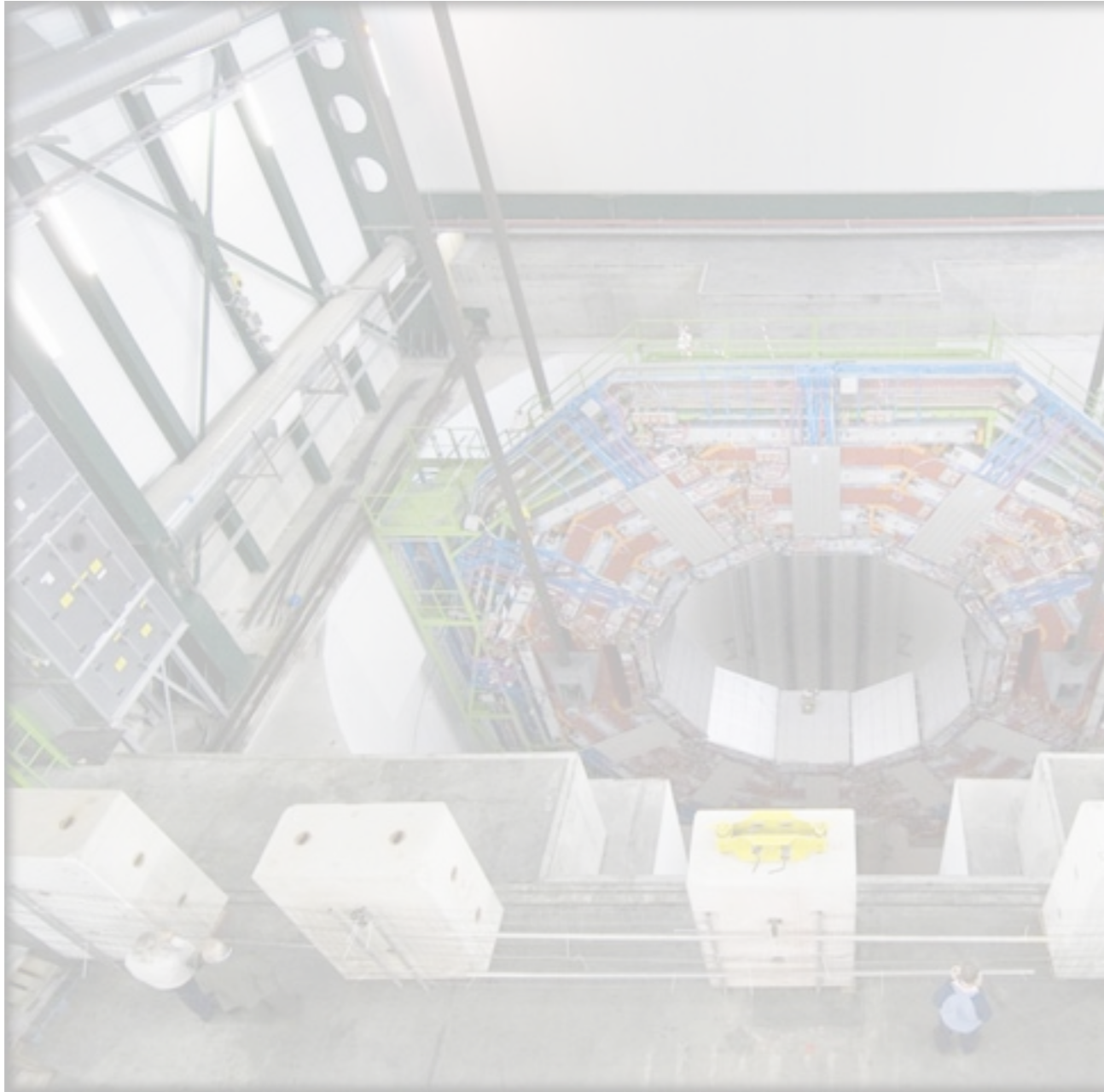
Construction of CMS



Construction of CMS

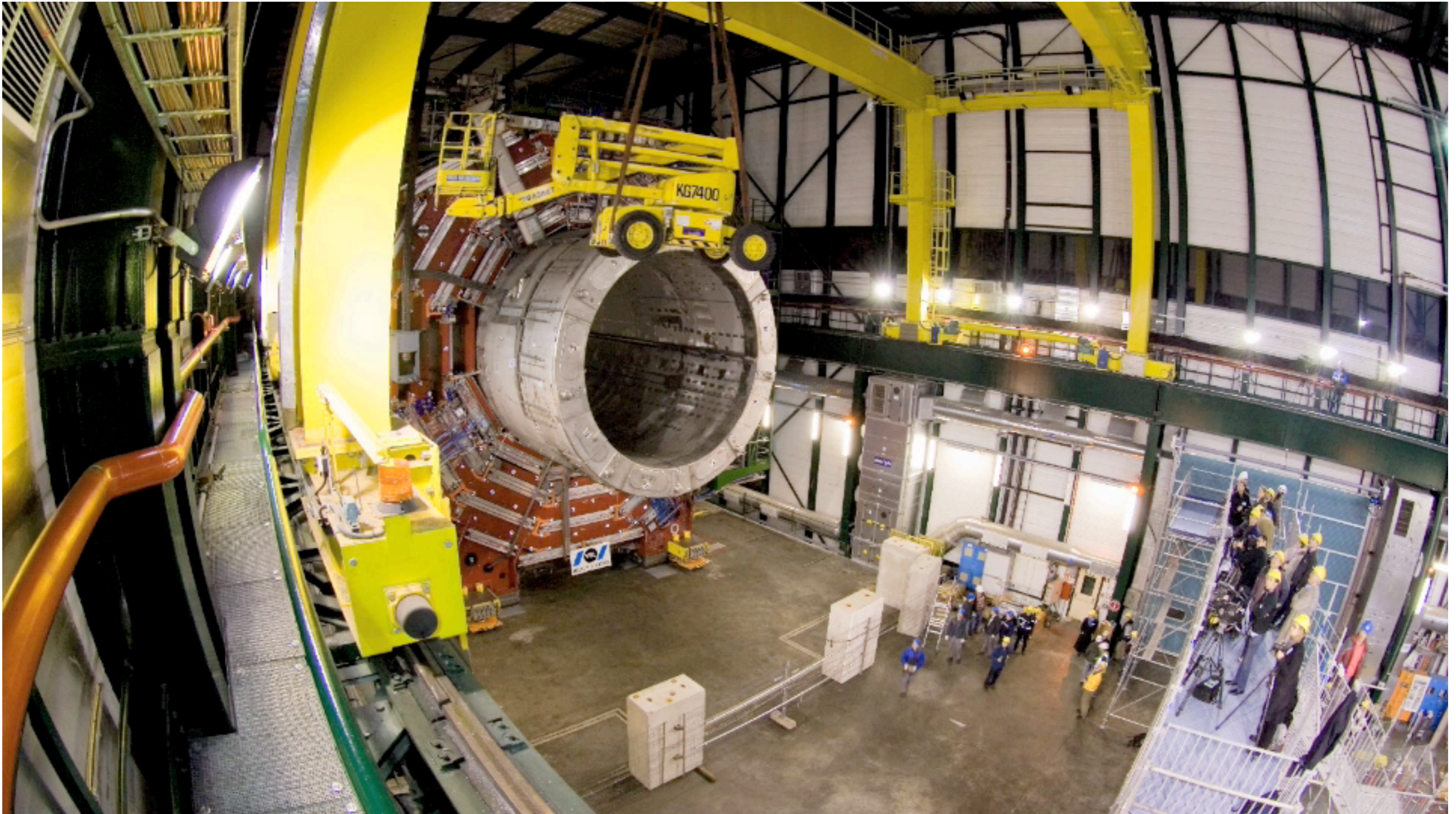


Construction of CMS

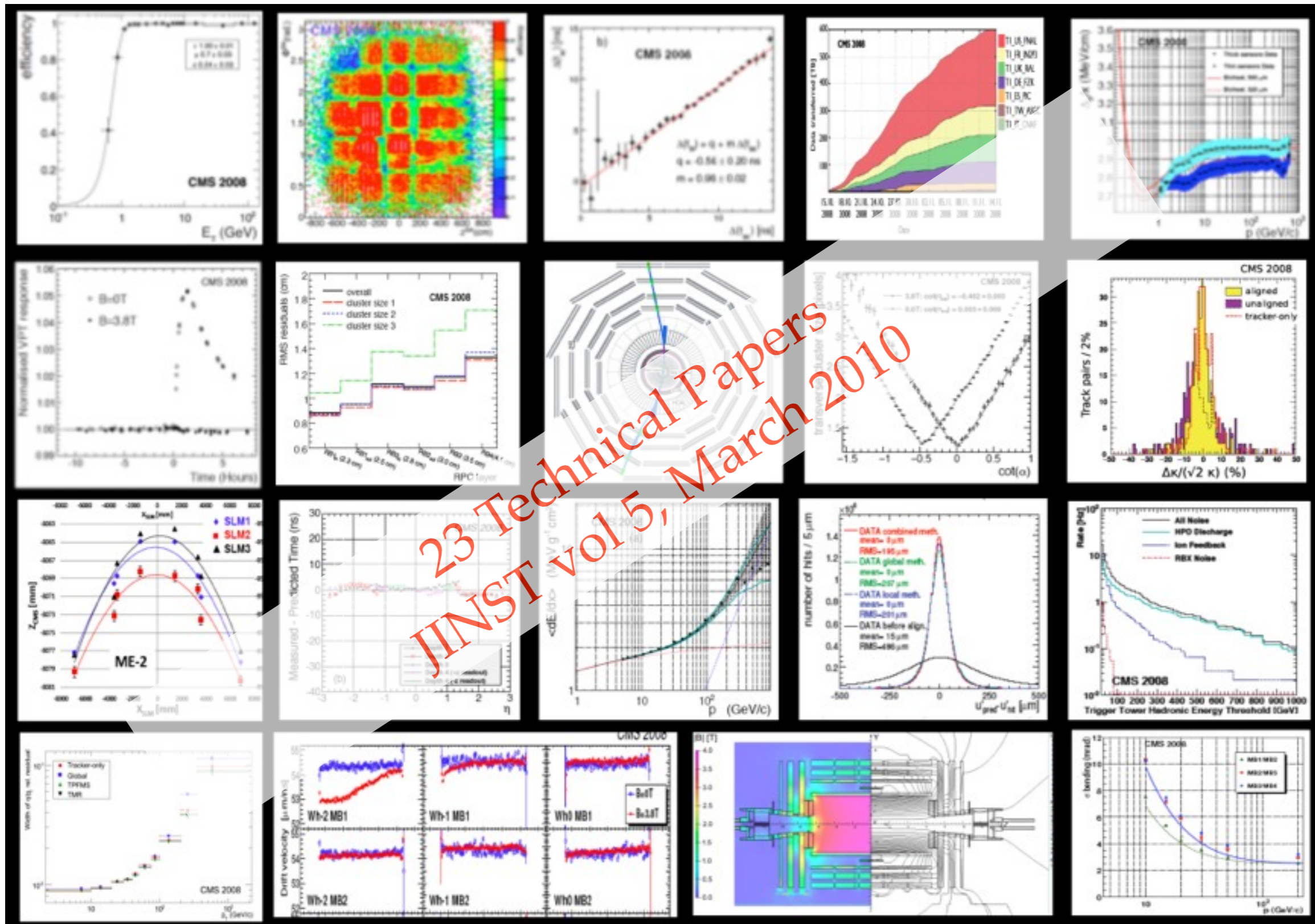


Going Underground

Going Underground

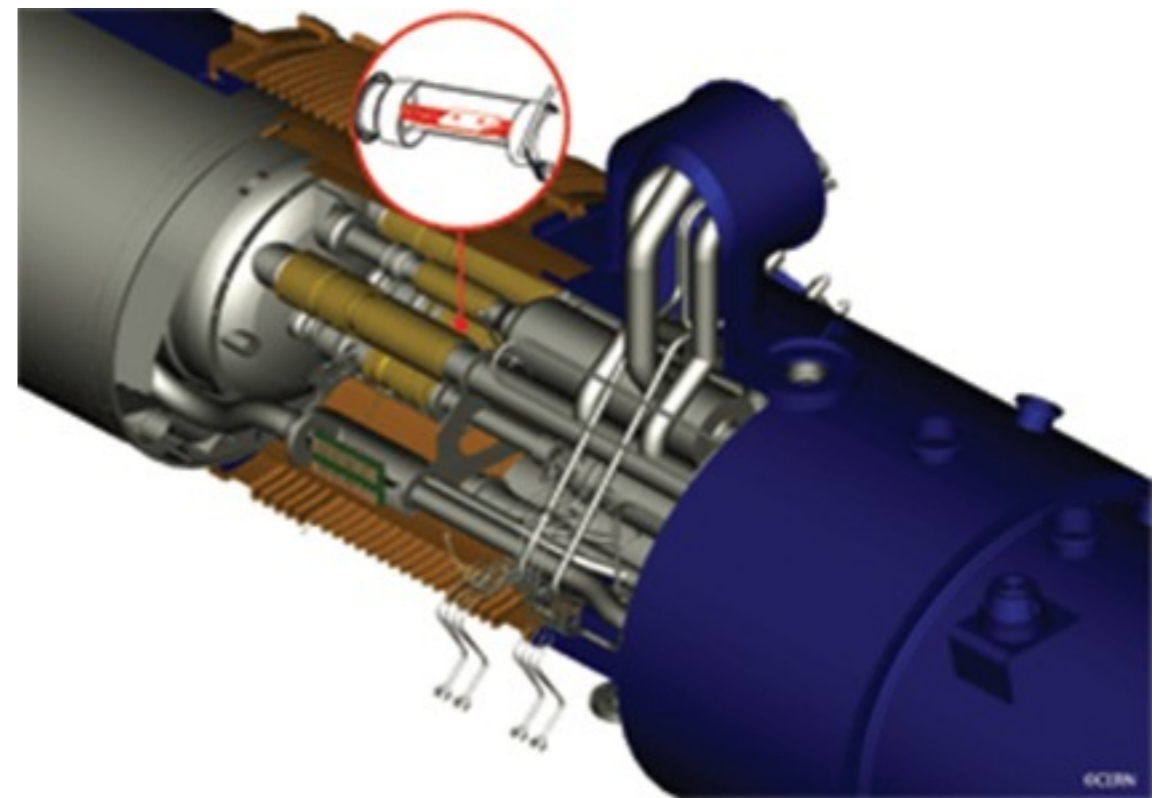
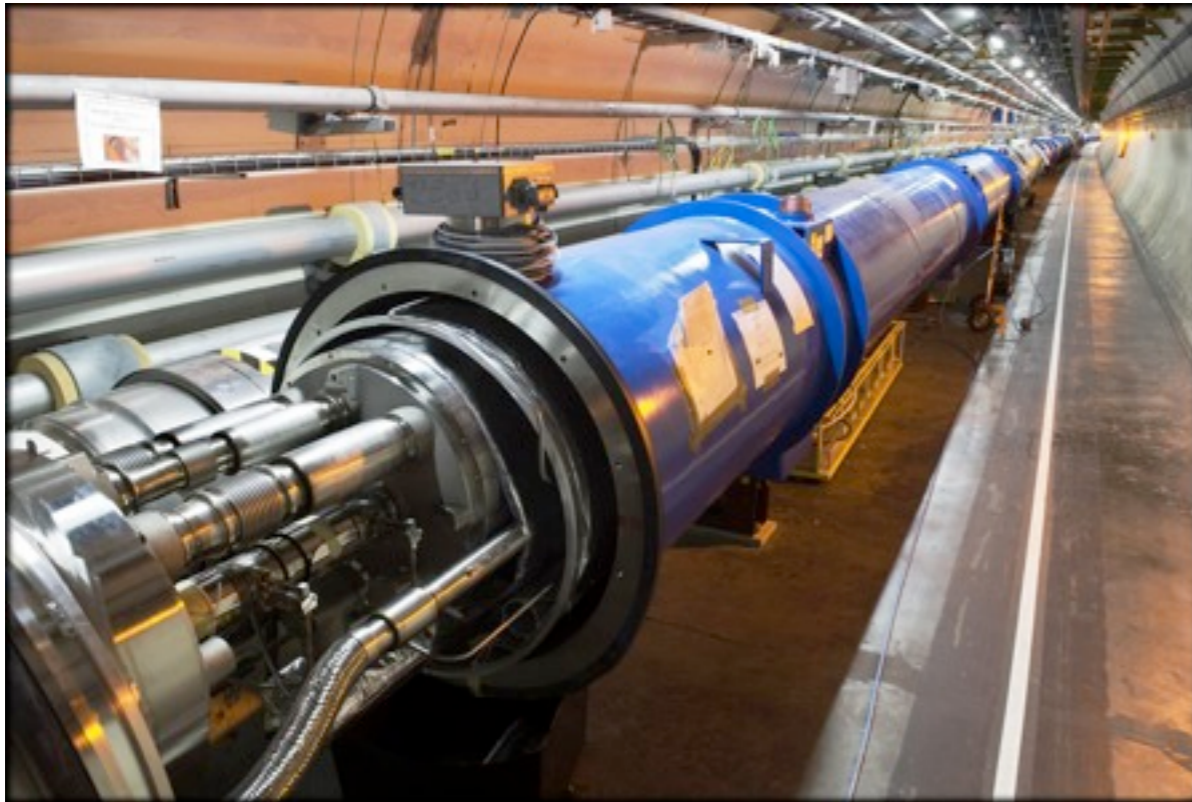


Before Collisions (~1950's Physics)



23 Technical Papers
JINST vol 5, March 2010

LHC Startup: Annus Horribilis

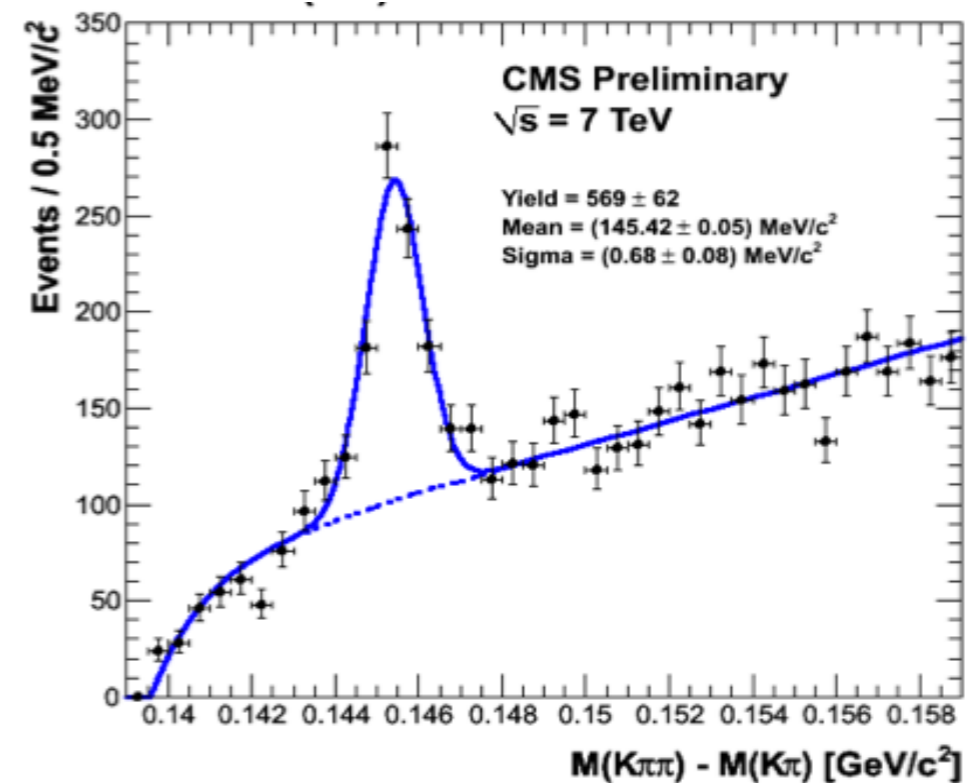
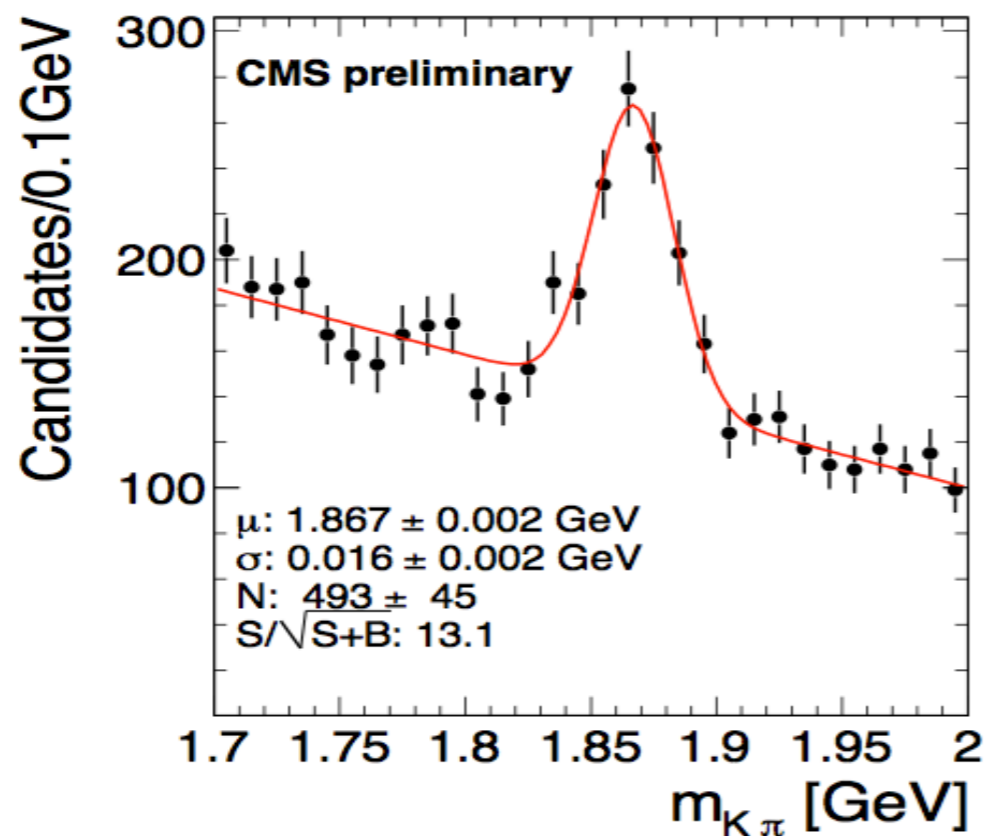
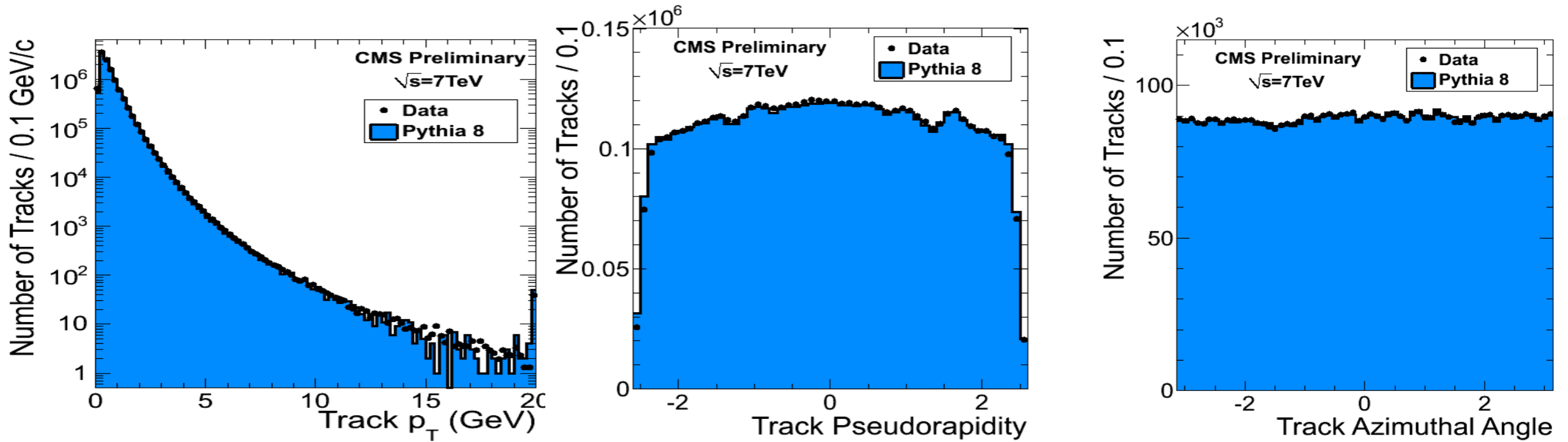


- ▶ LHC starts spectacularly on Sep 10th '08
 - ▶ LHC fail occurs on Sep 19th
 - ▶ Tiny imperfection in a soldered joint caused an electrical arc
 - ▶ Helium released, but safety systems prevented further damage
 - ▶ Remember, beam energies in 2011 $> 100\text{MJ}$ ($\sim 1\text{TW}$ on a target)
- ▶ Fixing the machine is like a mission into space – 18 months

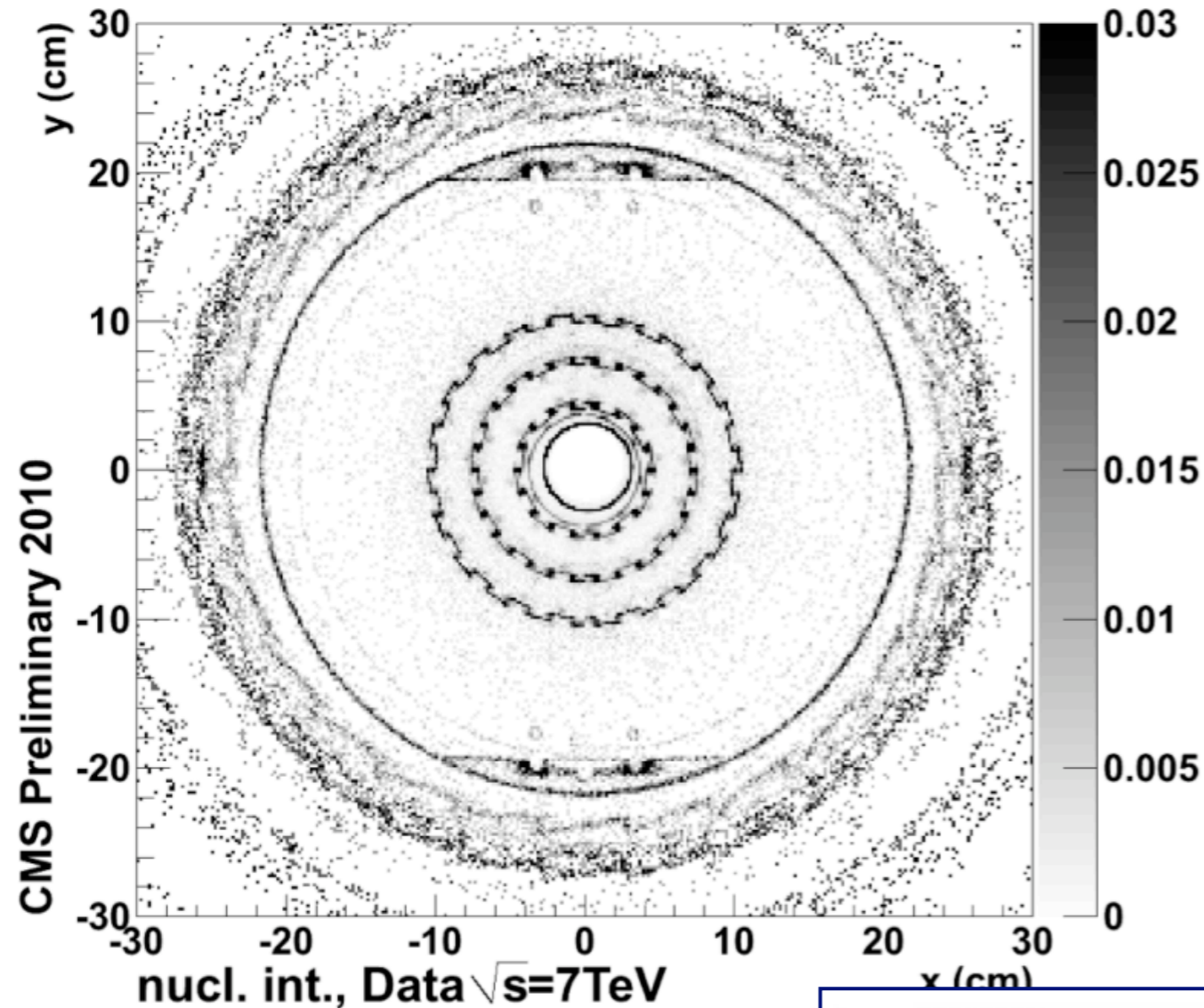
The Wrong Kind of Big Bang



Understanding Detector: Tracking



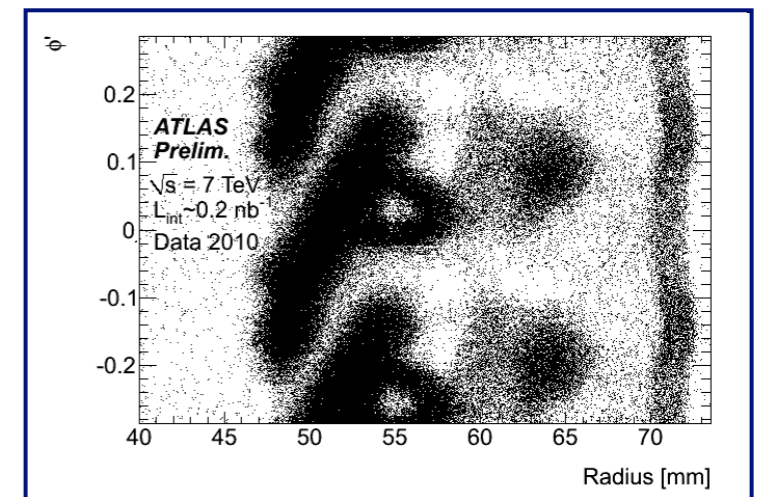
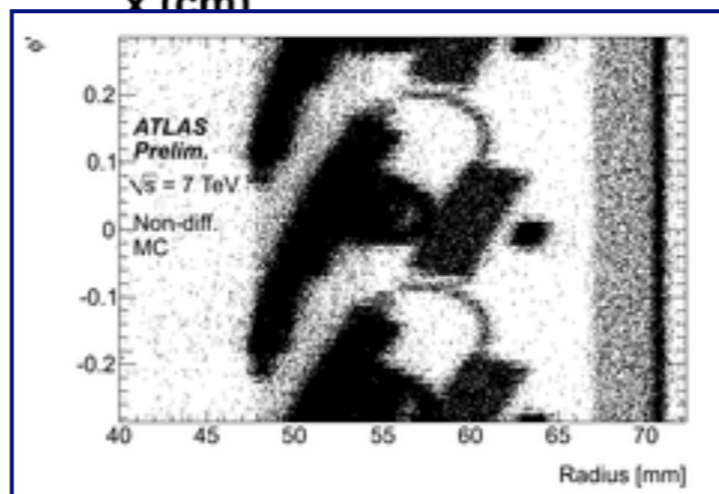
Understanding Detector: Material Plots



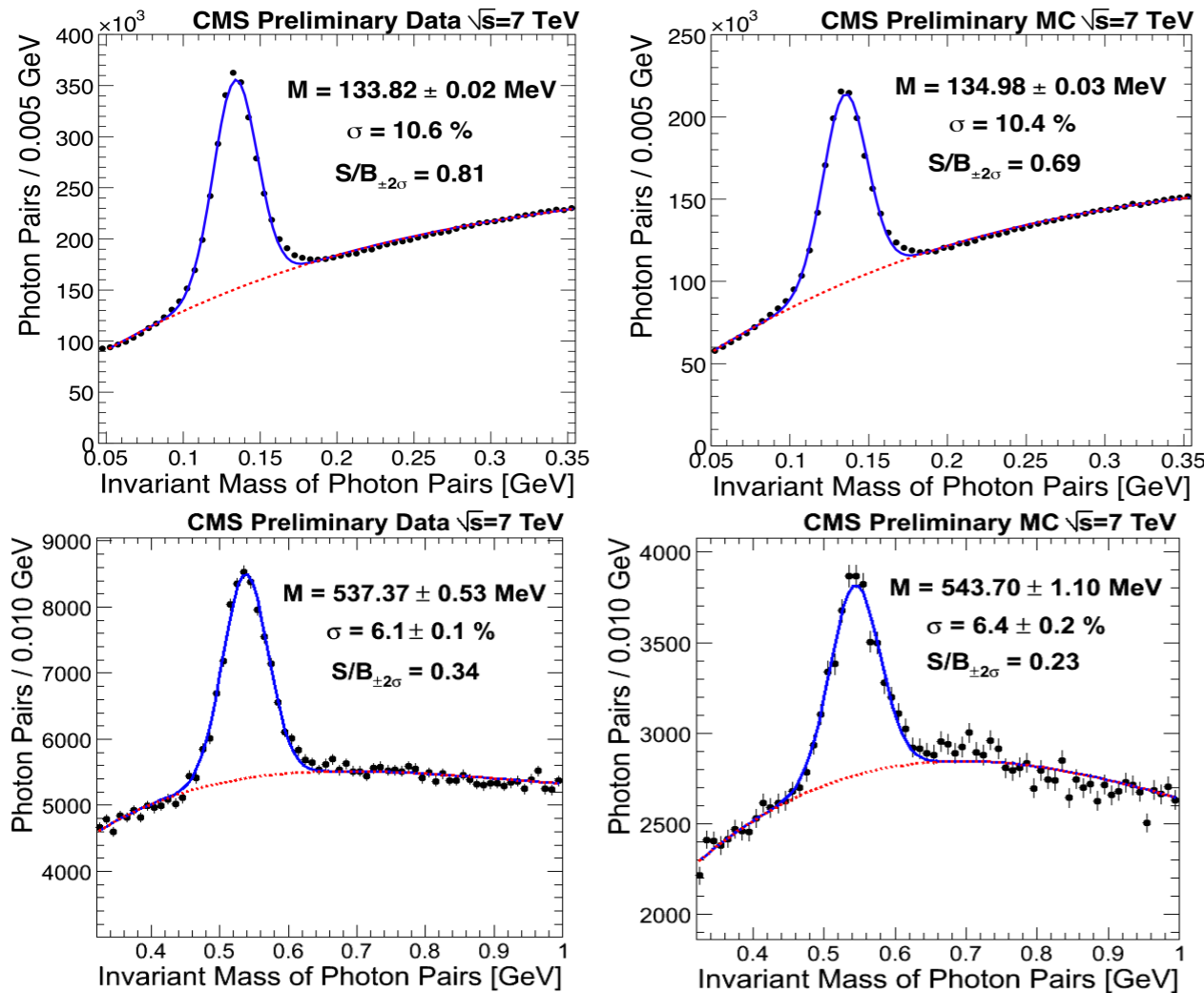
- ▶ Material distribution
 - ▶ Crucial test of detector MC accuracy
 - ▶ Essential in recovery of photon conversions
- ▶ Plot EM (γ) and hadronic interaction vertices

MC Sim

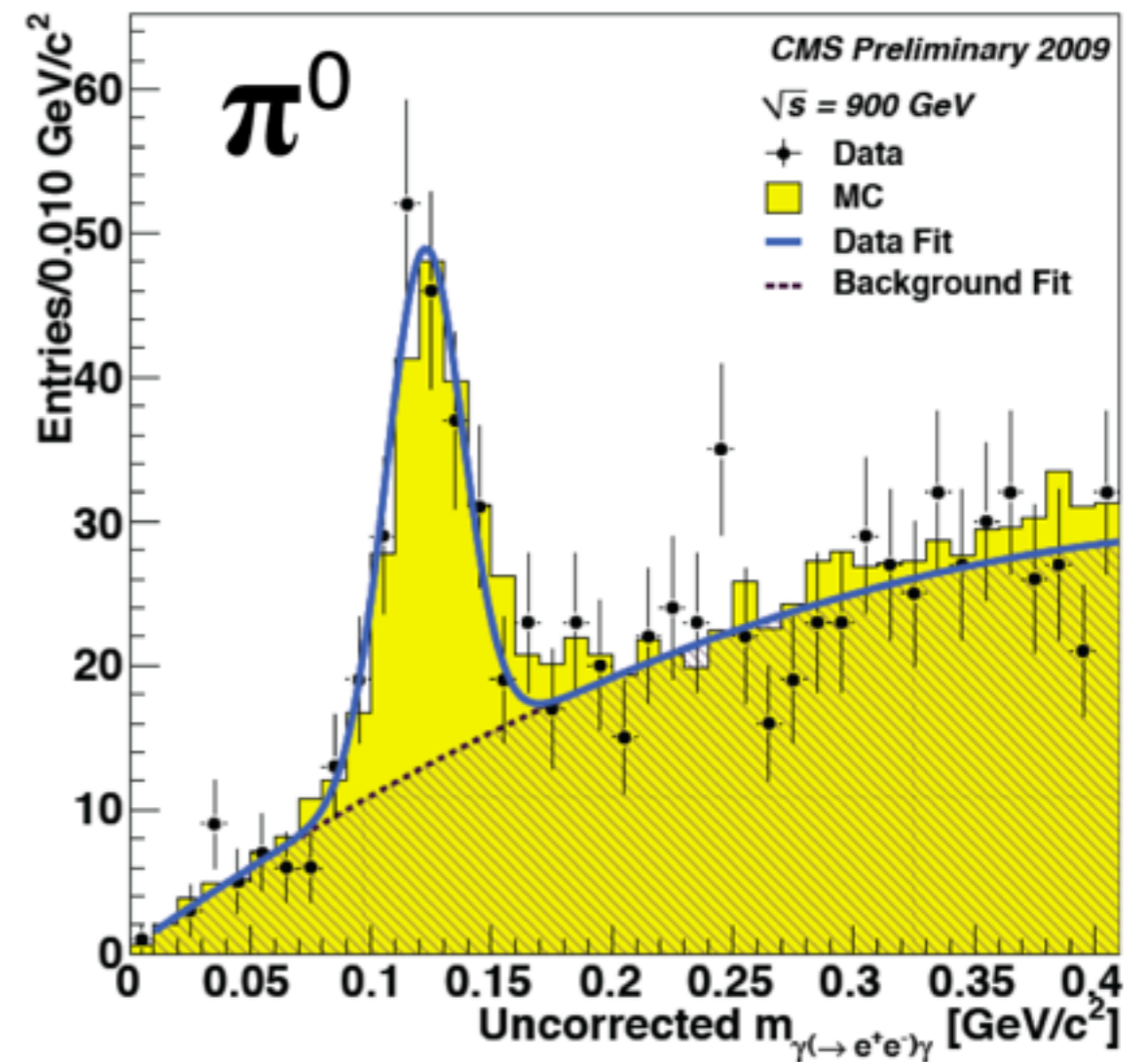
Data



Understanding Detector: ECAL

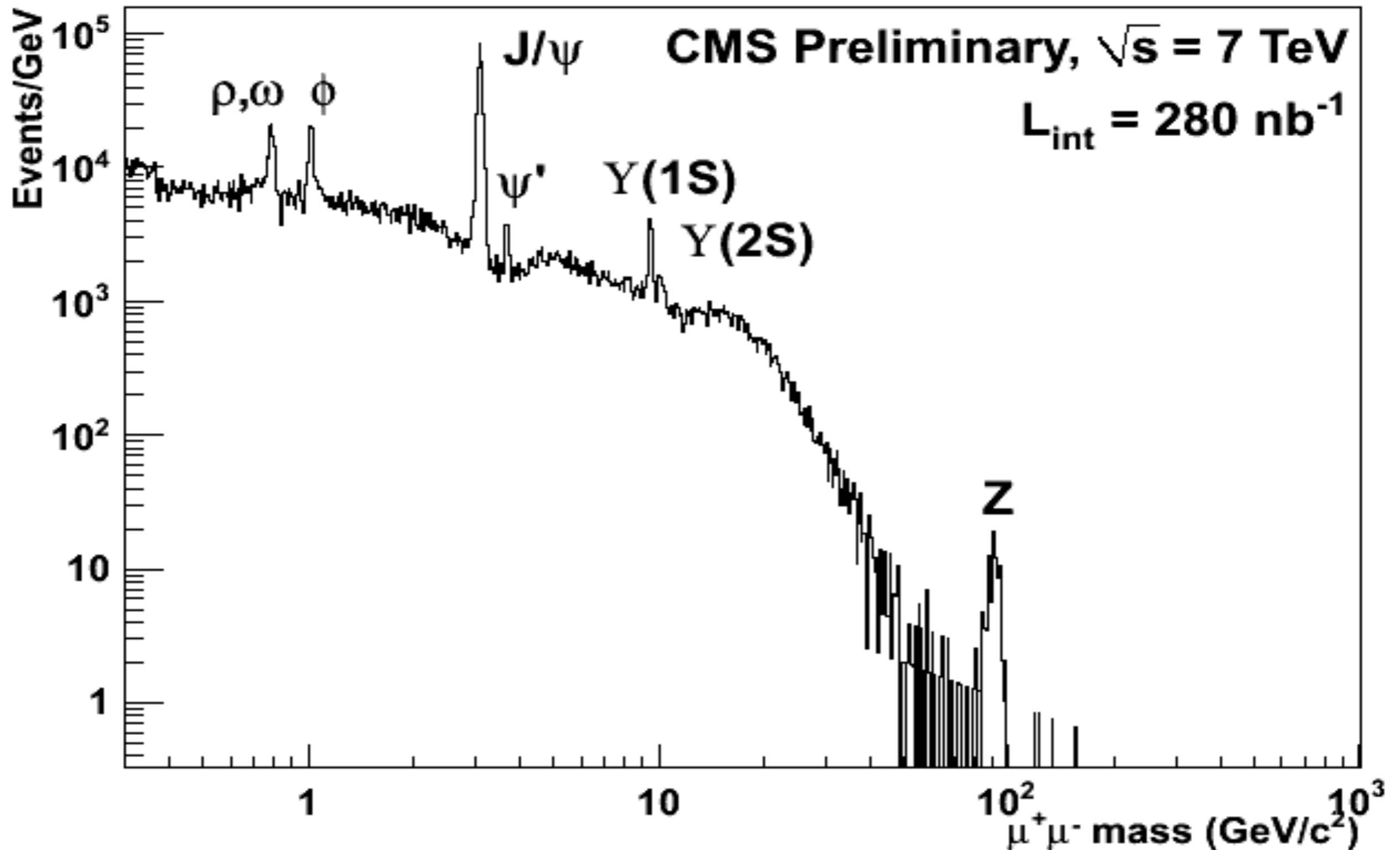


With one photon converted!

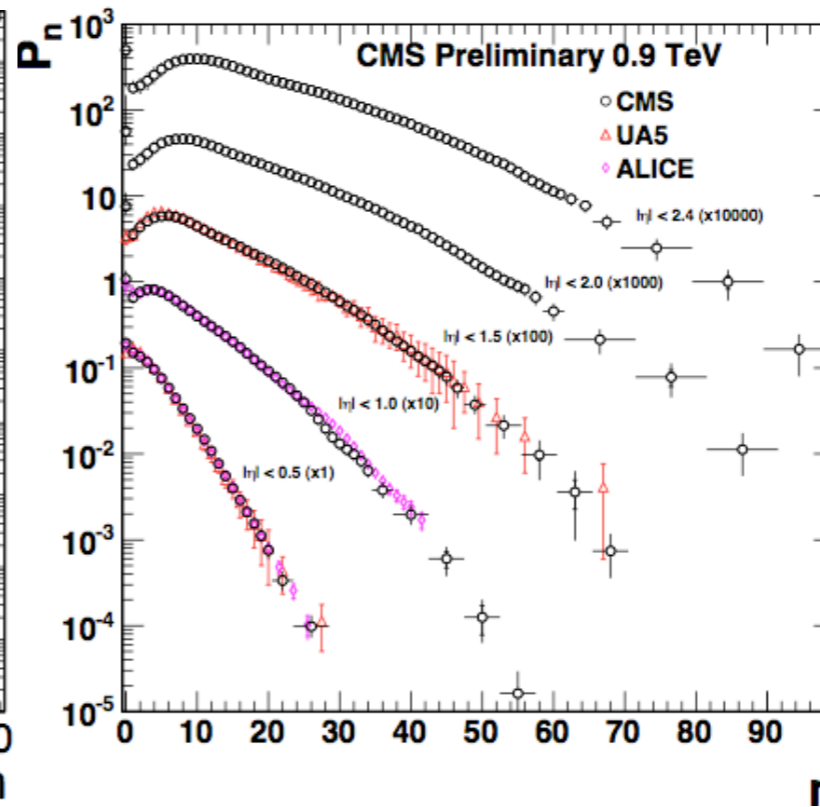
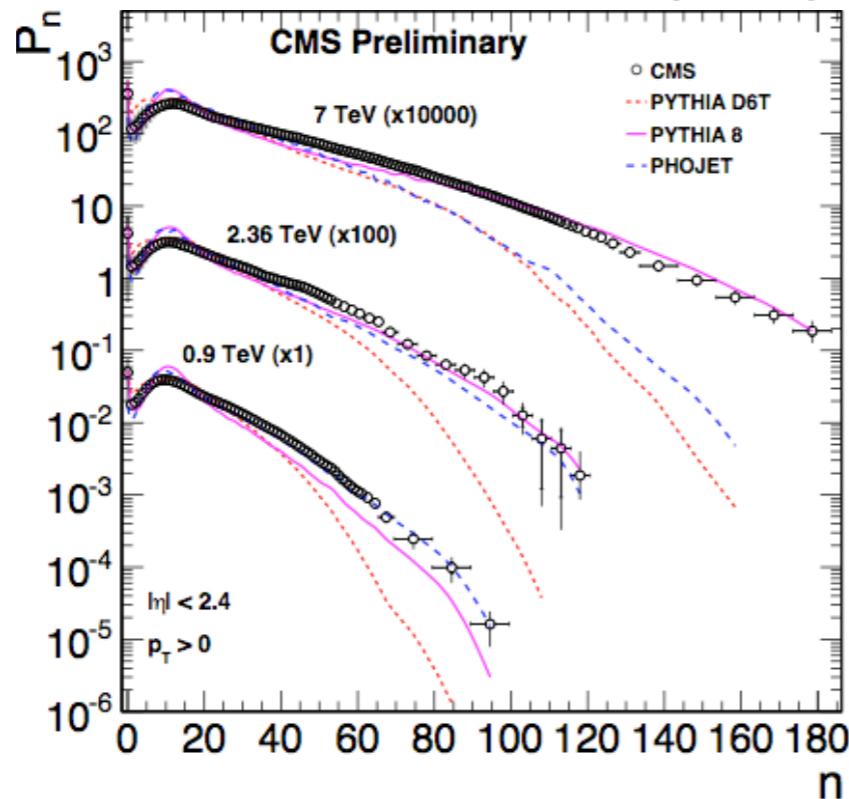
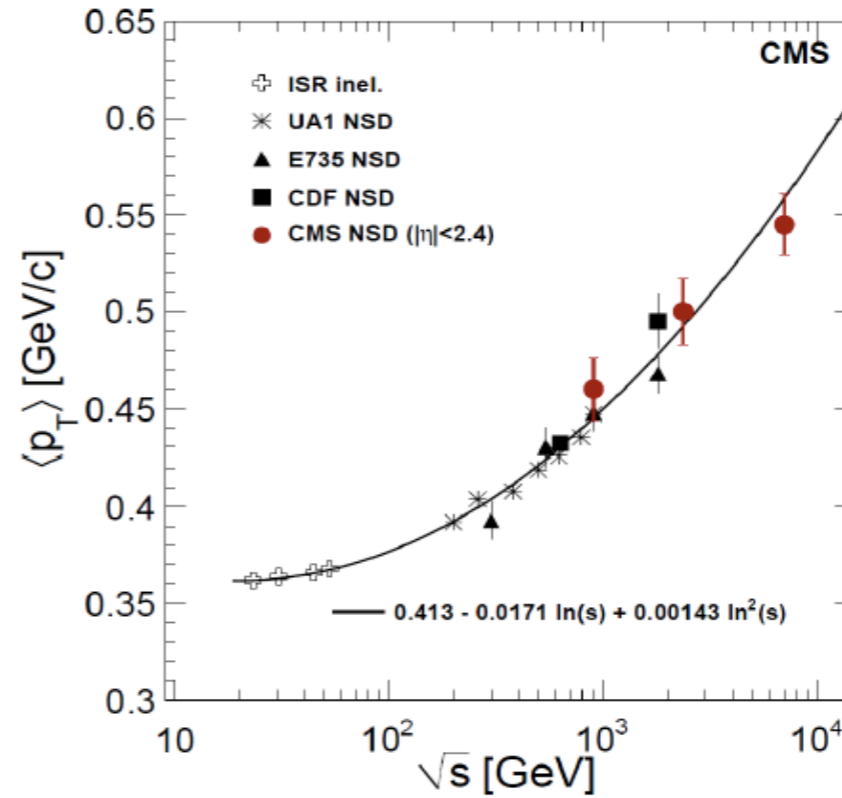
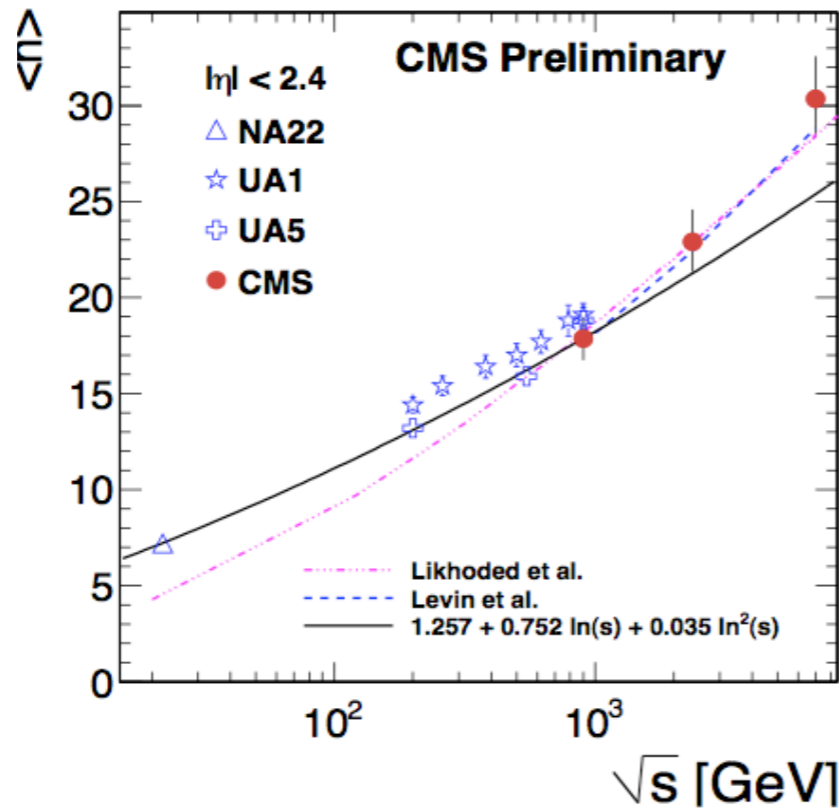


- ▶ Excellent agreement with MC
 - ▶ Calibration and energy scale are already good, and will improve

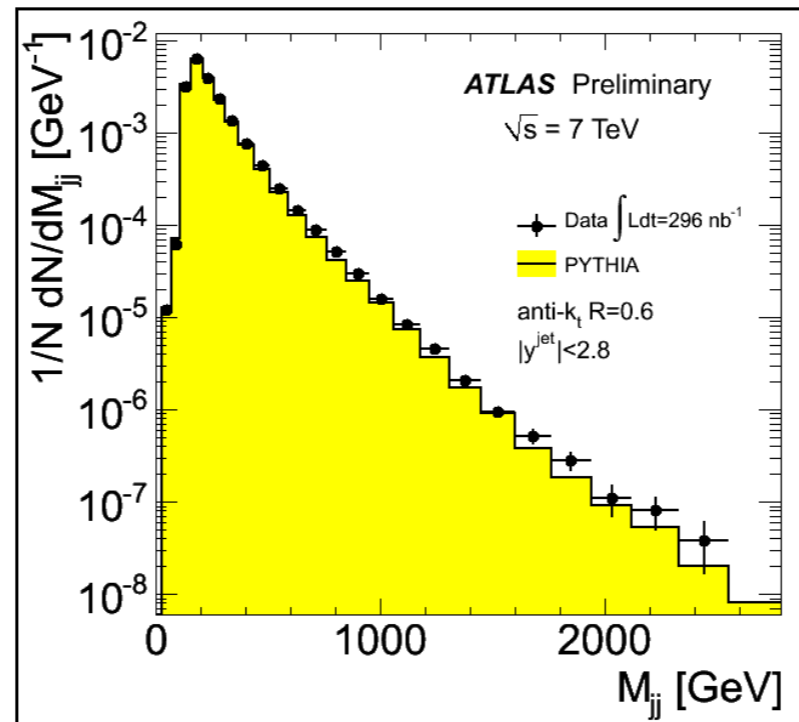
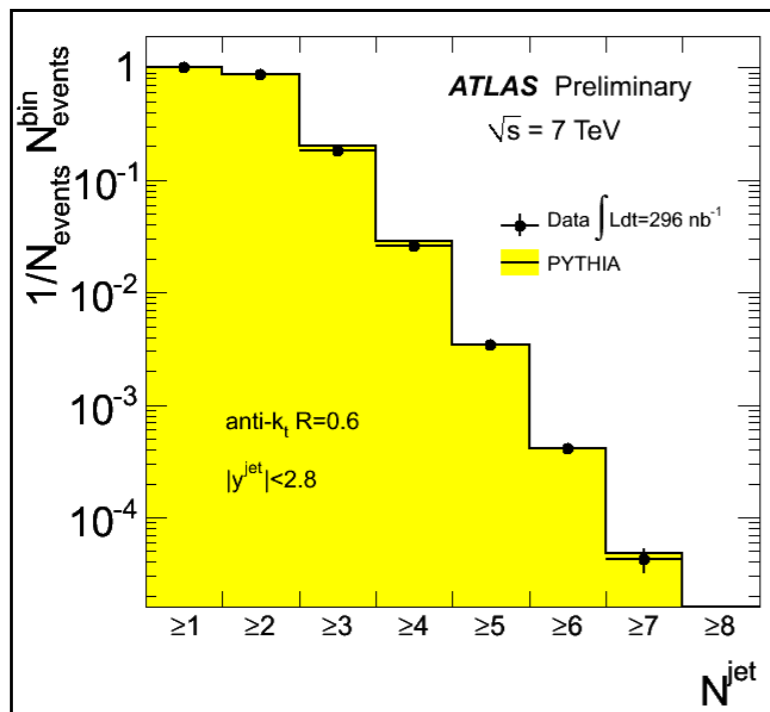
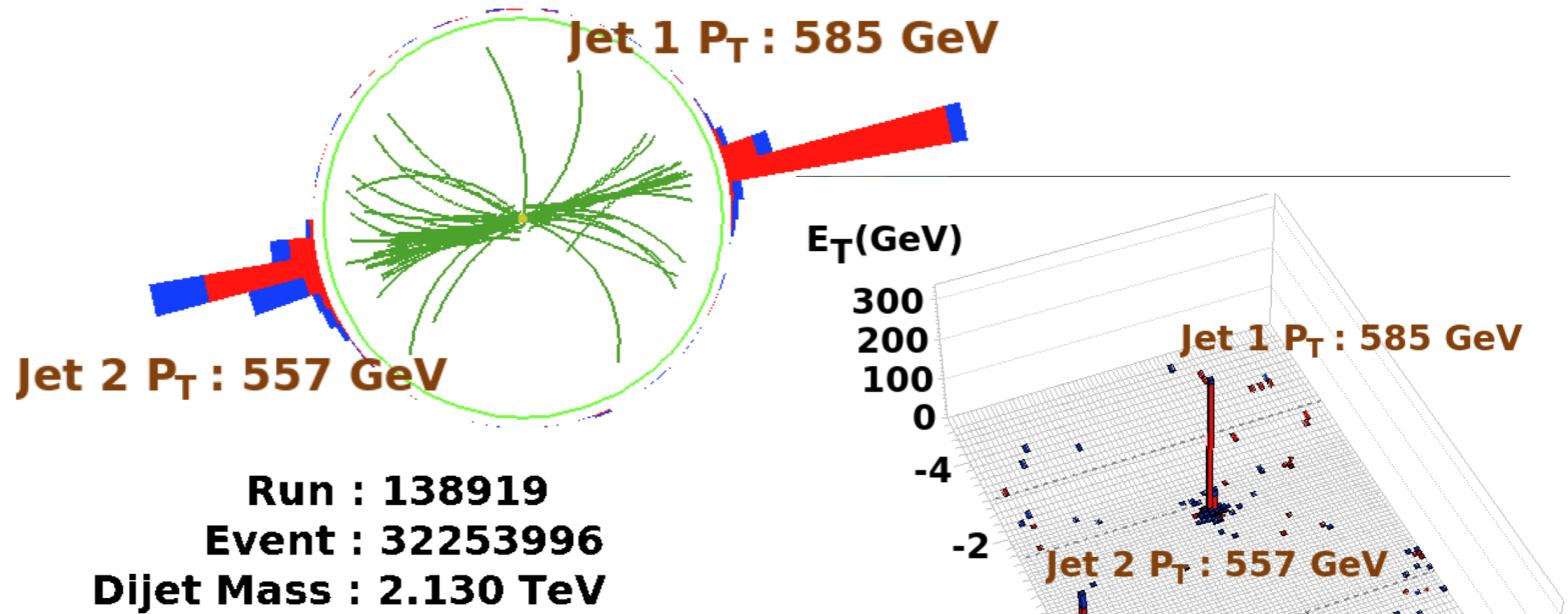
Understanding Detector: Muons



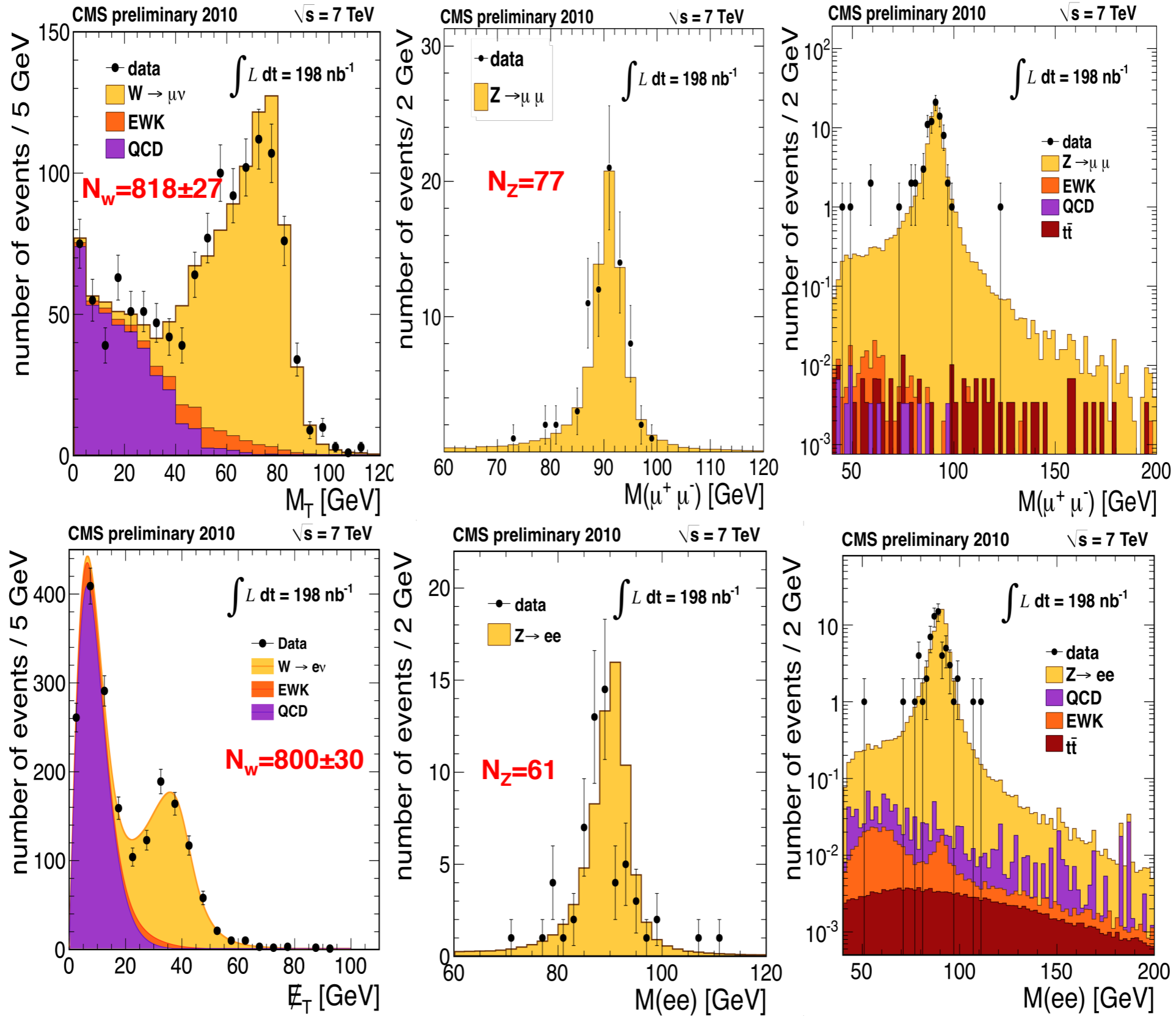
First Data: Minimum Bias (~1960's Physics)



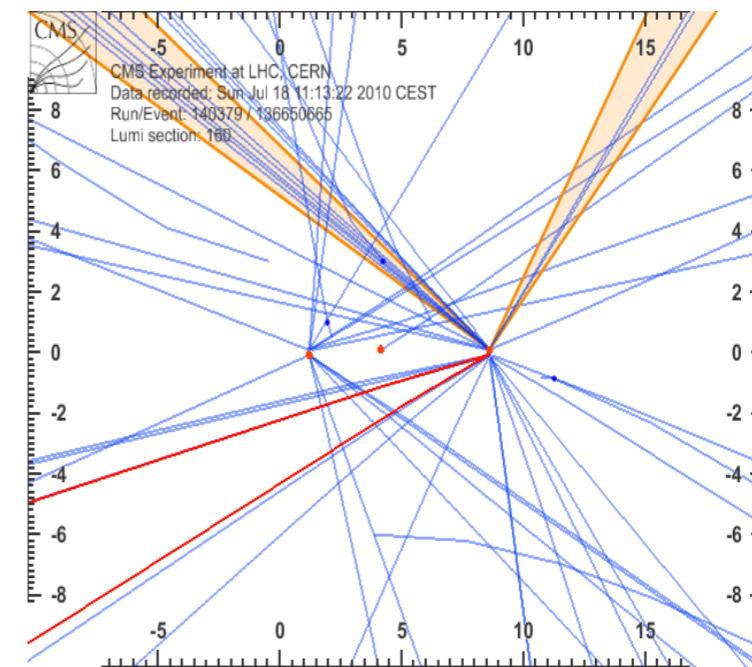
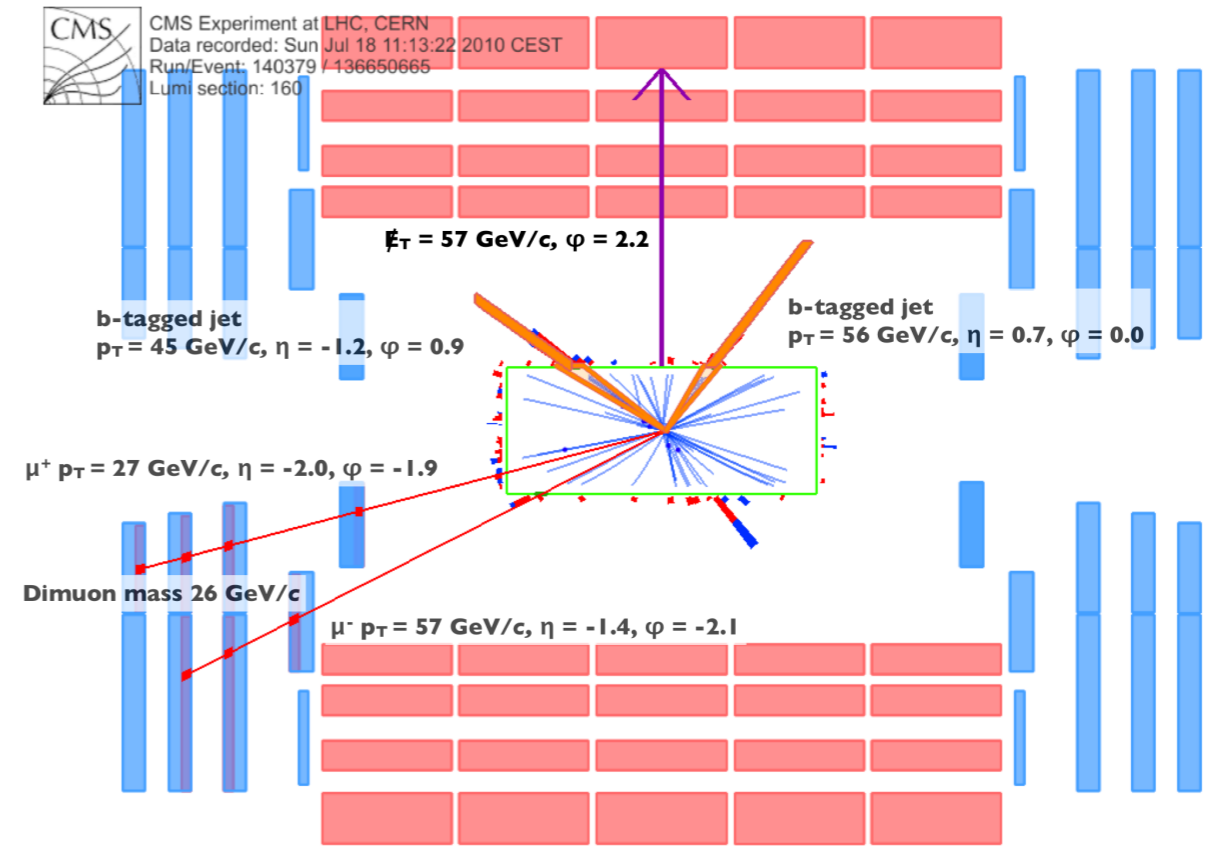
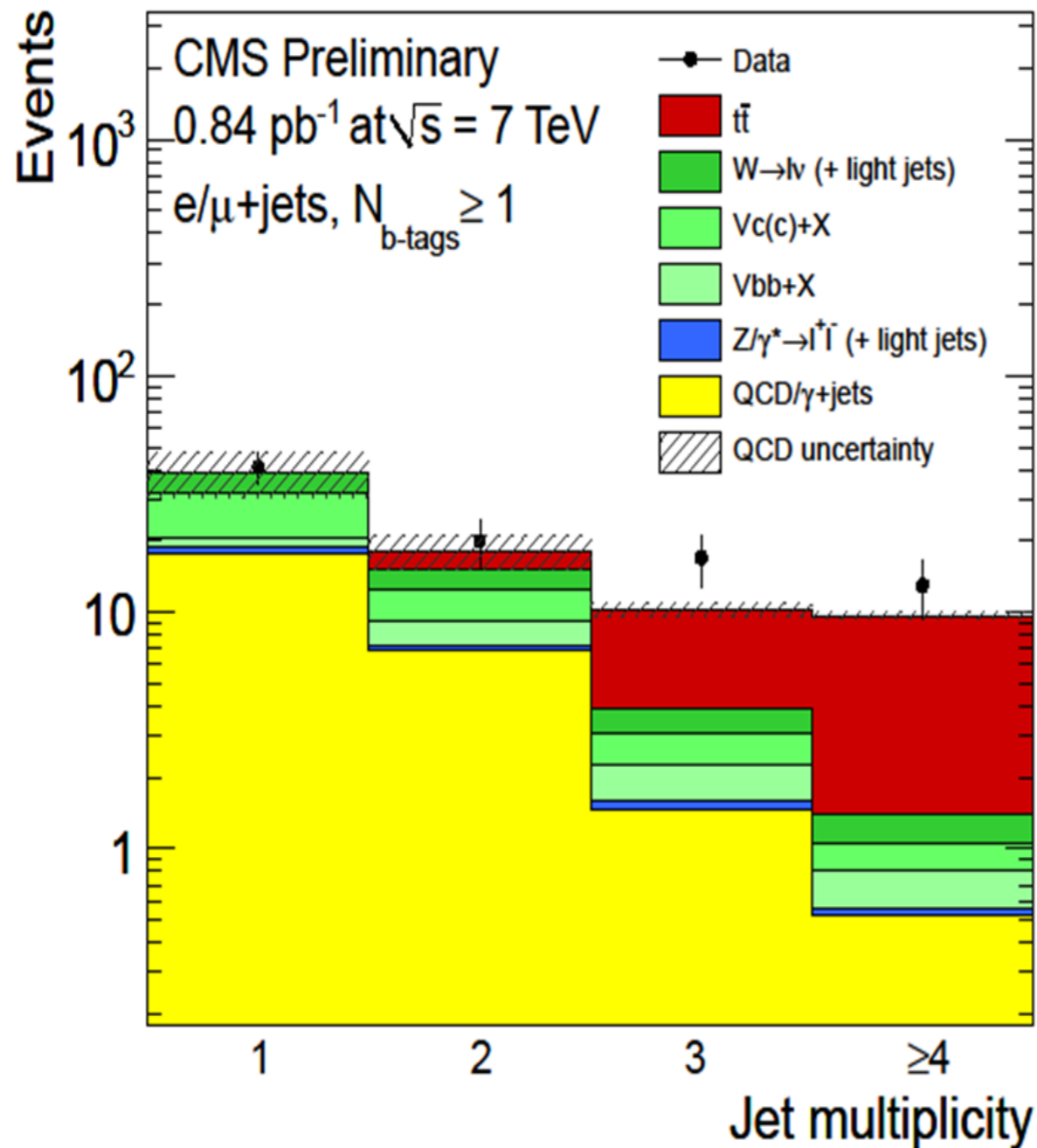
First Data: Jets (\sim 1970's Physics)



First Data: Electroweak (~1980's Physics)



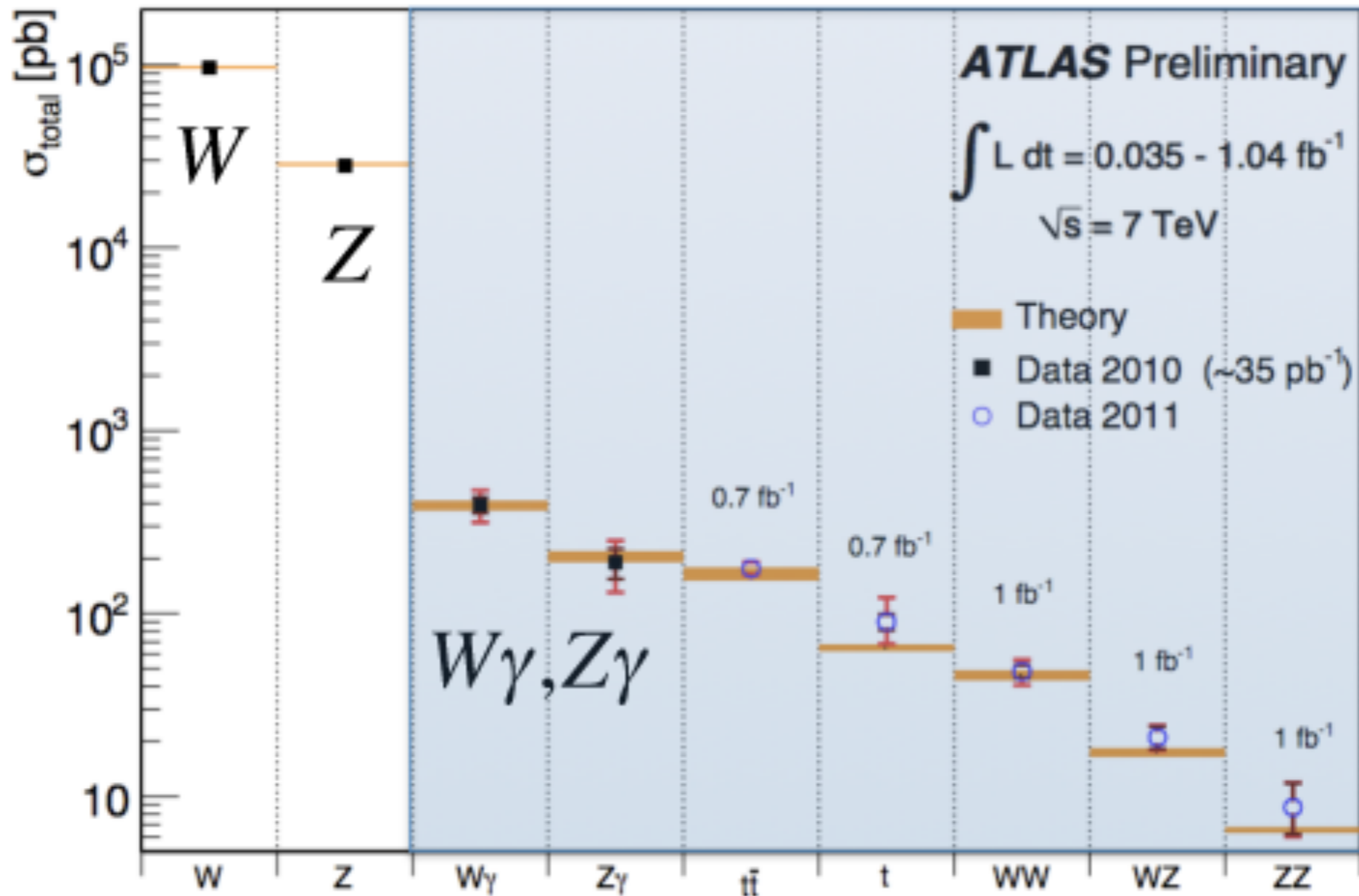
First Data: Top (~1990's Physics)



LHC Physics Programme

- ▶ Where we are today
 - ▶ Surprisingly well understood detectors and environment
 - ▶ Exponentially increasing luminosity
 - ▶ Huge number of channels to examine
- ▶ The emphasis
 - ▶ Increasingly detailed and precise SM measurements
 - ▶ Broad inclusive searches for a range of BSM phenomena
 - ▶ Few significant attempts at interpretation yet
- ▶ The challenges
 - ▶ The environment (esp pileup levels) is rapidly changing
 - ▶ *And not for the better*
 - ▶ Trigger conditions have rapidly evolved
 - ▶ *Already, some interesting physics is ~inaccessible due to trigger constraints*
 - ▶ Data handling – computing now becoming an issue
 - ▶ Organisation and prioritisation

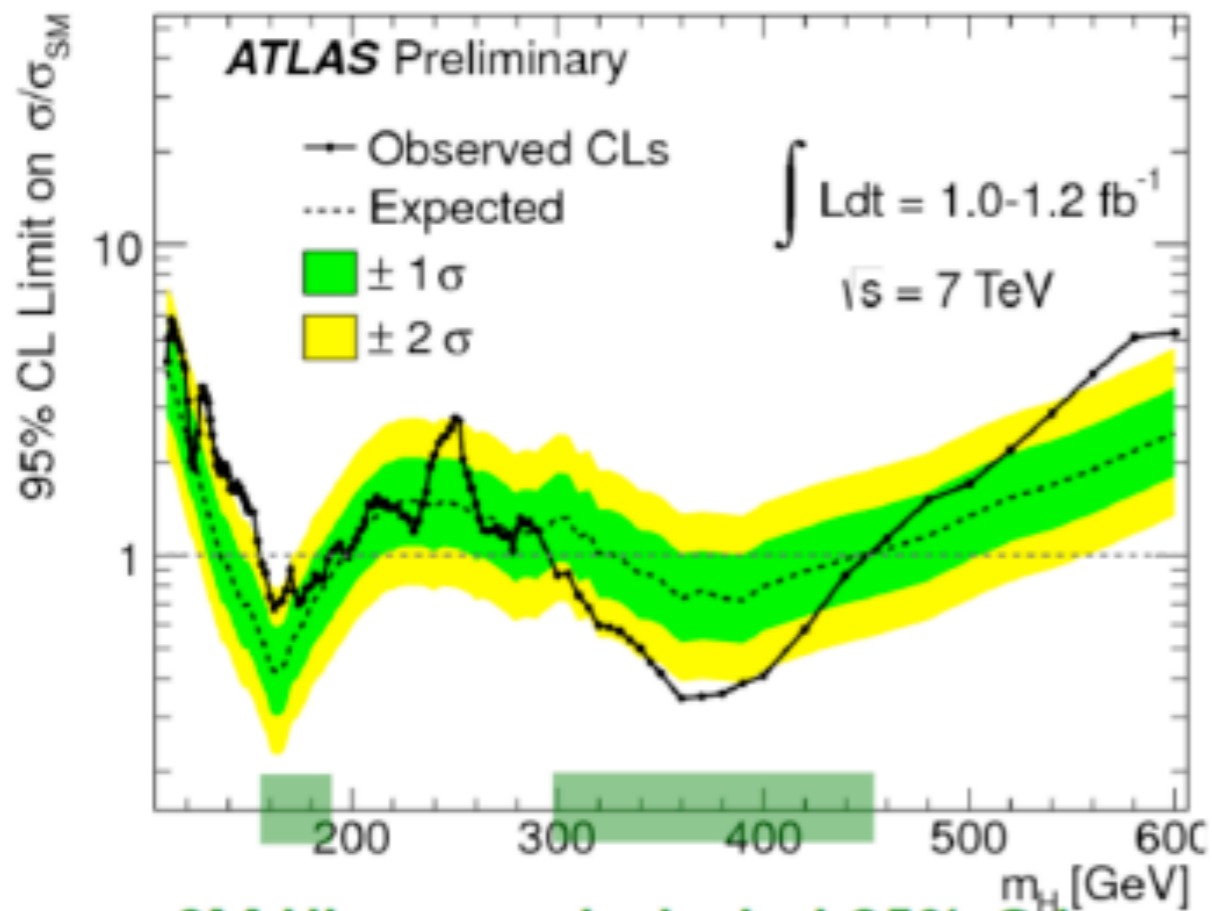
LHC Electroweak Measurements



← ~ SM Higgs

► Theory at NLO

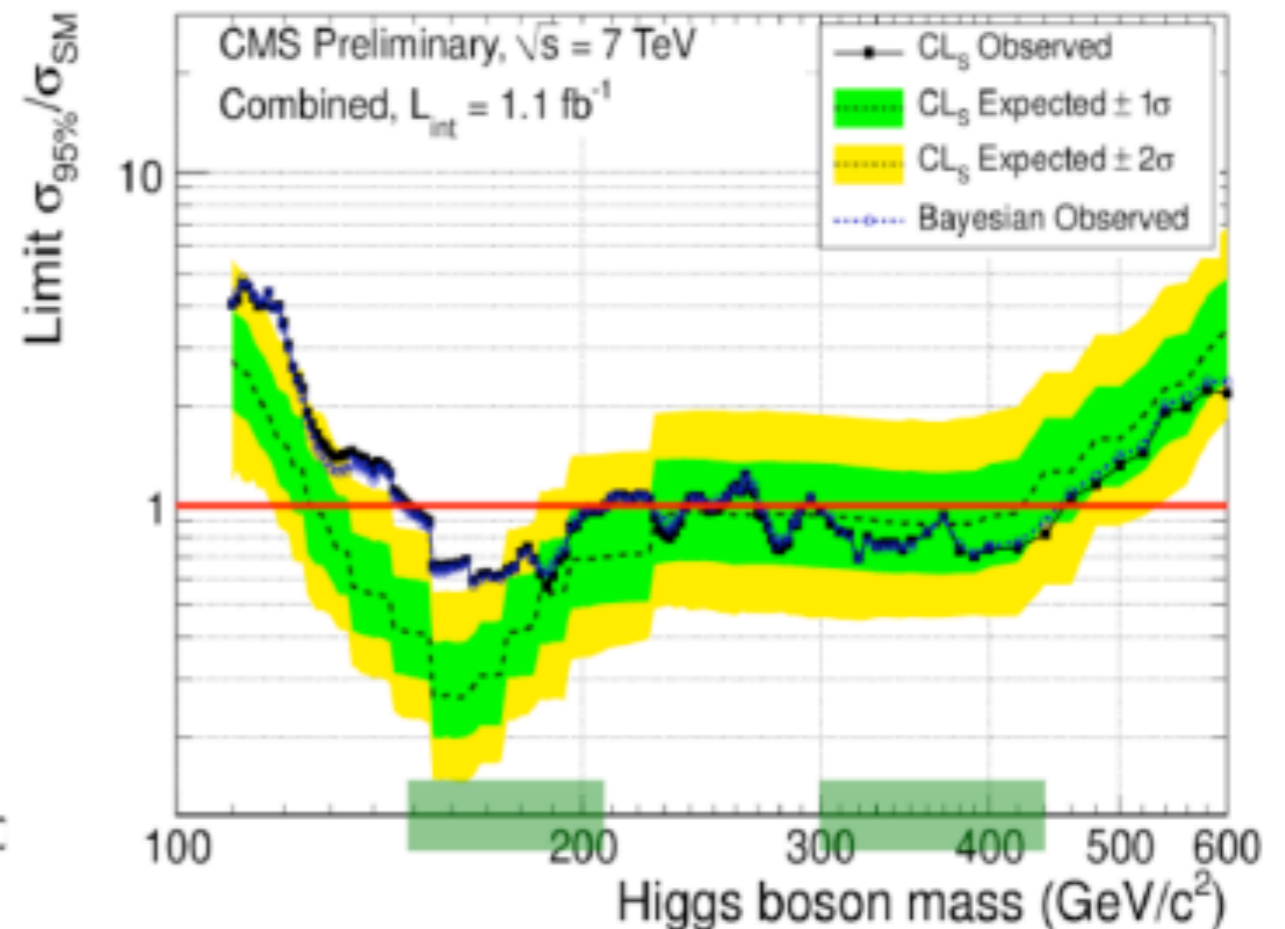
LHC Higgs Limits



SM Higgs excluded at 95% C.L.

$155 < M_H < 190 \text{ GeV}$

$295 < M_H < 450 \text{ GeV}$

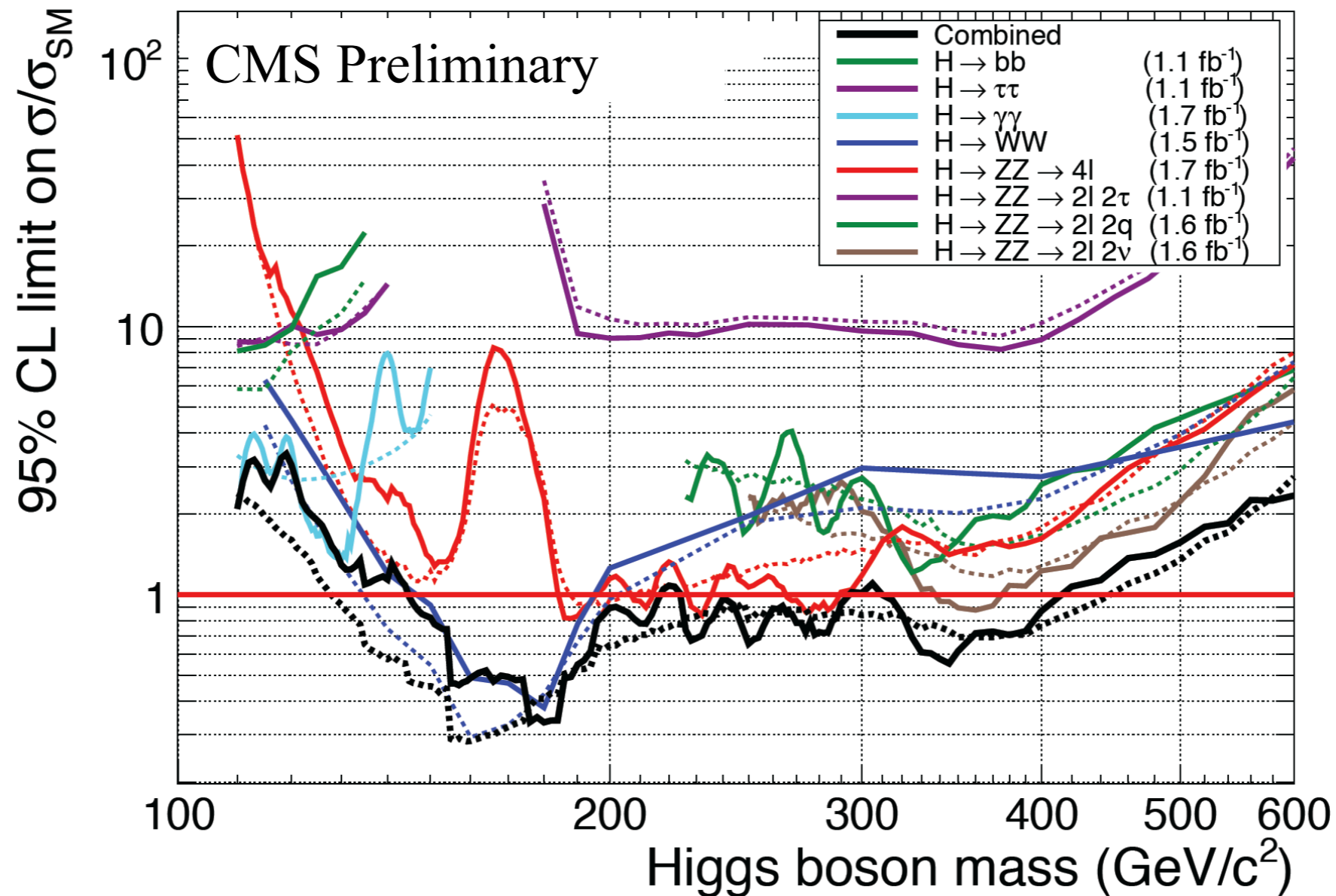


SM Higgs excluded at 95% C.L.

$149 < M_H < 206 \text{ GeV}$

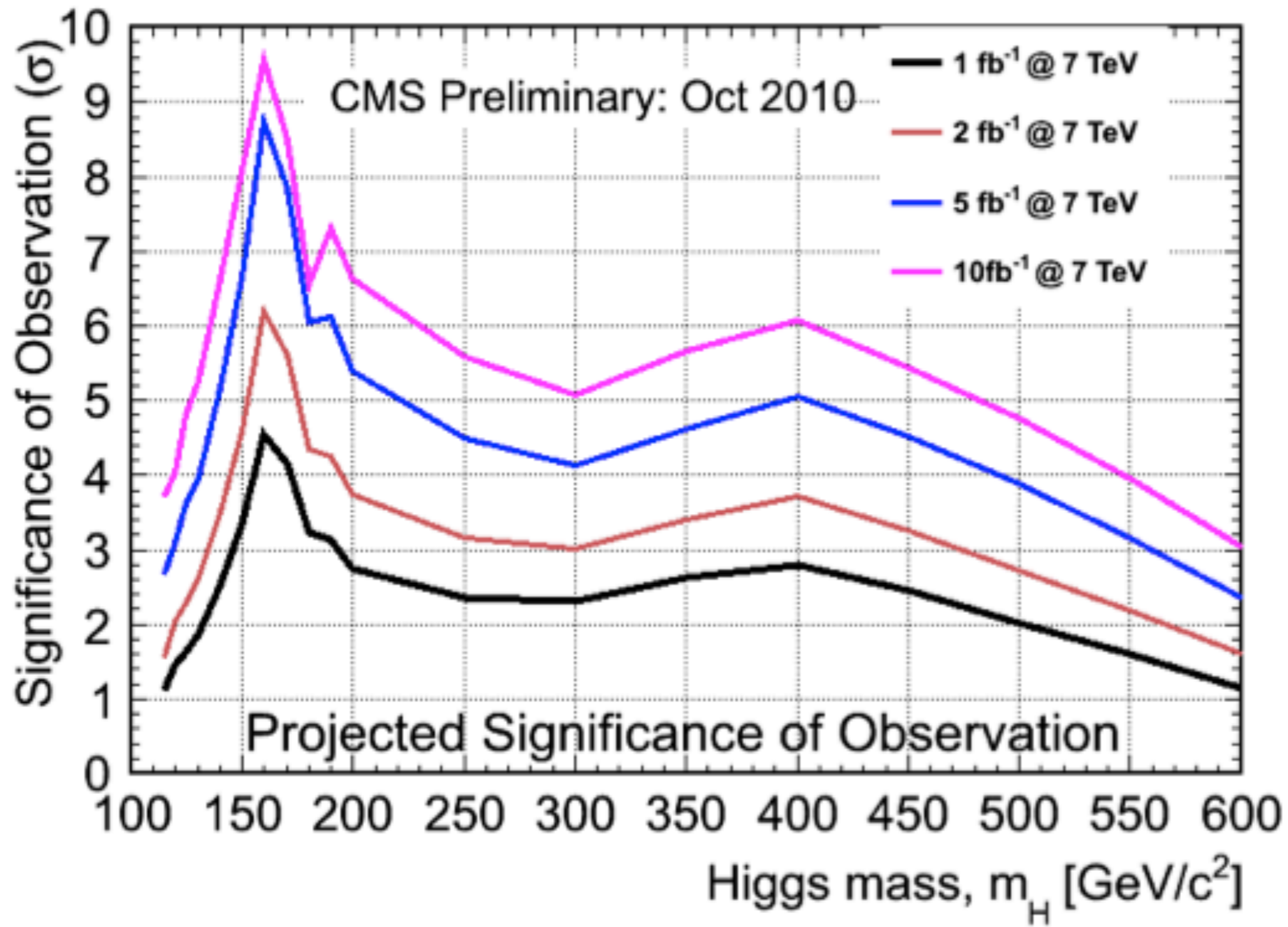
$300 < M_H < 440 \text{ GeV}$

Higgs Combinations

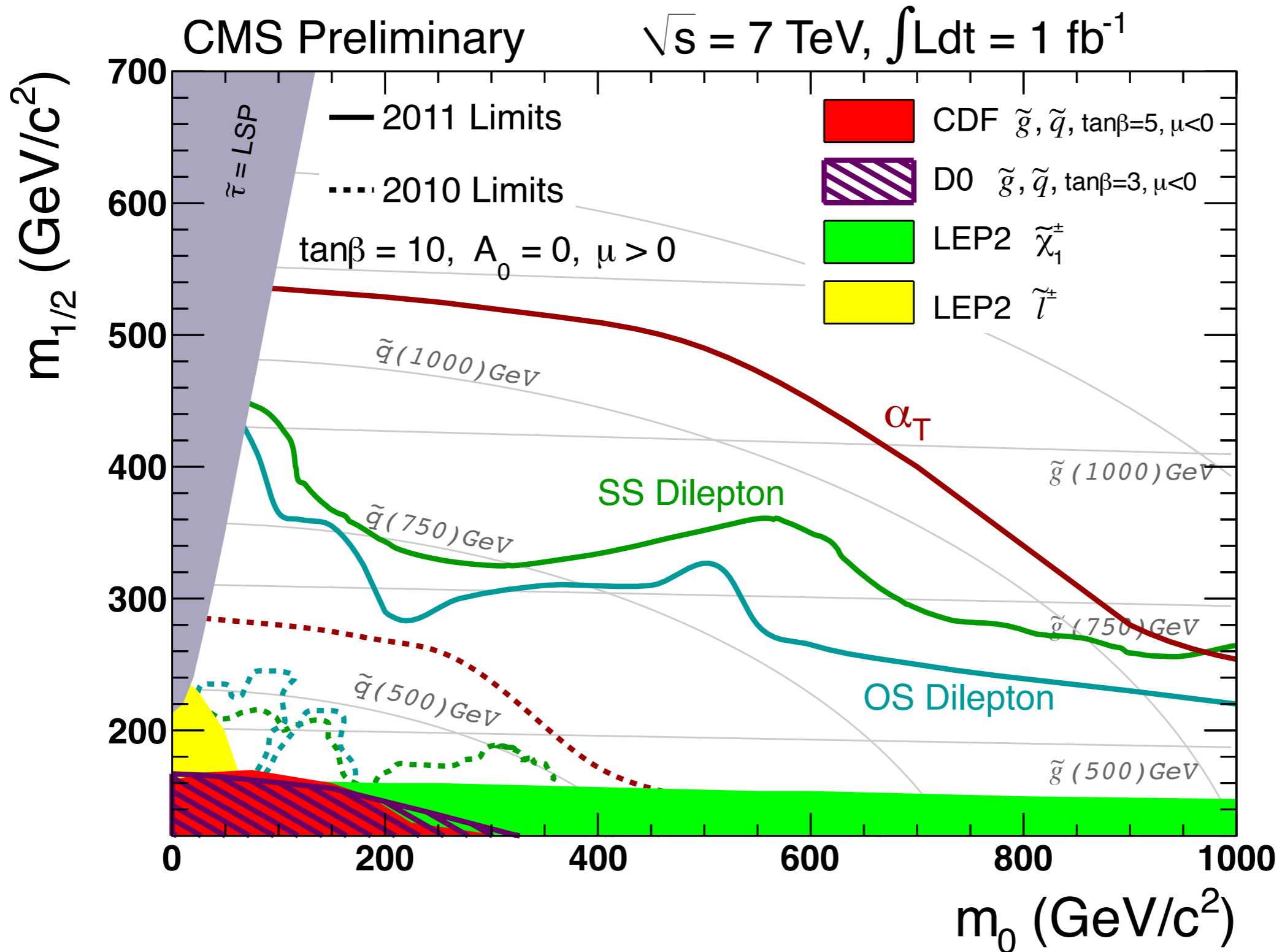


- ▶ A deep understanding of uncertainties and correlations required to form these combinations

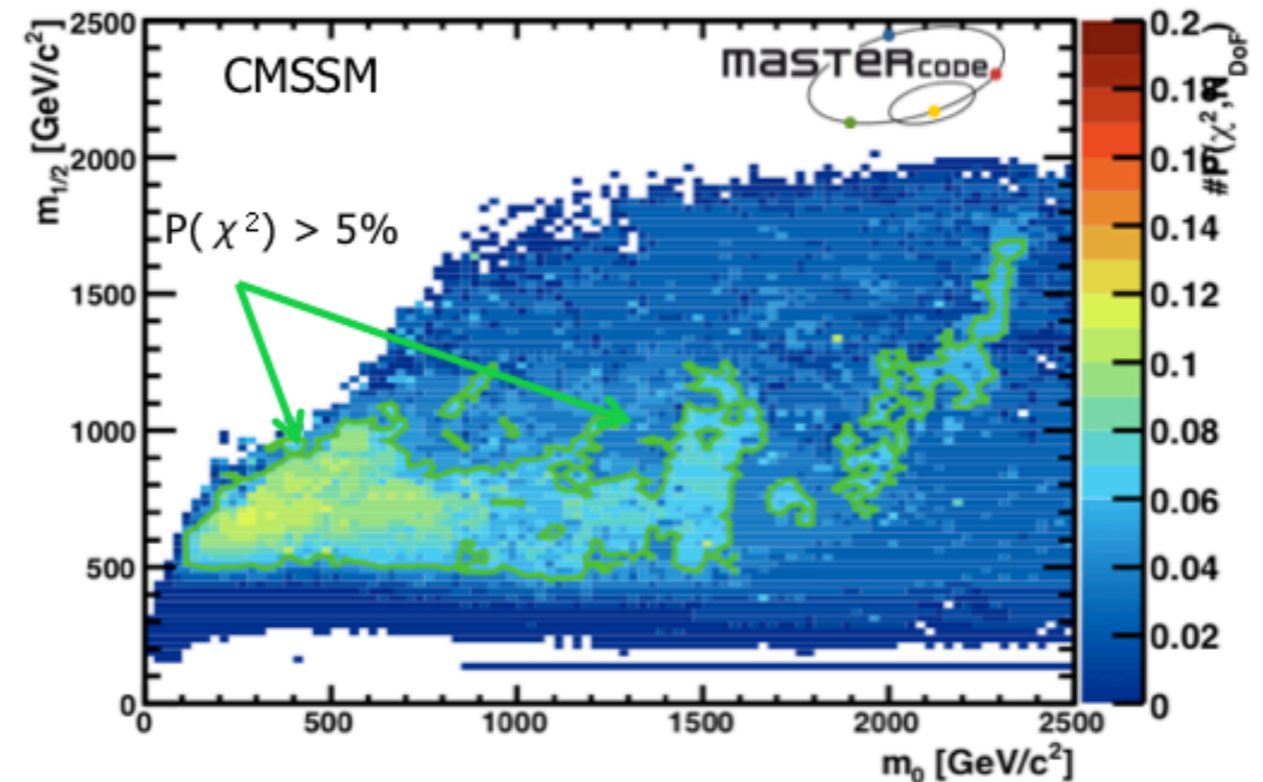
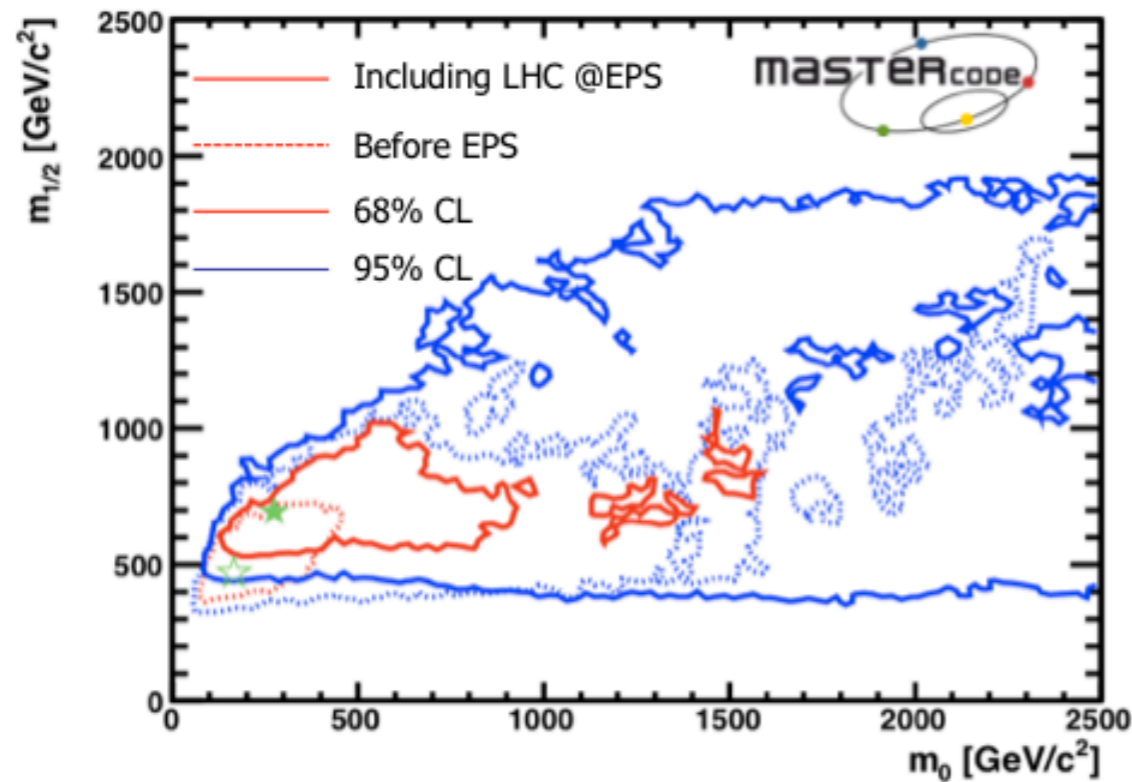
Higgs Outlook



CMSSM Limits

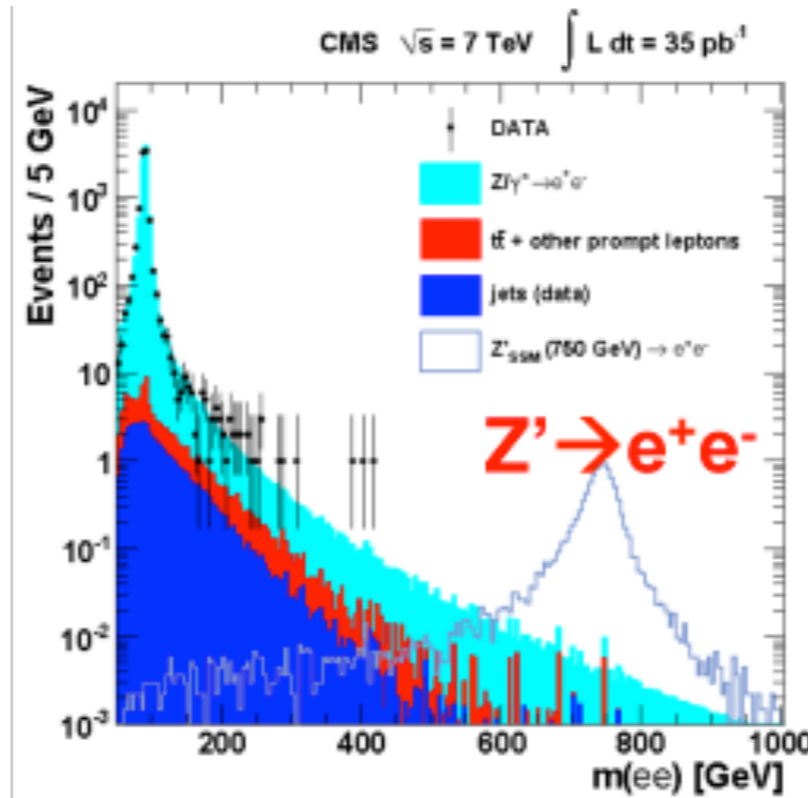
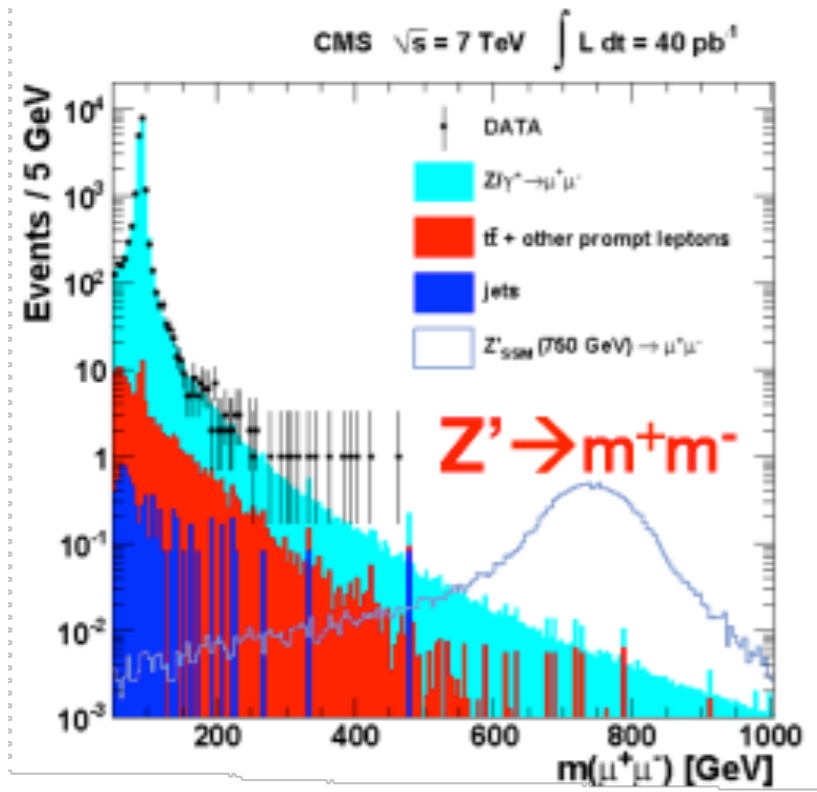


CMSSM Fits

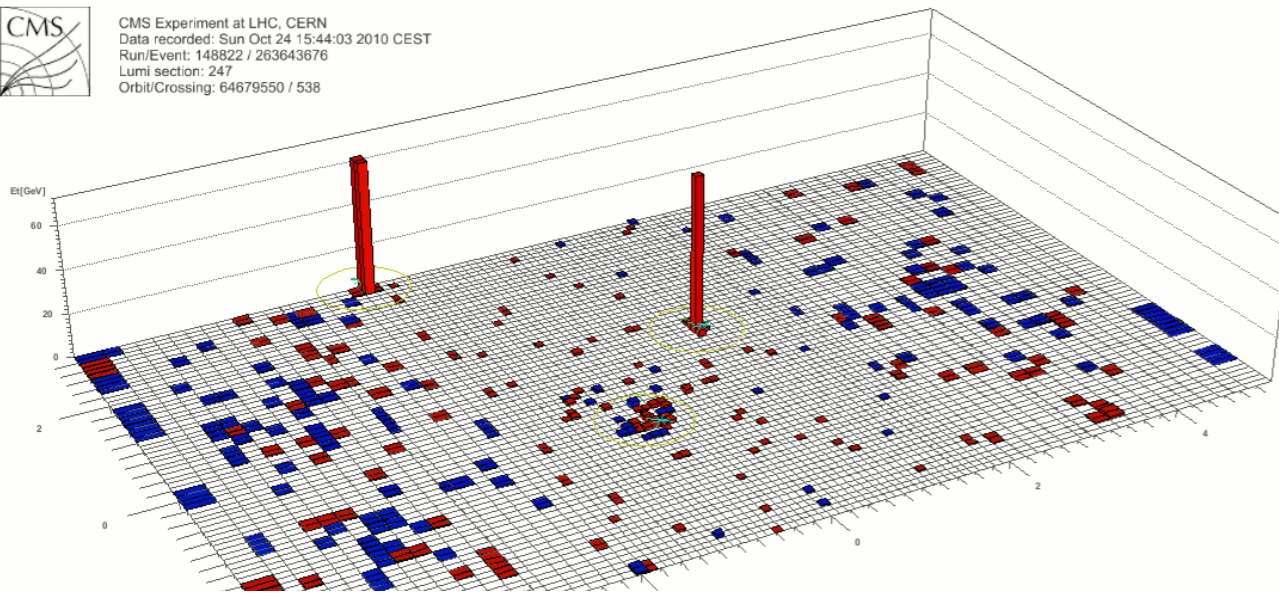


- ▶ Of course, plenty of theory space left yet
 - ▶ Constraints from other sources will be crucial here

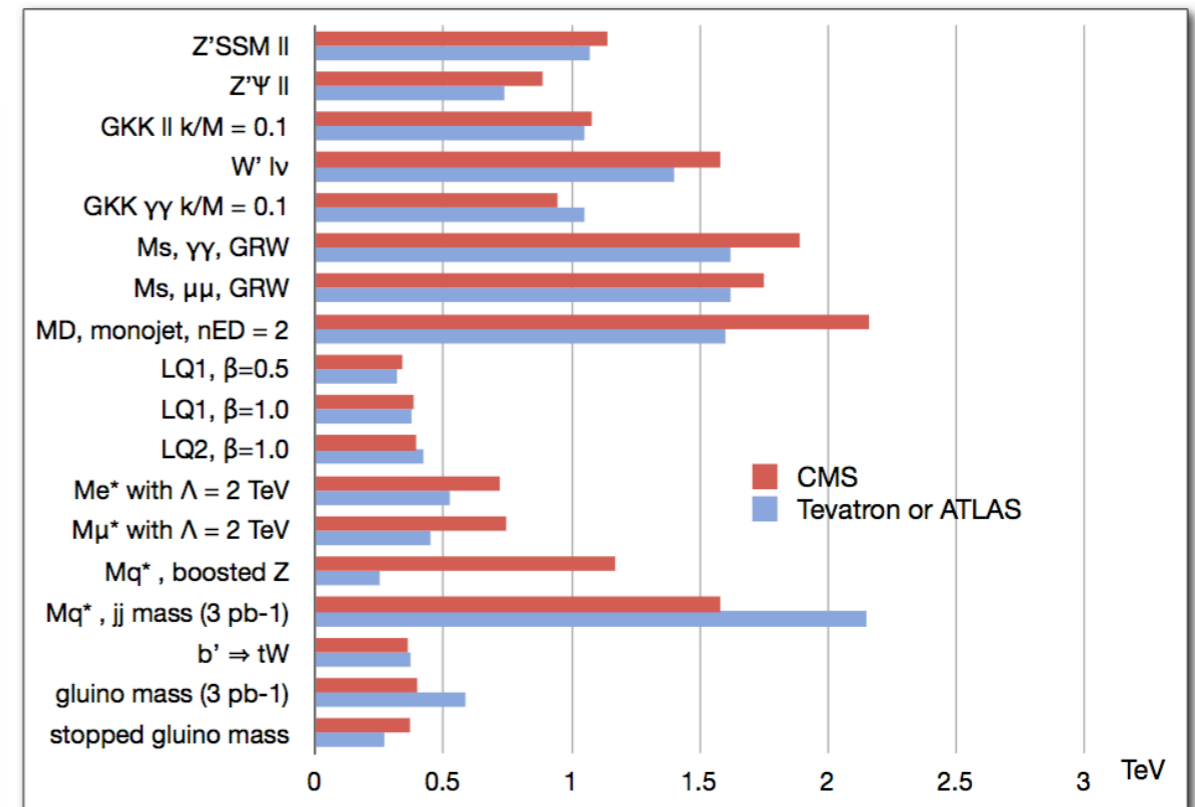
BSM - e.g. U(1) Extensions



CMS Experiment at LHC, CERN
 Data recorded: Sun Oct 24 15:44:03 2010 CEST
 Run/Event: 148822 / 263643676
 Lumi section: 247
 Orbit/Crossing: 64679550 / 538

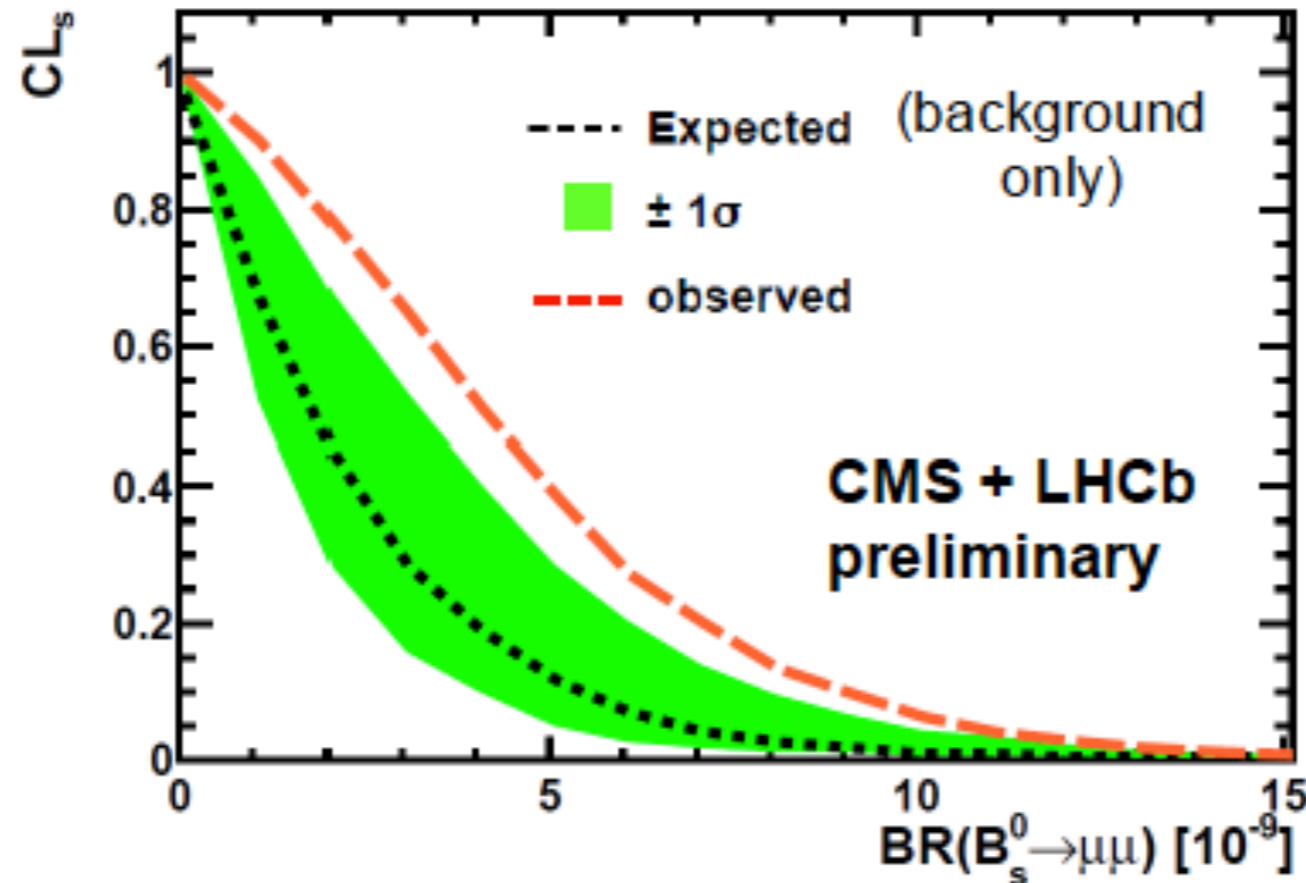


Very clean $\sim 60+80 \text{ GeV}$ E_t dielectron



$B_s \rightarrow \mu\mu$

A preliminary CMS-LHCb combination on $BR(B_s \rightarrow \mu^+\mu^-)$ has been performed, again using the CLs approach, & taking LHCb value of f_s/f_d as common input

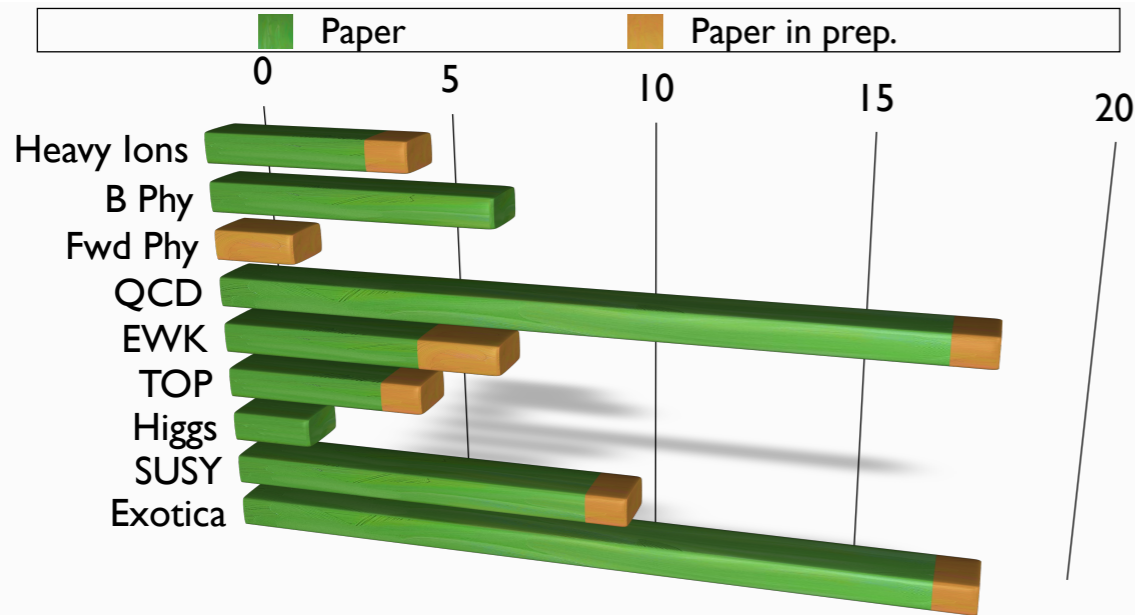


Observed limit at 95% (90%): $1.1 (0.9) \times 10^{-9}$

This is 3.4 times the expected SM value

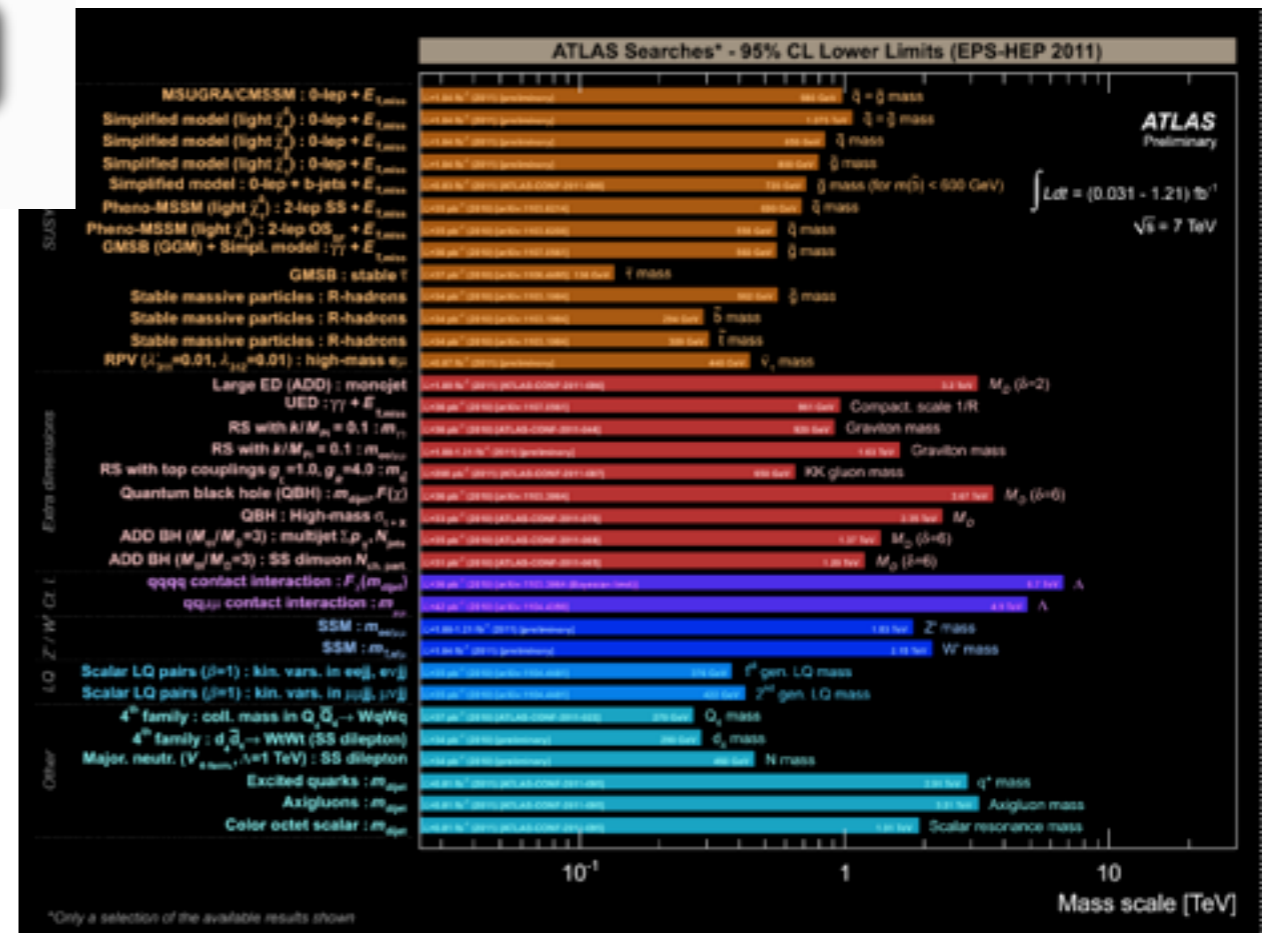
A BR of 1.8×10^{-8} has a CLs value of $\sim 0.3\%$

No Time To Mention...



In total : 65 papers on physics analyses, submitted, accepted or published
9 papers close to submission

<https://twiki.cern.ch/twiki/bin/view/CMSPublic/PhysicsResults>



A Warning From History

► Shown at BUSSTEPP2004...

Ladbrokes.com

http://www.ladbrokes.com/lbr_portal?action=do_lang_splash&form_name=lang_splash&LANG=en&STYLE=

Open account Banking Password? i Help

Account ID: Password: Login

Home Horses Greyhounds Football Sports Specials Poker Casino Games Lottos

View odds as decimals Change language Quick Menu to Betting >>>> (UK time) 9:26:46

Specials - Science Specials

Physics Breakthroughs

SELECTION	ODDS	BET NOW OR ADD TO MULTIPLE
Click here or on ODDS to change view order		
Understanding the origin of cosmic rays by 2010	4/1	Bet
The ATLAS experiment at CERN finding the Higgs Boson by 2010	6/1	Bet
The Laser Interferometer Gravitational Wave Observatory (LIGO) detecting gravitational waves by 2010	10/1	Bet
Building a fusion power station by 2010	100/1	Bet

2004-08-30 17:00:00

Selections will be settled on the basis of reports published in **New Scientist** magazine.

[Bet Help](#) [Bet Calculator](#)

After the Champagne..?

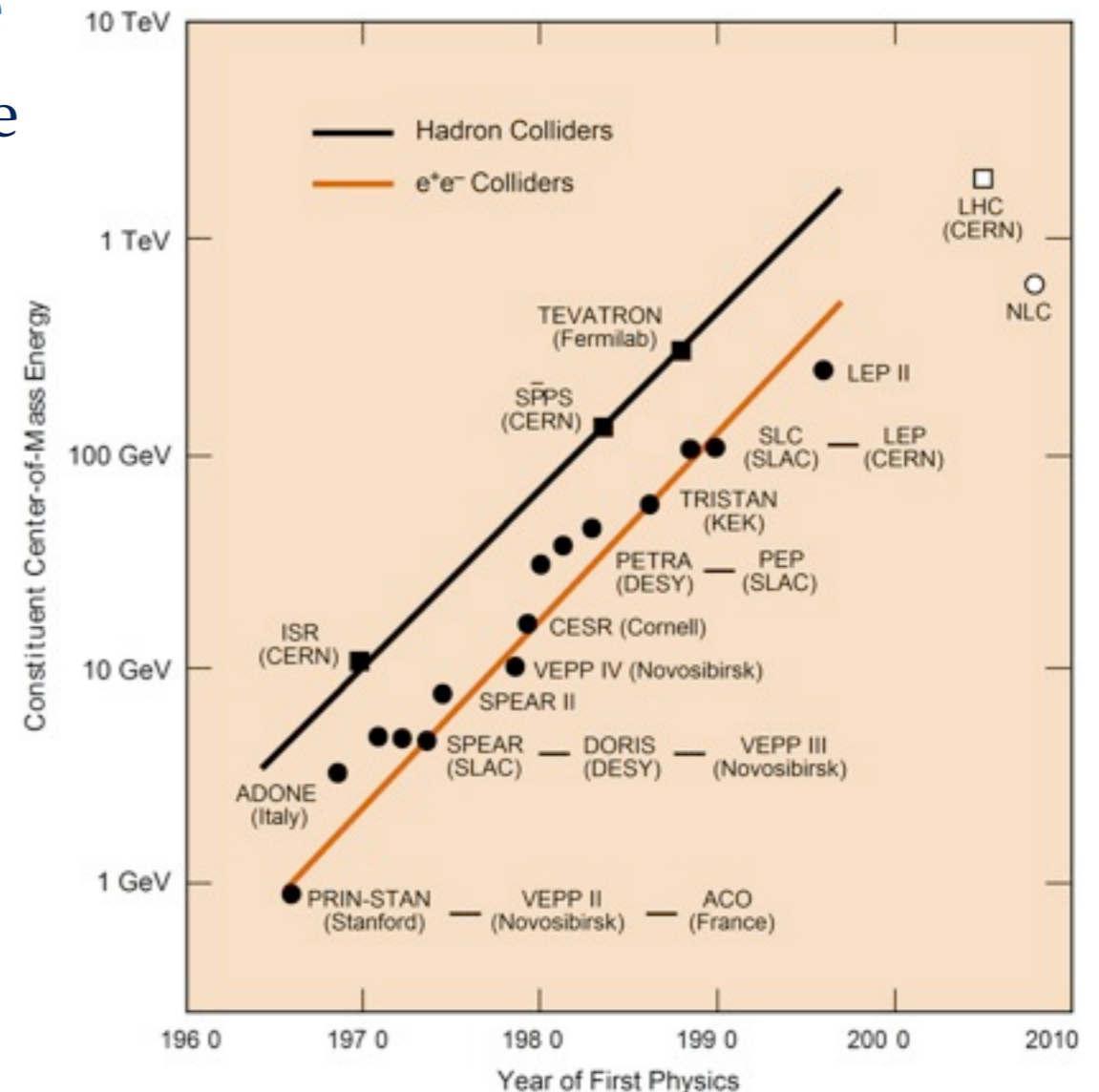
- ▶ The LHC is a discovery machine
 - ▶ It may not be a measurement machine
 - ▶ Though one always finds a way...

- ▶ Exploring further

- ▶ How do we exploit our new knowledge
- ▶ e.g. measure Higgs couplings
- ▶ e.g. measure SUSY spectrum
- ▶ e.g. find a 2nd KK resonance
- ▶ e.g. produce copious black holes

- ▶ The usual story applies

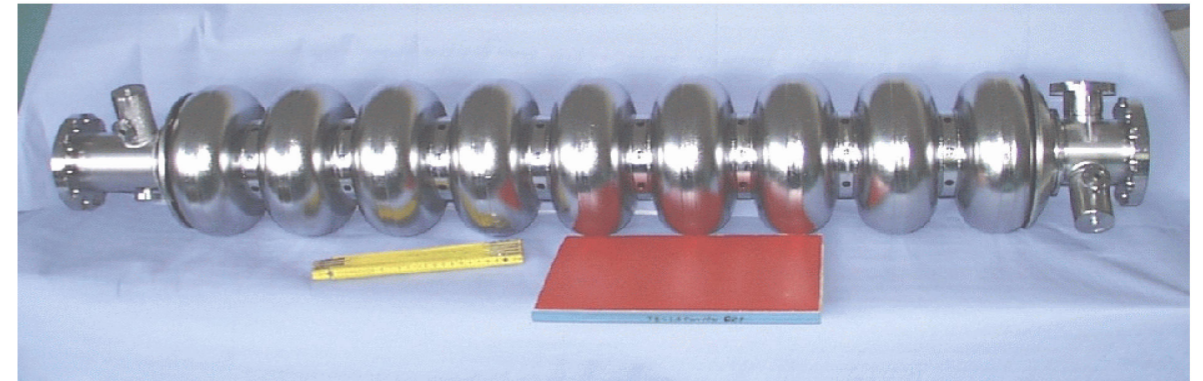
- ▶ More events: HL-SLHC (10x luminosity)
- ▶ More energy: HE-SLHC (3x energy)
- ▶ More precision, cleaner environment: Next Linear Collider (e+e-)



Future prospects

▶ Next Linear Collider

- ▶ Avoid problems with synchrotron radiation with H/E electron beam
- ▶ ~30km straight collider
- ▶ ILC: superconducting cavity technology
- ▶ CLIC: Two-beam acceleration (higher energy, tricky)
- ▶ What energy? GigaZ / Higgs / SUSY scan?

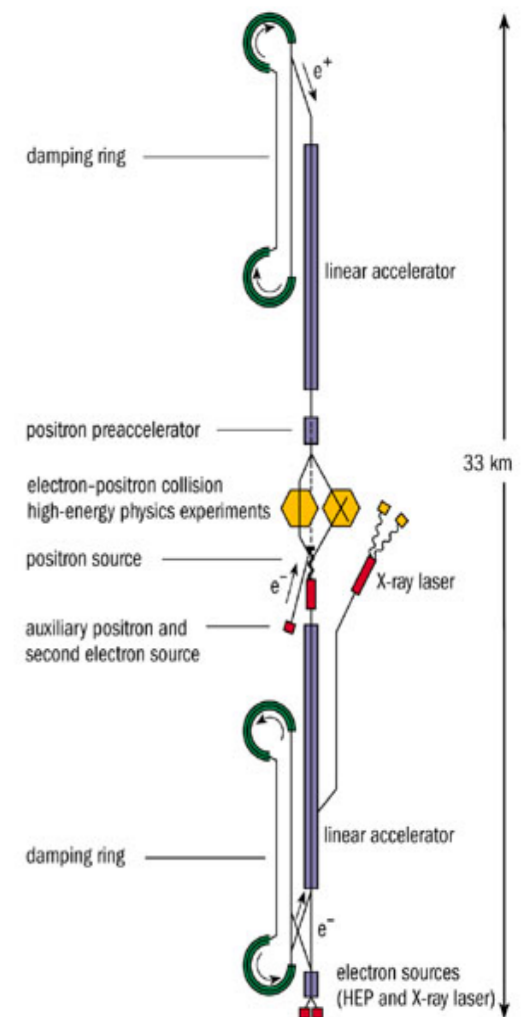


▶ SuperLHC

- ▶ Raise luminosity to $L=10^{35}$
- ▶ This is going to be experimentally tough
 - ▶ As hard as the original LHC development
- ▶ Work now under way on major GPD upgrades

▶ The far future? One can only speculate:

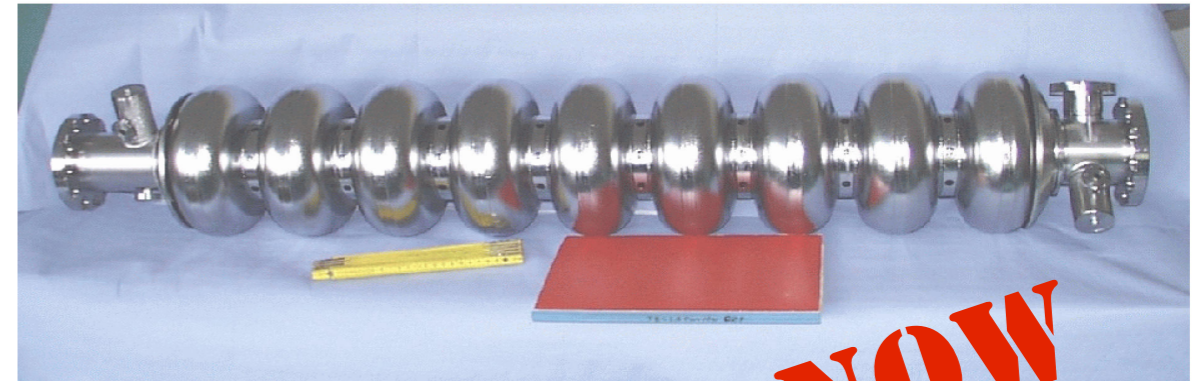
- ▶ Neutrino superbeam? Mu-mu collider?
- ▶ How can all of this be made affordable? How can we *downsize*?



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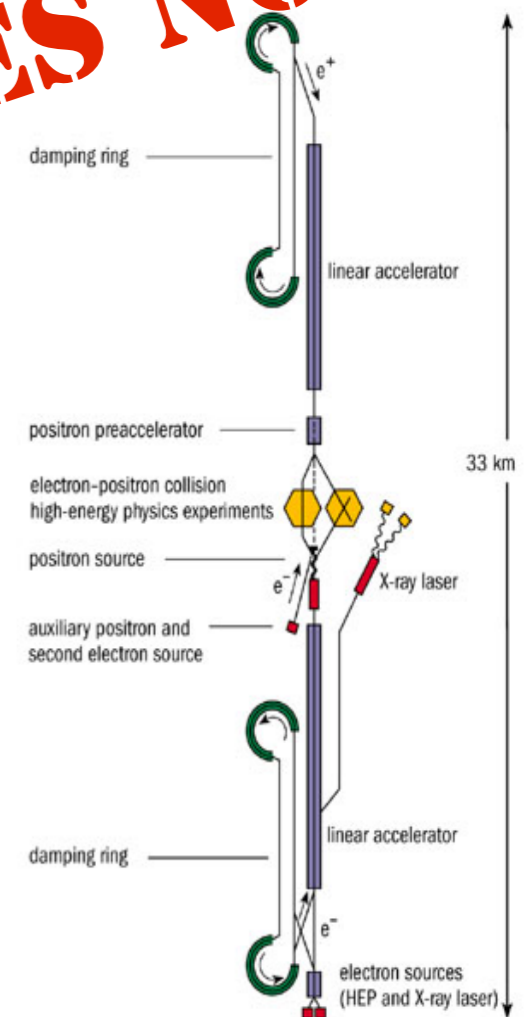


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STOP PRESS: HIGH LUMI (AB) B FACTORIES NOW APPROVED

Final Thoughts

- ▶ Theory and experiment
 - ▶ We have been waiting a long time to test theoretical ideas
 - ▶ Probably has not been a healthy situation
- ▶ Now is a great time to be an experimentalist!
 - ▶ The next (last?) big energy frontier is being crossed
 - ▶ BUT - looks like no easy ride to new discoveries
 - ▶ 1/fb down, 2999/fb to go! Though LHC is by no means the only story
 - ▶ Lots of work and smart thinking required in the coming years
- ▶ Now is a great time to be a theorist!
 - ▶ Many beautiful ideas in circulation – which are correct? If any?
 - ▶ Something entirely unexpected could happen
 - ▶ Lots of work and smart thinking required in the coming years
- ▶ We must work together to exploit the new opportunities
 - ▶ The best of luck!